Supersymmetric Particle Searches

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SUPERSYMMETRIC MODEL ASSUMPTIONS

The exclusion of particle masses within a mass range $(m_1,\ m_2)$ will be denoted with the notation "none $m_1 - m_2$ " in the VALUE column of the

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$\tilde{\chi}_1^0$ (Lightest Neutralino) MASS LIMIT

 $\widetilde{\chi}_1^0$ is often assumed to be the lightest supersymmetric particle (LSP). See also the $\widetilde{\chi}_2^{\bar{0}}$, $\widetilde{\chi}_3^{0}$, $\widetilde{\chi}_4^{0}$ section below.

We have divided the $\widetilde{\chi}_1^0$ listings below into five sections:

- 1) Accelerator limits for stable $\tilde{\chi}_1^0$,
- 2) Bounds on $\widetilde{\chi}_1^0$ from dark matter searches, 3) Bounds on $\widetilde{\chi}_1^0$ elastic cross sections from dark matter searches,
- 4) Other bounds on $\tilde{\chi}_1^0$ from astrophysics and cosmology, and
- 5) Bounds on unstable $\widetilde{\chi}_1^0$.

- Accelerator limits for stable $\widetilde{\chi}_1^0$

Unless otherwise stated, results in this section assume spectra, production rates, decay modes, and branching ratios as evaluated in the MSSM, with gaugino and sfermion mass unification at the GUT scale. These papers generally study production of $\widetilde{\chi}_{i}^{0}\widetilde{\chi}_{i}^{0}$ ($i \geq 1, j \geq 2$), $\widetilde{\chi}_{1}^{+}\widetilde{\chi}_{1}^{-}$, and (in the case of hadronic collisions) $\widetilde{\chi}_1^+\widetilde{\chi}_2^0$ pairs. The mass limits on $\widetilde{\chi}_1^0$ are either direct, or follow indirectly from the constraints set by the non-observation of $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_2^0$ states on the gaugino and higgsino MSSM parameters M_2 and μ . In some cases, information is used from the nonobservation of slepton decays.

Obsolete limits obtained from e^+e^- collisions up to \sqrt{s} =184 GeV have been removed from this compilation and can be found in the 2000 Edition (The European Physical Journal C15 1 (2000)) of this Review. $\Delta m_0 = m_{\widetilde{\chi}_2^0} - m_{\widetilde{\chi}_1^0}$

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>40	95	¹ ABBIENDI	04H	OPAL	all $ aneta$, $\Delta m_0 >$ 5 GeV, $m_0 >$ 500 GeV, $A_0 =$ 0
>42.4	95	² HEISTER			all tan eta , all Δm_0 , all m_0
>39.2	95	³ ABDALLAH	03м	DLPH	all tan eta , $m_{\widetilde{ u}}>$ 500 GeV

>46 95 4 ABDALLAH 03M DLPH all
$$\tan\beta$$
, all Δm_0 , all m_0 >32.5 95 5 ACCIARRI 00D L3 $\tan\beta > 0.7$, $\Delta m_0 > 3$ GeV, all

• • We do not use the following data for averages, fits, limits, etc. • •

6
 ABBOTT 98C D0 $p\overline{p}
ightarrow \, \widetilde{\chi}_{1}^{\pm} \, \widetilde{\chi}_{2}^{0}$ $>$ 41 95 7 ABE 98J CDF $p\overline{p}
ightarrow \, \widetilde{\chi}_{1}^{\pm} \, \widetilde{\chi}_{2}^{0}$

- 1 ABBIENDI 04H search for charginos and neutralinos in events with acoplanar leptons+jets and multi-jet final states in the 192–209 GeV data, combined with the results on leptonic final states from ABBIENDI 04. The results hold for a scan over the parameter space covering the region 0 < M_2 <5000 GeV, -1000 < μ <1000 GeV and tan β from 1 to 40. This limit supersedes ABBIENDI 00H.
- 2 HEISTER 04 data collected up to 209 GeV. Updates earlier analysis of selectrons from HEISTER 02E, includes a new analysis of charginos and neutralinos decaying into stau and uses results on charginos with initial state radiation from HEISTER 02J. The limit is based on the direct search for charginos and neutralinos, the constraints from the slepton search and the Higgs mass limits from HEISTER 02 using a top mass of 175 GeV, interpreted in a framework with universal gaugino and sfermion masses. Assuming the mixing in the stau sector to be negligible, the limit improves to 43.1 GeV. Under the assumption of MSUGRA with unification of the Higgs and sfermion masses, the limit improves to 50 GeV, and reaches 53 GeV for $A_0\,=\,0$. These limits include and update the results of BARATE 01.
- ³ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV. A limit on the mass of $\widetilde{\chi}_1^0$ is derived from direct searches for neutralinos combined with the chargino search. Neutralinos are searched in the production of $\widetilde{\chi}_1^0\widetilde{\chi}_2^0$, $\widetilde{\chi}_1^0\widetilde{\chi}_3^0$, as well as $\widetilde{\chi}_2^0\widetilde{\chi}_3^0$ and $\widetilde{\chi}_2^0\widetilde{\chi}_4^0$ giving rise to cascade decays, and $\widetilde{\chi}_1^0\widetilde{\chi}_2^0$ and $\widetilde{\chi}_1^0\widetilde{\chi}_2^0$, followed by the decay $\widetilde{\chi}_2^0 \to \widetilde{\tau}\tau$. The results hold for the parameter space defined by values of $M_2 < 1$ TeV, $|\mu| \leq 2$ TeV with the $\widetilde{\chi}_1^0$ as LSP. The limit is obtained for $\tan\beta = 1$ and large m_0 , where $\widetilde{\chi}_2^0\widetilde{\chi}_4^0$ and chargino pair production are important. If the constraint from Higgs searches is also imposed, the limit improves to 49.0 GeV in the M_h^{max} scenario with $m_t=174.3$ GeV. These limits update the results of ABREU 00J.
- ⁴ ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV. An indirect limit on the mass of $\widetilde{\chi}_1^0$ is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays and $\widetilde{\tau}\tau$ final states), for charginos (for all Δm_+) and for sleptons, stop and sbottom. The results hold for the full parameter space defined by values of $M_2 < 1$ TeV, $|\mu| \leq 2$ TeV with the $\widetilde{\chi}_1^0$ as LSP. Constraints from the Higgs search in the M_h^{max} scenario assuming m_t =174.3 GeV are included. The limit is obtained for $\tan\beta \geq 5$ when stau mixing leads to mass degeneracy between $\widetilde{\tau}_1$ and $\widetilde{\chi}_1^0$ and the limit is based on $\widetilde{\chi}_2^0$ production followed by its decay to $\widetilde{\tau}_1\tau$. In the pathological scenario where m_0 and $|\mu|$ are large, so that the $\widetilde{\chi}_2^0$ production cross section is negligible, and where there is mixing in the stau sector but not in stop nor sbottom, the limit is based on charginos with soft decay products and an ISR photon. The limit then degrades to 39 GeV. See Figs 40–42 for the dependence of the limit on $\tan\beta$ and $m_{\widetilde{\nu}}$. These limits update the results of ABREU 00W.
- 5 ACCIARRI 00D data collected at $\sqrt{s}{=}189$ GeV. The results hold over the full parameter space defined by 0.7 \leq tan β \leq 60, 0 \leq M_2 \leq 2 TeV, m_0 \leq 500 GeV, $|\mu|$ \leq 2 TeV The minimum mass limit is reached for tan $\beta{=}1$ and large m_0 . The results of slepton searches from ACCIARRI 99W are used to help set constraints in the region of small m_0 . The limit improves to 48 GeV for m_0 \gtrsim 200 GeV and tan β \gtrsim 10. See their Figs. 6–8 for the tan β and m_0 dependence of the limits. Updates ACCIARRI 98F.

- ⁶ ABBOTT 98C searches for trilepton final states $(\ell=e,\mu)$. See footnote to ABBOTT 98C in the Chargino Section for details on the assumptions. Assuming a negligible decay rate of $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_2^0$ to quarks, they obtain $m_{\widetilde{\chi}_2^0} \gtrsim 51$ GeV.
- ⁷ ABE 98J searches for trilepton final states ($\ell=e,\mu$). See footnote to ABE 98J in the Chargino Section for details on the assumptions. The quoted result corresponds to the best limit within the selected range of parameters, obtained for $m_{\widetilde{q}} > m_{\widetilde{g}}$, $\tan\beta=2$, and $\mu=-600$ GeV.

- Bounds on $\widetilde{\chi}^0_1$ from dark matter searches -

These papers generally exclude regions in the $M_2-\mu$ parameter plane assuming that $\widetilde{\chi}_1^0$ is the dominant form of dark matter in the galactic halo. These limits are based on the lack of detection in laboratory experiments or by the absence of a signal in underground neutrino detectors. The latter signal is expected if $\widetilde{\chi}_1^0$ accumulates in the Sun or the Earth and annihilates into high-energy ν 's.

 VALUE
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 TECN

 • • • We do not use the following data for averages, fits, limits, etc. • •

⁸ ACHTERBERG 06 ⁹ ACKERMANN 06 **AMND** ¹⁰ DEBOER **RVUE** ¹¹ DESAI **SKAM** ¹¹ AMBROSIO 99 **MCRO** ¹² LOSECCO 95 **RVUE** 13 MORI KAMI ¹⁴ BOTTINO 92 COSM ¹⁵ BOTTINO 91 **RVUE** ¹⁶ GELMINI COSM ¹⁷ KAMIONKOW.91 **RVUE** ¹⁸ MORI KAMI ¹⁹ OLIVE **COSM**

none 4-15 GeV

- ⁸ ACHTERBERG 06 is based on data collected during 421.9 effective days with the AMANDA detector. They looked for interactions of ν_{μ} s from the centre of the Earth over a background of atmospheric neutrinos and set 90 % CL limits on the muon flux. Their limit is compared with the muon flux expected from neutralino annihilations into W^+W^- and $b\overline{b}$ at the centre of the Earth for MSSM parameters compatible with the relic dark matter density, see their Fig. 7.
- 9 ACKERMANN 06 is based on data collected during 143.7 days with the AMANDA-II detector. They looked for interactions of $\nu_{\mu} s$ from the Sun over a background of atmospheric neutrinos and set 90 % CL limits on the muon flux. Their limit is compared with the muon flux expected from neutralino annihilations into W^+W^- in the Sun for SUSY model parameters compatible with the relic dark matter density, see their Fig. 3.
- ¹⁰ DEBOER 06 interpret an excess of diffuse Galactic gamma rays observed with the EGRET satellite as originating from π^0 decays from the annihilation of neutralinos into quark jets. They analyze the corresponding parameter space in a supergravity inspired MSSM model with radiative electroweak symmetry breaking, see their Fig. 3 for the preferred region in the $(m_0, m_{1/2})$ plane of a scenario with large $\tan\beta$.
- 11 AMBROSIO 99 and DESAI 04 set new neutrino flux limits which can be used to limit the parameter space in supersymmetric models based on neutralino annihilation in the Sun and the Earth.

 12 LOSECCO 95 reanalyzed the IMB data and places lower limit on $m_{\widetilde{\chi}^0_1}$ of 18 GeV if the LSP is a photino and 10 GeV if the LSP is a higgsino based on LSP annihilation in the sun producing high-energy neutrinos and the limits on neutrino fluxes from the IMB detector.

 13 MORI 93 excludes some region in $M_2-\mu$ parameter space depending on $\tan\beta$ and lightest scalar Higgs mass for neutralino dark matter $m_{\widetilde{\chi}0}>\!\!m_W$, using limits on upgoing muons produced by energetic neutrinos from neutralino annihilation in the Sun and the Earth.

- 14 BOTTINO 92 excludes some region $M_2\text{-}\mu$ parameter space assuming that the lightest neutralino is the dark matter, using upgoing muons at Kamiokande, direct searches by Ge detectors, and by LEP experiments. The analysis includes top radiative corrections on Higgs parameters and employs two different hypotheses for nucleon-Higgs coupling. Effects of rescaling in the local neutralino density according to the neutralino relic abundance are taken into account.
- 15 BOTTINO 91 excluded a region in $M_2-\mu$ plane using upgoing muon data from Kamioka experiment, assuming that the dark matter surrounding us is composed of neutralinos and that the Higgs boson is not too heavy.
- $^{16}\,\mathrm{GELMINI}$ 91 exclude a region in $M_2-\mu$ plane using dark matter searches.
- 17 KAMIONKOWSKI 91 excludes a region in the $M_2-\mu$ plane using IMB limit on upgoing muons originated by energetic neutrinos from neutralino annihilation in the sun, assuming that the dark matter is composed of neutralinos and that $m_{H_1^0} \lesssim 50$ GeV. See Fig. 8 in the paper.
- 18 MORI 91B exclude a part of the region in the $M_2-\mu$ plane with $m_{\widetilde{\chi}^0_1}\lesssim 80$ GeV using a limit on upgoing muons originated by energetic neutrinos from neutralino annihilation in the earth, assuming that the dark matter surrounding us is composed of neutralinos and that $m_{H^0_1}\lesssim 80$ GeV.
- ¹⁹ OLIVE 88 result assumes that photinos make up the dark matter in the galactic halo. Limit is based on annihilations in the sun and is due to an absence of high energy neutrinos detected in underground experiments. The limit is model dependent.

$----\widetilde{\chi}^0_1$ -ho elastic cross section extstyle -

Experimental results on the $\widetilde{\chi}_1^0$ -p elastic cross section are evaluated at $m_{\widetilde{\chi}_1^0}{=}100$ GeV. The experimental results on the cross section are often mass dependent. Therefore, the mass and cross section results are also given where the limit is strongest, when appropriate. Results are quoted separately for spin-dependent interactions (based on an effective 4-Fermi Lagrangian of the form $\overline{\chi}\gamma^\mu\gamma^5\chi\overline{q}\gamma_\mu\gamma^5q$) and spin-independent interactions ($\overline{\chi}\chi\overline{q}\,q$). For calculational details see GRIEST 88B, ELLIS 88D, BAR-BIERI 89c, DREES 93B, ARNOWITT 96, BERGSTROM 96, and BAER 97 in addition to the theory papers listed in the Tables. For a description of the theoretical assumptions and experimental techniques underlying most of the listed papers, see the review on "Dark matter" in this "Review of Particle Physics," and references therein. Most of the following papers use galactic halo and nuclear interaction assumptions from (LEWIN 96).

Spin-dependent interactions

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• • • We do not use the following data for averages, fits, limits, etc. • • •

< 5 20 AKERIB 06 CDMS Ge

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<sup>21</sup> SHIMIZU
                                                           06A CNTR CaF<sub>2</sub>
< 2
                                      <sup>22</sup> ALNER
                                                                 NAIA Nal Spin Dep.
< 0.4
                                     <sup>23</sup> BARNABE-HE..05
< 2
                                                                PICA
                                     <sup>24</sup> GIRARD
                                                               SMPL F, CI
< 1.4
                                     <sup>25</sup> KLAPDOR-K... 05 HDMS Ge
< 4
        \times 10^{-11} to 1 \times 10^{-4}
                                     <sup>26</sup> ELLIS
                                                                THEO \mu > 0
                                     <sup>27</sup> GIULIANI
                                                          04 SIMP
< 16
                                     <sup>28</sup> AHMED
                                                          03 NAIA
< 0.8
                                                                          Nal Spin Dep.
                                     <sup>29</sup> TAKEDA
                                                          03 BOLO NaF Spin Dep.
< 40
                                     <sup>30</sup> ANGLOHER
< 10
                                                          02 CRES Saphire
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                                                           01C THEO tan \beta < 10
                                     <sup>32</sup> BERNABEI
                                                          00D DAMA Xe
< 3.8
                                     <sup>33</sup> COLLAR
                                                                 SMPL F
< 15
                                         SPOONER
                                                          00 UKDM NaI
< 0.8
                                     <sup>34</sup> BELLI
                                                           99C DAMA F
< 4.8
                                     <sup>35</sup> OOTANI
<100
                                                                BOLO LiF
                                         BERNABEI
< 0.6
                                                           98C DAMA Xe
                                     <sup>34</sup> BERNABEI
                                                          97
                                                                DAMA F
<
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²⁰ The strongest upper limit is 4 pb and occurs at $m_{\chi} \simeq 60$ GeV. The limit on the neutron spin-dependent elastic cross section is 0.07 pb.

 $^{^{21}}$ The strongest upper limit is 1.2 pb and occurs at $m_\chi\simeq 40$ GeV. The limit on the neutron spin-dependent cross section is 35 pb.

 $^{^{22}\,\}mathrm{The}$ strongest upper limit is 0.35 pb and occurs at $m_\chi\,\simeq\,$ 60 GeV.

 $^{^{23}\,\}mathrm{The}$ strongest upper limit is 1.2 pb and occurs $m_\chi \simeq 30$ GeV.

²⁴ The strongest upper limit is 1.2 pb and occurs $m_{\chi}^{\sim} \simeq 40$ GeV.

 $^{^{\}rm 25}\,{\rm Limit}$ applies to neutron elastic cross section.

 $^{^{26}}$ ELLIS 04 calculates the $\chi-p$ elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry, but without universal scalar masses. In the case of universal squark and slepton masses, but non-universal Higgs masses, the limit becomes 2×10^{-4} , see ELLIS 03E.

²⁷ The strongest upper limit is 10 pb and occurs at $m_{\chi} \simeq$ 30 GeV.

 $^{^{28}\,\}mathrm{The}$ strongest upper limit is 0.75 pb and occurs at $\overset{\curvearrowright}{m_\chi}\approx$ 70 GeV.

 $^{^{29}}$ The strongest upper limit is 30 pb and occurs at $m_{\chi}^{\sim} \approx 20$ GeV.

 $^{^{31}}$ ELLIS 01C calculates the χ -p elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. In models with nonuniversal Higgs masses, the upper limit to the cross section is 6×10^{-4} .

 $^{^{32}}$ The strongest upper limit is 3 pb and occurs at $m_\chi \simeq 60$ GeV. The limits are for inelastic scattering $\chi^0 + ^{129}$ Xe $\to \chi^0 + ^{129}$ Xe* (39.58 keV).

 $^{^{33}\,\}mathrm{The}$ strongest upper limit is 9 pb and occurs at $m_\chi\simeq 30$ GeV.

 $^{^{34}\,\}mathrm{The}$ strongest upper limit is 4.4 pb and occurs at $m_\chi \simeq 60$ GeV.

 $^{^{35}}$ The strongest upper limit is about 35 pb and occurs at $m_\chi \simeq 15$ GeV.

Spin-independent interactions

• • • We do not use the following data for averages, fits, limits, etc. • • • • • • • • • • • • • • • • • • •	VALUE (pb)	DOCUMENT	· ID	TECN	COMMENT
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$< 1.4 \times 10^{-6}$	SANGLAR		EDEL	Ge
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$< 4 \times 10^{-7}$	⁴¹ AKERIB	04	CDMS	Ge
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$2 \times 10^{-11} \text{ to } 8 \times 10^{-6} 42$	^{1,43} ELLIS	04	THEO	$\mu > 0$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$< 5 \times 10^{-8}$	⁴⁴ PIERCE	04A	THEO	
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$ \begin{array}{cccccccccccccccccccccccccccccccccccc$				THEO	
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$		⁴⁹ ABRAMS	02	CDMS	Ge
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$< 1.4 \times 10^{-6}$	⁵⁰ BENOIT	02	EDEL	Ge
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		⁴² KIM		THEO	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		⁵¹ MORALES	02 B	CSME	Ge
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	\10	⁵² MORALES	020	IGEX	Ge
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$< 7 imes 10^{-6}$ BERNABEI 98C DAMA Xe		BERNABE	I 99		
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36 ALVEDID ACCOUNTS IN THE CALVEDID AS THE COUNTY IN THE CALVEDID ACCOUNTS IN THE CALVED ACCOUNTS ACCO					Xe

 $^{^{36}}$ AKERIB 06A updates the results of AKERIB 05. The strongest upper limit is 1.6 imes 10^{-7} pb and occurs at $m_\chi \approx 60$ GeV. 37 The strongest upper limit is 8×10^{-6} pb and occurs at $m_\chi \simeq 70$ GeV.

 $^{^{38}}$ AKERIB 05 is incompatible with the DAMA most likely value. The strongest upper limit is 4 \times 10 $^{-7}$ pb and occurs at $m_\chi~\simeq~60$ GeV.

The strongest upper limit is also close to 1.0×10^{-6} pb and occurs at $m_\chi\simeq70$ GeV. BENOIT 06 claim that the discrimination power of ZEPLIN-I measurement (ALNER 05A) is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06 is not reliable enough to obtain a limit better than 1×10^{-3} pb. do not agree with the criticisms of BENOIT 06.

- ⁴⁰ The strongest upper limit is also close to 1.4×10^{-6} pb and occurs at $m_{_Y} \simeq 70$ GeV.
- 41 AKERIB 04 is incompatible with BERNABEI 00 most likely value, under the assumption of standard WIMP-halo interactions. The strongest upper limit is 4×10^{-7} pb and occurs at $m_{_Y} \simeq 60$ GeV.
- 42 KIM 02 and ELLIS 04 calculate the χ p elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry, but without universal scalar masses.
- 43 In the case of universal squark and slepton masses, but non-universal Higgs masses, the limit becomes 2×10^{-6} (2×10^{-11} when constraint from the BNL g-2 experiment are included), see ELLIS 03E. ELLIS 05 display the sensitivity of the elastic scattering cross section to the π -Nucleon Σ term.
- 44 PIERCE 04A calculates the $\chi-p$ elastic scattering cross section in the framework of models with very heavy scalar masses. See Fig. 2 of the paper.
- 45 The strongest upper limit is 1.8×10^{-5} pb and occurs at $m_{_Y} \approx 80$ GeV.
- $^{
 m 46}$ Under the assumption of standard WIMP-halo interactions, Akerib 03 is incompatible with BERNABEI 00 most likely value at the 99.98% CL. See Fig. 4.
- 47 BAER 03A calculates the $\chi-p$ elastic scattering cross section in several models including the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- ⁴⁸ The strongest upper limit is 7×10^{-6} pb and occurs at $m_\chi \simeq 30$ GeV.
- $^{49}\,\mathrm{ABRAMS}$ 02 is incompatible with the DAMA most likely value at the 99.9% CL. The strongest upper limit is 3×10^{-6} pb and occurs at $m_{\gamma} \simeq 30$ GeV.
- 50 BENOIT 02 excludes the central result of DAMA at the 99.8%CL. 51 The strongest upper limit is 2 \times 10 $^{-5}$ pb and occurs at $m_\chi \simeq$ 40 GeV.
- 52 The strongest upper limit is 7×10^{-6} pb and occurs at $m_\chi^\sim \simeq$ 46 GeV.
- 53 The strongest upper limit is 1.8×10^{-5} pb and occurs at $\overset{\smallfrown}{m_\chi} \simeq$ 32 GeV
- 54 BOTTINO 01 calculates the χ -p elastic scattering cross section in the framework of the following supersymmetric models: N=1 supergravity with the radiative breaking of the electroweak gauge symmetry, N=1 supergravity with nonuniversal scalar masses and an effective MSSM model at the electroweak scale.
- ⁵⁵ Calculates the χ -p elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 56 ELLIS 01C calculates the χ -p elastic scattering cross section in the framework of $N\!\!=\!\!1$ supergravity models with radiative breaking of the electroweak gauge symmetry. EL-LIS 02B find a range 2×10^{-8} – 1.5×10^{-7} at $\tan \beta$ =50. In models with nonuniversal Higgs masses, the upper limit to the cross section is 4×10^{-7} .
- 57 ACCOMANDO 00 calculate the χ -p elastic scattering cross section in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. The limit is relaxed by at least an order of magnitude when models with nonuniversal scalar masses are considered. A subset of the authors in ARNOWITT 02 updated the limit to $< 9 \times 10^{-8}$ (tan $\beta < 55$).
- $^{58}\,\mathrm{BERNABEI}$ 00 search for annual modulation of the WIMP signal. The data favor the hypothesis of annual modulation at 4σ and are consistent, for a particular model framework quoted there, with $m_{\chi^0}=44^{+12}_{-9}$ GeV and a spin-independent χ^0 -proton cross section of (5.4 \pm 1.0) \times 10⁻⁶ pb. See also BERNABEI 01 and BERNABEI 00C.
- 59 FENG 00 calculate the χ -p elastic scattering cross section in the framework of $N\!\!=\!\!1$ supergravity models with radiative breaking of the electroweak gauge symmetry with a particular emphasis on focus point models. At $tan\beta=50$, the range is 8×10^{-8} – 4×10^{-7} .
- $^{60}\,\mathsf{BERNABEI}$ 99 search for annual modulation of the WIMP signal. The data favor the hypothesis of annual modulation at 99.6%CL and are consistent, for the particular model framework considered there, with $m_{\chi^0} = 59^{+17}_{-14}$ GeV and spin-independent χ^0 -proton cross section of $(7.0^{+0.4}_{-1.2}) \times 10^{-6}$ pb $(1 \sigma \text{ errors})$.

 61 BERNABEI 98 search for annual modulation of the WIMP signal. The data are consistent, for the particular model framework considered there, with $m_{\chi 0} = 59 ^{+36}_{-19}$ GeV and spin-independent χ^0 -proton cross section of $(1.0 ^{+0.1}_{-0.4}) \times 10^{-5}$ pb (1 σ errors).

- Other bounds on $\widetilde{\chi}^0_1$ from astrophysics and cosmology —

Most of these papers generally exclude regions in the M_2 – μ parameter plane by requiring that the $\widetilde{\chi}^0_1$ contribution to the overall cosmological density is less than some maximal value to avoid overclosure of the Universe. Those not based on the cosmological density are indicated. Many of these papers also include LEP and/or other bounds.

VALUE	DOCUMENT ID		TECN	COMMENT
>46 GeV	⁶² ELLIS	00	RVUE	
• • • We do not use	the following data for a	verage	es, fits, l	imits, etc. • • •
> 6 GeV	63,64 BELANGER	04	THEO	
	⁶⁵ ELLIS	04 B	COSM	
	⁶⁶ PIERCE	04A	COSM	
	⁶⁷ BAER	03	COSM	
> 6 GeV	⁶³ BOTTINO	03	COSM	
	⁶⁷ CHATTOPAD.	03	COSM	
	⁶⁸ ELLIS	03	COSM	
	⁶⁹ ELLIS	03 B	COSM	
	⁶⁷ ELLIS	03 C	COSM	
> 18 GeV	⁶³ HOOPER	03	COSM	$arOmega_\chi = 0.05 – 0.3$
	⁶⁷ LAHANAS	03	COSM	X
	⁷⁰ BAER	02	COSM	
	⁷¹ ELLIS	02	COSM	
	⁷² LAHANAS	02	COSM	
	⁷³ BARGER	01 C	COSM	
	⁷⁰ DJOUADI	01	COSM	
	⁷⁴ ELLIS		COSM	
	⁷⁰ ROSZKOWSK	l 01	COSM	
	⁶⁸ ВОЕНМ	00 B	COSM	
	⁷⁵ FENG	00	COSM	
	⁷⁶ LAHANAS	00	COSM	
$< 600 \; \text{GeV}$	77 ELLIS	98 B	COSM	
	⁷⁸ EDSJO	97	COSM	Co-annihilation
	⁷⁹ BAER	96	COSM	
	⁸⁰ BEREZINSKY	95	COSM	
	⁸¹ FALK	95	COSM	CP-violating phases
	82 DREES	93		Minimal supergravity
	⁸³ FALK	93	COSM	Sfermion mixing
	82 KELLEY	93		Minimal supergravity
	⁸⁴ MIZUTA	93		Co-annihilation
	⁸⁵ LOPEZ	92	COSM	Minimal supergravity, $m_0 = A = 0$
	⁸⁶ MCDONALD	92	COSM	U
	87 GRIEST	91	COSM	

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<sup>88</sup> NOJIRI
                                                                      COSM Minimal supergravity
                                     <sup>89</sup> OLIVE
                                                                      COSM
                                     <sup>90</sup> ROSZKOWSKI 91
                                                                      COSM
                                     <sup>91</sup> GRIEST
                                                                      COSM
                                     <sup>89</sup> OLIVE
                                                                      COSM
                                         SREDNICKI
none 100 eV - 15 GeV
                                                               88
                                                                      COSM \tilde{\gamma}; m_{\tilde{f}} = 100 \text{ GeV}
none 100 eV-5 GeV
                                                               84
                                                                      COSM \widetilde{\gamma}; for m_{\widetilde{f}} = 100 \text{ GeV}
                                         ELLIS
                                         GOLDBERG
                                                                       COSM \tilde{\gamma}
                                                               83
                                     <sup>92</sup> KRAUSS
                                                               83
                                                                      COSM \tilde{\gamma}
                                         VYSOTSKII
                                                                      COSM \tilde{\gamma}
                                                               83
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- 62 ELLIS 00 updates ELLIS 98. Uses LEP $e^+\,e^-$ data at $\sqrt{s}{=}202$ and 204 GeV to improve bound on neutralino mass to 51 GeV when scalar mass universality is assumed and 46 GeV when Higgs mass universality is relaxed. Limits on $\tan\beta$ improve to > 2.7 $(\mu > 0), >$ 2.2 $(\mu < 0)$ when scalar mass universality is assumed and > 1.9 (both signs of μ) when Higgs mass universality is relaxed.
- 63 HOOPER 03, BOTTINO 03 (see also BOTTINO 03A and BOTTINO 04) , and BE-LANGER 04 do not assume gaugino or scalar mass unification.
- ⁶⁴ Limit assumes a pseudo scalar mass < 200 GeV. For larger pseudo scalar masses, $m_{\chi} > 18(29)$ GeV for $\tan\beta = 50(10)$. Bounds from WMAP, $(g-2)_{\mu}$, $b \rightarrow s\gamma$, LEP.
- ⁶⁵ ELLIS 04B places constraints on the SUSY parameter space in the framework of *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry including supersymmetry breaking relations between A and B parameters. See also ELLIS 03D.
- ⁶⁶ PIERCE 04A places constraints on the SUSY parameter space in the framework of models with very heavy scalar masses.
- 67 BAER 03, CHATTOPADHYAY 03, ELLIS 03C and LAHANAS 03 place constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry based on WMAP results for the cold dark matter density.
- 68 BOEHM 00B and ELLIS 03 place constraints on the SUSY parameter space in the framework of minimal $N\!\!=\!\!1$ supergravity models with radiative breaking of the electroweak gauge symmetry. Includes the effect of $\chi\!\!-\!\!\tilde{t}$ co-annihilations.
- 69 BEREZINSKY 95 and ELLIS 03B places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry but non-Universal Higgs masses.
- ⁷⁰ DJOUADI 01, ROSZKOWSKI 01, and BAER 02 place constraints on the SUSY parameter space in the framework of minimal *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- ⁷¹ ELLIS 02 places constraints on the soft supersymmetry breaking masses in the framework of minimal *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 72 LAHANAS 02 places constraints on the SUSY parameter space in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. Focuses on the role of pseudo-scalar Higgs exchange.
- 73 BARGER 01C use the cosmic relic density inferred from recent CMB measurements to constrain the parameter space in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- ⁷⁴ ELLIS 01B places constraints on the SUSY parameter space in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. Focuses on models with large $\tan \beta$.
- 75 FENG 00 explores cosmologically allowed regions of MSSM parameter space with multi-
- 76 LAHANAS 00 use the new cosmological data which favor a cosmological constant and its implications on the relic density to constrain the parameter space in the framework

- of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 77 ELLIS 98B assumes a universal scalar mass and radiative supersymmetry breaking with universal gaugino masses. The upper limit to the LSP mass is increased due to the inclusion of $\chi-\widetilde{\tau}_R$ coannihilations.
- ⁷⁸ EDSJO 97 included all coannihilation processes between neutralinos and charginos for any neutralino mass and composition.
- 79 Notes the location of the neutralino Z resonance and h resonance annihilation corridors in minimal supergravity models with radiative electroweak breaking.
- 80 BEREZINSKY 95 and ELLIS 02C places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry but non-Universal Higgs masses.
- 81 Mass of the bino (=LSP) is limited to $m_{\widetilde{R}} \lesssim$ 350 GeV for $m_t =$ 174 GeV.
- 82 DREES 93, KELLEY 93 compute the cosmic relic density of the LSP in the framework of minimal *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- ⁸³ FALK 93 relax the upper limit to the LSP mass by considering sfermion mixing in the MSSM.
- 84 MIZUTA 93 include coannihilations to compute the relic density of Higgsino dark matter.
- $^{85}\, \text{LOPEZ}$ 92 calculate the relic LSP density in a minimal SUSY GUT model.
- ⁸⁶ MCDONALD 92 calculate the relic LSP density in the MSSM including exact tree-level annihilation cross sections for all two-body final states.
- ⁸⁷ GRIEST 91 improve relic density calculations to account for coannihilations, pole effects, and threshold effects.
- ⁸⁸ NOJIRI 91 uses minimal supergravity mass relations between squarks and sleptons to narrow cosmologically allowed parameter space.
- ⁸⁹ Mass of the bino (=LSP) is limited to $m_{\widetilde{B}}\lesssim 350$ GeV for $m_t\leq 200$ GeV. Mass of the higgsino (=LSP) is limited to $m_{\widetilde{H}}\lesssim 1$ TeV for $m_t\leq 200$ GeV.
- $^{90}\,\mathrm{ROSZKOWSKI}$ 91 calculates LSP relic density in mixed gaugino/higgsino region.
- ⁹¹ Mass of the bino (=LSP) is limited to $m_{\widetilde{B}} \lesssim 550$ GeV. Mass of the higgsino (=LSP) is limited to $m_{\widetilde{H}} \lesssim 3.2$ TeV.
- 92 KRAUSS 83 finds $m_{\widetilde{\gamma}}$ not 30 eV to 2.5 GeV. KRAUSS 83 takes into account the gravitino decay. Find that limits depend strongly on reheated temperature. For example a new allowed region $m_{\widetilde{\gamma}}=$ 4–20 MeV exists if $m_{\rm gravitino}$ $\,$ <40 TeV. See figure 2.

Unstable $\widetilde{\chi}_1^0$ (Lightest Neutralino) MASS LIMIT

Unless otherwise stated, results in this section assume spectra and production rates as evaluated in the MSSM. Unless otherwise stated, the goldstino or gravitino mass $m_{\widetilde{G}}$ is assumed to be negligible relative to all other masses. In the following, \widetilde{G} is assumed to be undetected and to give rise to a missing energy (E) signature.

VALUE (GeV) CL%	DOCUMENT ID		TECN	COMMENT
• • • We do not use	e the following dat	a for a	averages,	fits, limits, etc. ● ●
	⁹³ ABAZOV	06 D	D0	Ŗ, LL E
	⁹⁴ ABAZOV	06 P	D0	R, λ_{122}

> 96.8 >108	95 95	⁹⁵ ABBIENDI ⁹⁶ ABAZOV		OPAL D0	$e^+e^- ightarrow \widetilde{B}\widetilde{B}, (\widetilde{B} ightarrow \widetilde{G}\gamma)$ $p\overline{p} ightarrow \widetilde{\chi}\widetilde{\chi}, \widetilde{\chi} = \widetilde{\chi}_2^0, \widetilde{\chi}_1^{\pm}, \widetilde{\chi}_1^0 ightarrow \gamma \widetilde{G},$
		⁹⁷ ABDALLAH	ΛEΡ	UI DH	$\begin{array}{ccc} GMSB & & & \\ e^{+}e^{-} & \to & \widetilde{G}\widetilde{\chi}_{1}^{0},(\widetilde{\chi}_{1}^{0} \to & \widetilde{G}\gamma) \end{array}$
> 06	OF	98 ABDALLAH	05В	DLPH	
> 96	95	99 ACOSTA			$e^+e^- \rightarrow \widetilde{B}\widetilde{B}, (\widetilde{B} \rightarrow \widetilde{G}\gamma)$
> 93	95	33 ACOSTA	05E	CDF	$p\overline{p} \rightarrow \widetilde{\chi}\widetilde{\chi}, \ \widetilde{\chi} = \widetilde{\chi}_2^0, \ \widetilde{\chi}_1^{\pm}, \widetilde{\chi}_1^0 \rightarrow \gamma \widetilde{G},$
		¹⁰⁰ AKTAS	05	H1	$e^{\pm} \stackrel{GMSB}{p \to q} \widetilde{\chi}^0_1, \ \widetilde{\chi}^0_1 \to \gamma \widetilde{G},$
					$GMSB + R LQ\overline{D}$
		101 ABBIENDI	04N	OPAL	$e^+e^- o \gamma \gamma \not\!\! E$
> 66	95 102	2,103 ABDALLAH		DLPH	AMSB, $\mu > 0$
> 38.0	95 104	^{1,105} ABDALLAH		DLPH	$R(\overline{UDD})$
		¹⁰⁶ ACHARD	04E	L3	$e^+e^- ightarrow \ \widetilde{G} \widetilde{\chi}^0_1, \widetilde{\chi}^0_1 ightarrow \ \widetilde{G} \gamma$
> 99.5	95	¹⁰⁷ ACHARD	04E	L3	$e^+e^- ightarrow \ \widetilde{B}\widetilde{B},(\widetilde{B} ightarrow \ \widetilde{G}\gamma)$
> 89		¹⁰⁸ ABDALLAH	03 D	DLPH	$e^+e^- ightarrow~\widetilde{\chi}^0_1\widetilde{\chi}^0_1$, GMSB,
					$m(\widetilde{G})<1\mathrm{eV}$
		¹⁰⁹ HEISTER	03 C	ALEP	$e^+e^- \rightarrow \widetilde{B}\widetilde{B}, (\widetilde{B} \rightarrow \gamma \widetilde{G})$
		¹¹⁰ HEISTER	03 C	ALEP	$e^+e^- ightarrow \ \widetilde{G}\widetilde{\chi}^0_1$, $(\widetilde{\chi}^0_1 ightarrow \ \widetilde{G}\gamma)$
> 39.9	95	¹¹¹ ACHARD	02	L3	R, MSUGRA
> 92	95	112 HEISTER	02 R	ALEP	short lifetime
> 54	95	112 HEISTER	02 R	ALEP	any lifetime
> 85	95	¹¹³ ABBIENDI	01	OPAL	${ m e^+e^-} ightarrow~\widetilde{\chi}_1^0\widetilde{\chi}_1^0$, GMSB, tan $eta=2$
> 76	95	¹¹³ ABBIENDI	01	OPAL	$e^+e^- ightarrow~\widetilde{\chi}_1^{ar{0}}\widetilde{\chi}_1^{ar{0}}$, GMSB, tan $eta=$ 20
> 32.5	95	¹¹⁴ ACCIARRI	01	L3	R , all m_0 , $0.7 \le \tan \beta \le 40$
		¹¹⁵ ADAMS	01	NTEV	$\widetilde{\chi}^0 ightarrow \mu \mu u$, R , $LL\overline{E}$
> 29	95	¹¹⁶ ABBIENDI	99T	OPAL	$e^+e^- ightarrow \widetilde{\chi}_1^0 \widetilde{\chi}_1^0$, R , $m_0=500$ GeV, $\tan \beta > 1.2$
		¹¹⁷ ACCIARRI	99R	L3	Superseded by ACHARD 04E
> 88.2	95	¹¹⁸ ACCIARRI	99R		Superseded by ACHARD 04E
> 29	95	¹¹⁹ BARATE		ALEP	R , $LQ\overline{D}$, $tan\beta=1.41$, $m_0=500$ GeV
		¹²⁰ ABREU	98	DLPH	$e^+e^- \rightarrow \widetilde{\chi}_1^0\widetilde{\chi}_1^0 (\widetilde{\chi}_1^0 \rightarrow \gamma \widetilde{G})$
> 23	95	¹²¹ BARATE	985	ALEP	R , $LL\overline{E}$
		¹²² ELLIS	97	THEO	$e^+e^- ightarrow \widetilde{\chi}^0_1\widetilde{\chi}^0_1$, $\widetilde{\chi}^0_1 ightarrow \gamma\widetilde{G}$
		¹²³ CABIBBO	81	COSM	7171. 71

 $^{^{93}}$ ABAZOV 06D looked in 360 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with three leptons originating from the pair production of charginos and neutralinos, followed by R decays mediated by $LL\overline{E}$ couplings. One coupling is assumed to be dominant at a time. No significant excess was found compared to the background expectation in the $e\,e\,\ell,\,\mu\mu\ell$ nor $e\,e\,\tau$ ($\ell=e,\,\mu$) final states. Upper limits on the cross-section are extracted in a specific MSUGRA model and a MSSM model without unification of M_1 and M_2 at the GUT scale. A limit is derived on the masses of charginos and neutralinos for both scenarios assuming λ_{ijk} couplings such that the decay length is less than 1 cm, see their Table III and Fig. 4.

 $^{^{94}}$ ABAZOV 06P looked in 380 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with at least 2 opposite sign isolated muons which might arise from the decays of neutralinos into $\mu\mu\nu$ via R couplings $LL\overline{E}$. No events are observed in the decay region defined by a radius between 5 and 20 cm, in agreement with the SM expectation. Limits are set on the cross-section times branching ratio as a function of lifetime, shown in their Fig. 3. This limit excludes the SUSY interpretation of the NuTeV excess of dimuon events reported in ADAMS 01.

- 95 ABBIENDI 06B use 600 pb $^{-1}$ of data from $\sqrt{s}=$ 189–209 GeV. They look for events with diphotons $+\not\!\! E$ final states originating from prompt decays of pair-produced neutralinos in a GMSB scenario with $\widetilde{\chi}^0_1$ NLSP. Limits on the cross-section are computed as a function of m($\widetilde{\chi}^0_1$), see their Fig. 14. The limit on the $\widetilde{\chi}^0_1$ mass is for a pure Bino state assuming a prompt decay, with lifetimes up to $10^{-9} s$. Supersedes the results of ABBIENDI 04N.
- 96 ABAZOV 05A looked in 263 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for diphoton events with large $\not\!\!E_T$. They may originate from the production of $\widetilde{\chi}^\pm$ in pairs or associated to a $\widetilde{\chi}^0_2$, decaying to a $\widetilde{\chi}^0_1$ which itself decays promptly in GMSB to $\widetilde{\chi}^0_1 \to \gamma \, \widetilde{G}$. No significant excess was found at large $\not\!\!E_T$ compared to the background expectation. A limit is derived on the masses of SUSY particles in the GMSB framework for M=2 Λ , N=1, $\tan\beta=15$ and $\mu>0$, see Figure 2. It also excludes $\Lambda<79.6$ TeV. Very similar results are obtained for different choices of parameters, see their Table 2. Supersedes the results of ABBOTT 98.
- 97 ABDALLAH 05B use data from $\sqrt{s}=180$ –209 GeV. They look for events with single photons + $\cancel{\mathbb{F}}$ final states. Limits are computed in the plane (m(\widetilde{G}), m($\widetilde{\chi}_1^0$)), shown in their Fig. 9b for a pure Bino state in the GMSB framework and in Fig. 9c for a no-scale supergravity model. Supersedes the results of ABREU 00Z.
- 98 ABDALLAH 05B use data from $\sqrt{s}=130$ –209 GeV. They look for events with diphotons $+\not\!\!E$ final states and single photons not pointing to the vertex, expected in GMSB when the $\widetilde{\chi}_1^0$ is the NLSP. Limits are computed in the plane $(\mathsf{m}(\widetilde{G}),\,\mathsf{m}(\widetilde{\chi}_1^0))$, see their Fig. 10. The lower limit is derived on the $\widetilde{\chi}_1^0$ mass for a pure Bino state assuming a prompt decay and $m_{\widetilde{e}_R}=m_{\widetilde{e}_L}=2$ $m_{\widetilde{\chi}_1^0}$. It improves to 100 GeV for $m_{\widetilde{e}_R}=m_{\widetilde{e}_L}=1.1$ $m_{\widetilde{\chi}_1^0}$. and
 - the limit in the plane $(m(\tilde{\chi}_1^U), m(\tilde{e}_R))$ is shown in Fig. 10b. For long-lived neutralinos, cross-section limits are displayed in their Fig 11. Supersedes the results of ABREU 00Z.
- ACOSTA 05E looked in 202 pb $^{-1}$ of $p\overline{p}$ collisions at \sqrt{s} =1.96 TeV for diphoton events with large $\not\!\!E_T$. They may originate from the production of $\widetilde{\chi}^\pm$ in pairs or associated to a $\widetilde{\chi}^0_2$, decaying to a $\widetilde{\chi}^0_1$ which itself decays promptly in GMSB to $\gamma \, \widetilde{G}$. No events are selected at large $\not\!\!E_T$ compared to the background expectation. A limit is derived on the masses of SUSY particles in the GMSB framework for M=2 Λ , N=1, $\tan\beta=15$ and $\mu>0$, see Figure 2. It also excludes $\Lambda<69$ TeV. Supersedes the results of ABE 99I.
- AKTAS 05 data collected at 319 GeV with 64.3 pb $^{-1}$ of e^+p and 13.5 pb $^{-1}$ of e^-p . They look for $\not\!\!R$ resonant $\widetilde\chi^0_1$ production via t-channel exchange of a $\widetilde e$, followed by prompt GMSB decay of the $\widetilde\chi^0_1$ to $\gamma\,\widetilde G$. Upper limits at 95% on the cross section are derived, see their Figure 4, and compared to two example scenarios. In Figure 5, they display 95% exclusion limits in the plane of $M(\widetilde\chi^0_1)$ versus $M(\widetilde e_L)-M(\widetilde\chi^0_1)$ for the two scenarios and several values of the λ' Yukawa coupling.
- 101 ABBIENDI 04N use data from $\sqrt{s}=189$ –209 GeV, setting limits on $\sigma(e^+e^-\to XX)\times B^2(X\to Y\gamma)$, with Y invisible (see their Fig. 4). Limits on $\widetilde{\chi}_1^0$ masses for a specific model are given. Supersedes the results of ABBIENDI,G 00D.
- 102 ABDALLAH 04H use data from LEP 1 and $\sqrt{s}=192$ –208 GeV. They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region $1 < m_{3/2} < 50$ TeV, $0 < m_0 < 1000$ GeV, $1.5 < \tan\beta < 35$, both signs of μ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z width of 3.2 MeV. The limit is for $m_t=174.3$ GeV (see Table 2 for other m_t values).
- 103 The limit improves to 73 GeV for $\mu < 0$.
- 104 ABDALLAH 04M use data from $\sqrt{s}=192-208$ GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $\overline{U}\overline{D}\overline{D}$ couplings. The results are valid in the ranges 90< m_0 <500 GeV, 0.7<tan β <30, $-200<\mu$ <200 GeV, 0< M_2 <400 GeV. Supersedes the result of ABREU 01D and ABREU 00U.

- 105 The limit improves to 39.5 GeV for $LL\overline{E}$ couplings.
- ACHARD 04E use data from $\sqrt{s}=189$ –209 GeV. They look for events with single photons $+\not\!\! E$ final states. Limits are computed in the plane (m(\widetilde{G}), m($\widetilde{\chi}_1^0$)), shown in their Fig. 8c for a no-scale supergravity model, excluding, e.g., Gravitino masses below 10^{-5} eV for neutralino masses below 172 GeV. Supersedes the results of ACCIARRI 99R.
- 107 ACHARD 04E use data from $\sqrt{s}=189$ –209 GeV. They look for events with diphotons $+\not\!\!E$ final states. Limits are computed in the plane $(\mathsf{m}(\widetilde{\chi}_1^0),\,\mathsf{m}(\widetilde{e}_R)),$ see their Fig. 8d. The limit on the $\widetilde{\chi}_1^0$ mass is for a pure Bino state assuming a prompt decay, with $m_{\widetilde{e}_L}=1.1~m_{\widetilde{\chi}_1^0}$ and $m_{\widetilde{e}_R}=2.5~m_{\widetilde{\chi}_1^0}.$ Supersedes the results of ACCIARRI 99R.
- ABDALLAH 03D use data from $\sqrt{s}=161$ –208 GeV. They look for 4-tau + $\not\!\!E$ final states, expected in GMSB when the $\widetilde{\tau}_1$ is the NLSP, and 4-lepton + $\not\!\!E$ final states, expected in the co-NLSP scenario, and assuming a short-lived $\widetilde{\chi}_1^0$ (m(\widetilde{G})<1 eV). Limits are computed in the plane (m($\widetilde{\tau}_1$), m($\widetilde{\chi}_1^0$)) from a scan of the GMSB parameters space, after combining these results with the search for slepton pair production from the same paper to cover prompt decays and for the case of $\widetilde{\chi}_1^0$ NLSP from ABREU 00Z. The limit above is reached for a single generation of messengers and when the $\widetilde{\tau}_1$ is the NLSP. Stronger limits are obtained when more messenger generations are assumed or when the other sleptons are co-NLSP, see their Fig. 10. Supersedes the results of ABREU 01G.
- 109 HEISTER 03C use the data from $\sqrt{s}=189$ –209 GeV to search for $\gamma\not\!\!E_T$ final states with non-pointing photons and $\gamma\gamma\not\!\!E_T$ events. Interpreted in the framework of Minimal GMSB, a lower bound on the $\widetilde{\chi}_1^0$ mass is obtained as function of its lifetime. For a laboratory lifetime of less than 3 ns, the limit at 95% CL is 98.8 GeV. For other lifetimes, see their Fig. 5. These results are interpreted in a more general GMSB framework in HEISTER 02R.
- ¹¹⁰ HEISTER 03C use the data from $\sqrt{s}=189$ –209 GeV to search for $\gamma\not\!\!E_T$ final states. They obtained an upper bound on the cross section for the process $e^+e^-\to \widetilde G\widetilde\chi^0_1$, followed by the prompt decay $\widetilde\chi^0_1\to\gamma\widetilde G$, shown in their Fig. 4. These results supersede BARATE 98H.
- 111 ACHARD 02 searches for the production of sparticles in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit holds for \overline{UDD} couplings and increases to 40.2 GeV for $LL\overline{E}$ couplings. For L3 limits from $LQ\overline{D}$ couplings, see ACCIARRI 01.
- HEISTER 02R search for signals of GMSB in the 189–209 GeV data. For the $\widetilde{\chi}^0_1$ NLSP scenario, they looked for topologies consisting of $\gamma\gamma E$ or a single γ not pointing to the interaction vertex. For the $\widetilde{\ell}$ NLSP case, the topologies consist of $\ell\ell E$ or $4\ell E$ (from $\widetilde{\chi}^0_1\widetilde{\chi}^0_1$) production), including leptons with large impact parameters, kinks, or stable particles. Limits are derived from a scan over the GMSB parameters (see their Table 5 for the ranges). The limits are valid whichever is the NLSP. The absolute mass bound on the $\widetilde{\chi}^0_1$ for any lifetime includes indirect limits from the chargino search, and from the slepton search HEISTER 02E preformed within the MSUGRA framework. A bound for any NLSP and any lifetime of 77 GeV has also been derived by using the constraints from the neutral Higgs search in HEISTER 02. Limits on the universal SUSY mass scale Λ are also derived in the paper. Supersedes the results from BARATE 00G.
- ¹¹³ ABBIENDI 01 looked for final states with $\gamma\gamma \not\!\! E$, $\ell\ell \not\!\! E$, with possibly additional activity and four leptons $+\not\!\! E$ to search for prompt decays of $\widetilde{\chi}_1^0$ or $\widetilde{\ell}_1$ in GMSB. They derive limits in the plane $(m_{\widetilde{\chi}_1^0}, m_{\widetilde{\tau}_1})$, see Fig. 6, allowing either the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}_1$ to be the NLSP. Two

- scenarios are considered: $\tan\beta{=}2$ with the 3 sleptons degenerate in mass and $\tan\beta{=}20$ where the $\widetilde{\tau}_1$ is lighter than the other sleptons. Data taken at $\sqrt{s}{=}189$ GeV.
- ACCIARRI 01 searches for multi-lepton and/or multi-jet final states from R prompt decays with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings at $\sqrt{s}{=}189$ GeV. The search is performed for direct and indirect decays of neutralinos, charginos, and scalar leptons, with the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}$ as LSP and assuming one coupling to be nonzero at a time. Mass limits are derived using simultaneously the constraints from the neutralino, chargino, and slepton analyses; and the Z^0 width measurements from ACCIARRI 00C in a scan of the parameter space assuming MSUGRA with gaugino and scalar mass universality. Updates and supersedes the results from ACCIARRI 99I.
- ADAMS 01 looked for neutral particles with mass > 2.2 GeV, produced by 900 GeV protons incident on a Beryllium oxide target and decaying through weak interactions into $\mu\mu$, μe , or $\mu\pi$ final states in the decay channel of the NuTeV detector (E815) at Fermilab. The number of observed events is $3\,\mu\mu$, $0\,\mu e$, and $0\,\mu\pi$ with an expected background of 0.069 ± 0.010 , 0.13 ± 0.02 , and 0.14 ± 0.02 , respectively. The $\mu\mu$ events are consistent with the R decay of a neutralino with mass around 5 GeV. However, they share several aspects with ν -interaction backgrounds. An upper limit on the differential production cross section of neutralinos in pp interactions as function of the decay length is given in Fig. 3.
- ABBIENDI 99T searches for the production of neutralinos in the case of R-parity violation with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings using data from $\sqrt{s}{=}183$ GeV. They investigate topologies with multiple leptons, jets plus leptons, or multiple jets, assuming one coupling at the time to be non-zero and giving rise to direct or indirect decays. Mixed decays (where one particle has a direct, the other an indirect decay) are also considered for the \overline{UDD} couplings. Upper limits on the cross section are derived which, combined with the constraint from the Z^0 width, allow to exclude regions in the M_2 versus μ plane for any coupling. Limits on the neutralino mass are obtained for non-zero $LL\overline{E}$ couplings $> 10^{-5}$. The limit disappears for $\tan\beta < 1.2$ and it improves to 50 GeV for $\tan\beta > 20$.
- ¹¹⁷ ACCIARRI 99R searches for γE final states using data from \sqrt{s} =189 GeV. From limits on cross section times branching ratio, mass limits are derived in a no-scale SUGRA model, see their Fig. 5. Supersedes the results of ACCIARRI 98V.
- ¹¹⁸ ACCIARRI 99R searches for γE final states using data from \sqrt{s} =189 GeV. From a scan over the GMSB parameter space, a limit on the mass is derived under the assumption that the neutralino is the NLSP. Supersedes the results of ACCIARRI 98V.
- ¹¹⁹BARATE 99E looked for the decay of gauginos via *R*-violating couplings $LQ\overline{D}$. The bound is significantly reduced for smaller values of m_0 . Data collected at \sqrt{s} =130–172 GeV.
- ¹²⁰ ABREU 98 uses data at \sqrt{s} =161 and 172 GeV. Upper bounds on $\gamma\gamma E$ cross section are obtained. Similar limits on γE are also given, relevant for $e^+e^- \to \widetilde{\chi}_1^0 \widetilde{G}$ production.
- ¹²¹ BARATE 98S looked for the decay of gauginos via *R*-violating coupling $LL\overline{E}$. The bound improves to 25 GeV if the chargino decays into neutralino which further decays into lepton pairs. Data collected at \sqrt{s} =130–172 GeV.
- 122 ELLIS 97 reanalyzed the LEP2 (\sqrt{s} =161 GeV) limits of $\sigma(\gamma\gamma + E_{\rm miss}) < 0.2$ pb to exclude $m_{\widetilde{\chi}_1^0} <$ 63 GeV if $m_{\widetilde{e}_L} = m_{\widetilde{e}_R} <$ 150 GeV and $\widetilde{\chi}_1^0$ decays to γ \widetilde{G} inside detector.
- ¹²³ CABIBBO 81 consider $\widetilde{\gamma} \to \gamma + \text{goldstino}$. Photino must be either light enough (<30 eV) to satisfy cosmology bound, or heavy enough (>0.3 MeV) to have disappeared at early universe.

$\widetilde{\chi}^0_2,\,\widetilde{\chi}^0_3,\,\widetilde{\chi}^0_4$ (Neutralinos) MASS LIMITS

Neutralinos are unknown mixtures of photinos, z-inos, and neutral higgsinos (the supersymmetric partners of photons and of Z and Higgs bosons). The limits here apply only to $\widetilde{\chi}^0_2$, $\widetilde{\chi}^0_3$, and $\widetilde{\chi}^0_4$. $\widetilde{\chi}^0_1$ is the lightest supersymmetric particle (LSP); see $\widetilde{\chi}^0_1$

Mass Limits. It is not possible to quote rigorous mass limits because they are extremely model dependent; i.e. they depend on branching ratios of various $\widetilde{\chi}^0$ decay modes, on the masses of decay products $(\widetilde{e},\ \widetilde{\gamma},\ \widetilde{q},\ \widetilde{g})$, and on the \widetilde{e} mass exchanged in $e^+e^-\to \widetilde{\chi}^0_i\,\widetilde{\chi}^0_j$. Limits arise either from direct searches, or from the MSSM constraints set on the gaugino and higgsino mass parameters M_2 and μ through searches for lighter charginos and neutralinos. Often limits are given as contour plots in the $m_{\widetilde{\chi}^0}-m_{\widetilde{e}}$ plane vs other parameters. When specific assumptions are made, e.g, the neutralino is a pure photino $(\widetilde{\gamma})$, pure z-ino (\widetilde{Z}) , or pure neutral higgsino (\widetilde{H}^0) , the neutralinos will be labelled as such.

Limits obtained from e^+e^- collisions at energies up to 136 GeV, as well as other limits from different techniques, are now superseded and have not been included in this compilation. They can be found in the 1998 Edition (The European Physical Journal **C3** 1 (1998)) of this Review.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
> 78	95	124 ABBIENDI	04н	OPAL	$\widetilde{\chi}_2^0$, all tan β , $\Delta m_0 >$ 5 GeV,
> 62.4	95	¹²⁵ ABREU	00W I	DLPH	$m_0 >$ 500 GeV, $A_0 = 0$ $\widetilde{\chi}_2^0$, $1 \le \tan\beta \le 40$, all Δm_0 ,
> 99.9	95	¹²⁵ ABREU	00W I	DLPH	all m_0 $\widetilde{\chi}^0_3$, $1 \leq \tan\beta \leq 40$, all Δm_0 ,
>116.0	95	125 ABREU	00W I	DLPH	all m_0 $\widetilde{\chi}_4^0$, $1 \leq \tan\beta \leq 40$, all Δm_0 , all m_0

• • • We do not use the following data for averages, fits, limits, etc. • • •

		•		•	
		¹²⁶ ABDALLAH	05 B	DLPH	$e^+e^- \rightarrow \widetilde{\chi}_2^0\widetilde{\chi}_2^0, (\widetilde{\chi}_2^0 \rightarrow \widetilde{\chi}_1^0\gamma)$
		¹²⁷ ACHARD	04E	L3	$\begin{array}{ll} e^{+} e^{-} \rightarrow \ \widetilde{\chi}_{2}^{0} \widetilde{\chi}_{2}^{0}, \ (\widetilde{\chi}_{2}^{0} \rightarrow \ \widetilde{\chi}_{1}^{0} \gamma) \\ e^{+} e^{-} \rightarrow \ \widetilde{\chi}_{2}^{0} \widetilde{\chi}_{2}^{0}, \ (\widetilde{\chi}_{2}^{0} \rightarrow \ \widetilde{\chi}_{1}^{0} \gamma) \end{array}$
> 80.0	95	¹²⁸ ACHARD	02	L3	$\widetilde{\chi}_{2}^{0}$, R , MSUGRA
>107.2	95	¹²⁸ ACHARD	02	L3	$\widetilde{\chi}_{3}^{0}$, $\not R$, MSUGRA
		¹²⁹ ABREU	01 B	DLPH	$e^{+}e^{-} \rightarrow \widetilde{\chi}_{i}^{0}\widetilde{\chi}_{i}^{0}$
> 68.0	95	¹³⁰ ACCIARRI	01	L3	$\widetilde{\chi}_2^0$, R , all m_0 , $0.7 \le \tan \beta \le 40$
> 99.0	95	¹³⁰ ACCIARRI	01	L3	$\widetilde{\chi}_{3}^{\overline{0}}$, $\not \! R$, all m_{0} , $0.7 \leq aneta \leq 40$
> 50	95	¹³¹ ABREU	00 U	DLPH	7.2
		132 ABBIENDI		OPAL	$\begin{array}{l} 1 \leq \tan\beta \leq 30 \\ e^{+} e^{-} \rightarrow \widetilde{\chi}_{2}^{0} \widetilde{\chi}_{1}^{0} (\widetilde{\chi}_{2}^{0} \rightarrow \gamma \widetilde{\chi}_{1}^{0}) \\ e^{+} e^{-} \rightarrow \widetilde{\chi}_{2}^{0} \widetilde{\chi}_{2}^{0} (\widetilde{\chi}_{2}^{0} \rightarrow \gamma \widetilde{\chi}_{1}^{0}) \end{array}$
		133 ABBIENDI	99F	OPAL	
		134 ABBOTT		D0	$ \rho \overline{\rho} \rightarrow \widetilde{\chi}_{1}^{\pm} \widetilde{\chi}_{2}^{0} $
> 82.2	95	¹³⁵ ABE			$p\overline{p} \rightarrow \widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^{\overline{0}}$
> 92	95	¹³⁶ ACCIARRI	98F	L3	\widetilde{H}_2^0 , tan β =1.41, M_2 < 500 GeV
		¹³⁷ ACCIARRI	98V	L3	$e^{\stackrel{?}{+}}e^{-} \rightarrow \widetilde{\chi}_{2}^{0}\widetilde{\chi}_{1,2}^{0}$
					$(\widetilde{\chi}_2^0 \rightarrow \gamma \widetilde{\chi}_1^0)$
> 53	95	¹³⁸ BARATE	98н	ALEP	$e^+e^- \rightarrow \widetilde{\gamma}\widetilde{\gamma}(\widetilde{\gamma} \rightarrow \gamma\widetilde{H}^0)$
> 74	95	139 BARATE	98J	ALEP	
		¹⁴⁰ ABACHI	96	D0	$ ho \overline{ ho} ightarrow \ \widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0$
		¹⁴¹ ABE	96K	CDF	$p\overline{p} \rightarrow \widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^{\overline{0}}$

- ABBIENDI 04H search for charginos and neutralinos in events with acoplanar leptons+jets and multi-jet final states in the 192–209 GeV data, combined with the results on leptonic final states from ABBIENDI 04. The results hold for a scan over the parameter space covering the region 0 < M_2 <5000 GeV, $-1000 < \mu < 1000$ GeV and $\tan\beta$ from 1 to 40. This limit supersedes ABBIENDI 00H.
- ABREU 00W combines data collected at $\sqrt{s}{=}189$ GeV with results from lower energies. The mass limit is obtained by constraining the MSSM parameter space with gaugino and sfermion mass universality at the GUT scale, using the results of negative direct searches for neutralinos (including cascade decays and $\widetilde{\tau}\tau$ final states) from ABREU 01, for charginos from ABREU 00J and ABREU 00T (for all Δm_{+}), and for charged sleptons from ABREU 01B. The results hold for the full parameter space defined by all values of M_2 and $|\mu| \leq 2$ TeV with the $\widetilde{\chi}_1^0$ as LSP.
- 126 ABDALLAH 05B use data from $\sqrt{s}=130$ –209 GeV, looking for events with diphotons + \cancel{E} . Limits on the cross-section are computed in the plane (m($\widetilde{\chi}_2^0$), m($\widetilde{\chi}_1^0$)), see Fig. 12. Supersedes the results of ABREU 00Z.
- 127 ACHARD 04E use data from $\sqrt{s}=189$ –209 GeV, looking for events with diphotons + $\cancel{\mathbb{E}}$. Limits are computed in the plane (m($\widetilde{\chi}_2^0$), m(\widetilde{e}_R)), for $\Delta m_0>10$ GeV, see Fig. 7. Supersedes the results of ACCIARRI 99R.
- 128 ACHARD 02 searches for the production of sparticles in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit of $\widetilde{\chi}_2^0$ holds for \overline{UDD} couplings and increases to 84.0 GeV for $LL\overline{E}$ couplings. The same $\widetilde{\chi}_3^0$ limit holds for both $LL\overline{E}$ and \overline{UDD} couplings. For L3 limits from $LQ\overline{D}$ couplings, see ACCIARRI 01.
- ABREU 01B used data from \sqrt{s} =189 GeV to search for the production of $\widetilde{\chi}_{i}^{0}\widetilde{\chi}_{j}^{0}$. They looked for di-jet and di-lepton pairs with $\not\!\!E$ for events from $\widetilde{\chi}_{i}^{0}\widetilde{\chi}_{j}^{0}$ with the decay $\widetilde{\chi}_{j}^{0} \to f\overline{f}\widetilde{\chi}_{1}^{0}$; multi-jet and multi-lepton pairs with or without additional photons to cover the cascade decays $\widetilde{\chi}_{j}^{0} \to f\overline{f}\widetilde{\chi}_{2}^{0}$, followed by $\widetilde{\chi}_{j}^{0} \to f\overline{f}\widetilde{\chi}_{1}^{0}$ or $\widetilde{\chi}_{j}^{0} \to \gamma\widetilde{\chi}_{1}^{0}$; multi-tau final states from $\widetilde{\chi}_{2}^{0} \to \widetilde{\tau}\tau$ with $\widetilde{\tau} \to \tau\widetilde{\chi}_{1}^{0}$. See Figs. 9 and 10 for limits on the (μ, M_{2}) plane for $\tan\beta$ =1.0 and different values of m_{0} .
- 130 ACCIARRI 01 searches for multi-lepton and/or multi-jet final states from R prompt decays with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings at $\sqrt{s}{=}189$ GeV. The search is performed for direct and indirect decays of neutralinos, charginos, and scalar leptons, with the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}$ as LSP and assuming one coupling to be nonzero at a time. Mass limits are derived using simultaneously the constraints from the neutralino, chargino, and slepton analyses; and the Z^0 width measurements from ACCIARRI 00C in a scan of the parameter space assuming MSUGRA with gaugino and scalar mass universality. Updates and supersedes the results from ACCIARRI 99I.
- ABREU 000 searches for the production of charginos and neutralinos in the case of R-parity violation with $LL\overline{E}$ couplings, using data from \sqrt{s} =189 GeV. They investigate topologies with multiple leptons or jets plus leptons, assuming one coupling to be nonzero at the time and giving rise to direct or indirect decays. LImits are obtained in the M_2 versus μ plane and a limit on the neutralino mass is derived from a scan over the parameters m_0 and $\tan\beta$.
- ABBIENDI 99F looked for $\gamma E\!\!\!\!/$ final states at $\sqrt{s}{=}183$ GeV. They obtained an upper bound on the cross section for the production $e^+e^- \to \widetilde{\chi}_2^0 \widetilde{\chi}_1^0$ followed by the prompt decay $\widetilde{\chi}_2^0 \to \gamma \widetilde{\chi}_1^0$ of 0.075–0.80 pb in the region $m_{\widetilde{\chi}_2^0} + m_{\widetilde{\chi}_1^0} > m_Z$, $m_{\widetilde{\chi}_2^0} = 91$ –183 GeV, and $\Delta m_0 > 5$ GeV. See Fig. 7 for explicit limits in the $(m_{\widetilde{\chi}_2^0}, m_{\widetilde{\chi}_1^0})$ plane.

- 133 ABBIENDI 99F looked for $\gamma\gamma E$ final states at \sqrt{s} =183 GeV. They obtained an upper bound on the cross section for the production e $^+$ e $^ \rightarrow$ $~\widetilde{\chi}^0_2\,\widetilde{\chi}^0_2$ followed by the prompt decay $\widetilde{\chi}^0_2 \to \gamma \widetilde{\chi}^0_1$ of 0.08–0.37 pb for $m_{\widetilde{\chi}^0_2}$ =45–81.5 GeV, and $\Delta m_0 > 5$ GeV. See Fig. 11 for explicit limits in the $(m_{\widetilde{\chi}^0_2}, m_{\widetilde{\chi}^0_1})$ plane.
- 134 ABBOTT 98C searches for trilepton final states ($\ell = e, \mu$). See footnote to ABBOTT 98C in the Chargino Section for details on the assumptions. Assuming a negligible decay rate of $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_2^0$ to quarks, they obtain $m_{\widetilde{\chi}_2^0} \gtrsim$ 103 GeV.
- 135 ABE 98J searches for trilepton final states ($\ell = e, \mu$). See footnote to ABE 98J in the Chargino Section for details on the assumptions. The quoted result for $m_{\widetilde{\chi}^0_2}$ corresponds to the best limit within the selected range of parameters, obtained for $m_{\widetilde{q}} > m_{\widetilde{g}}$, $\tan \beta = 2$, and μ =-600 GeV.
- 136 ACCIARRI 98F is obtained from direct searches in the $e^+e^-
 ightarrow ~ \widetilde{\chi}^0_{1.2} \widetilde{\chi}^0_2$ production channels, and indirectly from $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_1^0$ searches within the MSSM. See footnote to ACCIARRI 98F in the chargino Section for further details on the assumptions. Data taken at $\sqrt{s} = 130-172$ GeV.
- 137 ACCIARRI 98V looked for $\gamma(\gamma) \not\! E$ final states at $\sqrt{s} = 183$ GeV. They obtained an upper bound on the cross section for the production $e^+e^- \to \widetilde{\chi}^0_2 \widetilde{\chi}^0_{1,2}$ followed by the prompt decay $\widetilde{\chi}_2^0 \to \gamma \widetilde{\chi}_1^0$. See Figs. 4a and 6a for explicit limits in the $(m_{\widetilde{\chi}_2^0}, m_{\widetilde{\chi}_1^0})$ plane.
- 138 BARATE 98H looked for $\gamma\gamma\not\sqsubseteq$ final states at $\sqrt{s}=161{,}172$ GeV. They obtained an upper bound on the cross section for the production $e^+e^- \to \tilde{\chi}_2^0 \tilde{\chi}_2^0$ followed by the prompt decay $\tilde{\chi}_2^0 \to \gamma \tilde{\chi}_1^0$ of 0.4–0.8 pb for $m_{\tilde{\chi}_2^0} = 10$ –80 GeV. The bound above is for the specific case of $\widetilde{\chi}_1^0=\widetilde{H}^0$ and $\widetilde{\chi}_2^0=\widetilde{\gamma}$ and $m_{\widetilde{e}_R}=100$ GeV. See Fig. 6 and 7 for explicit limits in the $(\widetilde{\chi}_2^0,\widetilde{\chi}_1^0)$ plane and in the $(\widetilde{\chi}_2^0,\widetilde{e}_R)$ plane.
- 139 BARATE 98J looked for $\gamma\gamma\not$ $ot\!\!E$ final states at $\sqrt{s}=161$ –183 GeV. They obtained an upper bound on the cross section for the production e^+ e^- \to $~\widetilde{\chi}^0_2~\widetilde{\chi}^0_2$ followed by the prompt decay $\widetilde{\chi}^0_2 \to \gamma \widetilde{\chi}^0_1$ of 0.08–0.24 pb for $m_{\widetilde{\chi}^0_2} <$ 91 GeV. The bound above is for the specific case of $\widetilde{\chi}_1^0 = \widetilde{H}^0$ and $\widetilde{\chi}_2^0 = \widetilde{\gamma}$ and $m_{\widetilde{e}_D} = 100$ GeV.
- $^{140}\,\mathsf{ABACHI}$ 96 searches for 3-lepton final states. Efficiencies are calculated using mass relations and branching ratios in the Minimal Supergravity scenario. Results are presented as lower bounds on $\sigma(\widetilde{\chi}_1^{\pm}\widetilde{\chi}_2^0) \times \mathsf{B}(\widetilde{\chi}_1^{\pm} \to \ell \nu_\ell \widetilde{\chi}_1^0) \times \mathsf{B}(\widetilde{\chi}_2^0 \to \ell^+ \ell^- \widetilde{\chi}_1^0)$ as a function of $m_{\widetilde{\chi}_1^0}$. Limits range from 3.1 pb $(m_{\widetilde{\chi}_1^0} = 45 \text{ GeV})$ to 0.6 pb $(m_{\widetilde{\chi}_1^0} = 100 \text{ GeV})$.
- 141 ABE 96K looked for trilepton events from chargino-neutralino production. They obtained lower bounds on $m_{\widetilde{\chi}0}$ as a function of μ . The lower bounds are in the 45–50 GeV range for gaugino-dominant $\tilde{\chi}_2^0$ with negative μ , if $\tan\!\beta <\!10$. See paper for more details of the assumptions.

 $\widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_2^{\pm}$ (Charginos) MASS LIMITS Charginos are unknown mixtures of w-inos and charged higgsinos (the supersymmetric partners of W and Higgs bosons). A lower mass limit for the lightest chargino $(\tilde{\chi}_1^{\pm})$ of approximately 45 GeV, independent of the field composition and of the decay mode, has been obtained by the LEP experiments from the analysis of the Z width and decays. These results, as well as other now superseded limits from e^+e^- collisions

at energies below 136 GeV, and from hadronic collisions, can be found in the 1998 Edition (The European Physical Journal **C3** 1 (1998)) of this Review.

Unless otherwise stated, results in this section assume spectra, production rates, decay modes and branching ratios as evaluated in the MSSM, with gaugino and sfermion mass unification at the GUT scale. These papers generally study production of $\widetilde{\chi}_1^0 \widetilde{\chi}_2^0$, $\widetilde{\chi}_1^+ \widetilde{\chi}_1^-$ and (in the case of hadronic collisions) $\widetilde{\chi}_1^+ \widetilde{\chi}_2^0$ pairs, including the effects of cascade decays. The mass limits on $\widetilde{\chi}_1^\pm$ are either direct, or follow indirectly from the constraints set by the non-observation of $\widetilde{\chi}_2^0$ states on the gaugino and higgsino MSSM parameters M_2 and μ . For generic values of the MSSM parameters, limits from high-energy e^+e^- collisions coincide with the highest value of the mass allowed by phase-space, namely $m_{\widetilde{\chi}_1^\pm} \lesssim \sqrt{s}/2$. At the time of this writing, preliminary and unpublished results from the 2000 run of LEP2 at \sqrt{s} up to \simeq 209 GeV give therefore a lower mass limit of approximately 104 GeV valid for general MSSM models. The limits become however weaker in special regions of the MSSM parameter space where the detection efficiencies or production cross sections are suppressed. For example, this may happen when: (i) the mass differences $\Delta m_+ = m_{\widetilde{\chi}_1^\pm} - m_{\widetilde{\chi}_1^0}$ or $\Delta m_\nu = m_{\widetilde{\chi}_1^\pm} - m_{\widetilde{\chi}_1^0}$ are very small, and the detection efficiency is reduced; (ii) the electron sneutrino mass is small, and the $\widetilde{\chi}_1^\pm$ production rate is suppressed due to a destructive

sneutrino mass is small, and the $\tilde{\chi}_1^\pm$ production rate is suppressed due to a destructive interference between s and t channel exchange diagrams. The regions of MSSM parameter space where the following limits are valid are indicated in the comment lines or in the footnotes.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>101	95	¹⁴² ABBIENDI	04н	OPAL	all tan β , $\Delta m_0 > 5$ GeV,
> 89	95	¹⁴³ ABBIENDI	03н	OPAL	$m_0 >$ 500 GeV, $A_0 = 0$ 0.5 $\leq \Delta m_+ \leq$ 5 GeV,higgsinolike, $\tan\!\beta =$ 1.5
> 97.1	95	¹⁴⁴ ABDALLAH	03м	DLPH	$\widetilde{\chi}_{1}^{\pm}$, $\Delta m_{+} \geq$ 3 GeV, $m_{\widetilde{ u}} > m_{\widetilde{\chi}^{\pm}}$
> 75	95	¹⁴⁴ ABDALLAH	03м	DLPH	$\tilde{\chi}_{1}^{\pm}$, higgsino, all Δm_{+} , $m_{\tilde{f}} > m_{\tilde{\chi}^{\pm}}$
> 70	95	¹⁴⁴ ABDALLAH	03м	DLPH	$\widetilde{\chi}_1^{\pm}$, all Δm_+ , $m_{\widetilde{ u}} >$ 500 GeV,
> 94	95	¹⁴⁵ ABDALLAH			$\widetilde{\chi}_1^{\pm}$, $ aneta \leq 2M_1 \leq 10M_2$ $\widetilde{\chi}_1^{\pm}$, $ aneta \leq 40$, $\Delta m_+ > 3$ GeV,all m_0
> 88	95	¹⁴⁶ HEISTER	02J	ALEP	$\tilde{\chi}_1^{\pm}$, all Δm_+ , large m_0
> 67.7	95	¹⁴⁷ ACCIARRI	00 D	L3	$ aneta > 0.7$, all Δm_+ , all m_0
> 69.4	95	¹⁴⁸ ACCIARRI	00K	L3	$e^+e^- ightarrow~\widetilde{\chi}^\pm\widetilde{\chi}^\mp$, all Δm_+ ,
					heavy scalars

• • • We do not use the following data for averages, fits, limits, etc. • • •

		¹⁴⁹ ABAZOV	06 D	D0	Ŗ, LL E
>195	95	¹⁵⁰ ABAZOV	05A	D0	$p\overline{p} \rightarrow \widetilde{\chi}\widetilde{\chi}, \widetilde{\chi} = \widetilde{\chi}_2^0, \widetilde{\chi}_1^{\pm}, \widetilde{\chi}_1^0 \rightarrow$
					$\gamma \widetilde{\textbf{\textit{G}}}$, GMSB
>117	95	¹⁵¹ ABAZOV	05 U	D0	$p\overline{p} \rightarrow \widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0$
>167	95	¹⁵² ACOSTA	05E	CDF	$p\overline{p} \rightarrow \widetilde{\chi}\widetilde{\chi}, \widetilde{\chi} = \widetilde{\chi}_{2}^{0}, \widetilde{\chi}_{1}^{\pm}, \widetilde{\chi}_{1}^{0} \rightarrow$
					$\gamma \widetilde{G}$, GMSB
> 66	95 153	^{3,154} ABDALLAH	04H	DLPH	AMSB, $\mu > 0$
>102.5	₉₅ 155	^{5,156} ABDALLAH	04M	DLPH	$R(\overline{U}\overline{D}\overline{D})$

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>100		¹⁵⁷ ABDALLAH	03 D	DLPH	$e^+e^- ightarrow ~ \widetilde{\chi}_1^{\pm} \widetilde{\chi}_1^{\mp} (\widetilde{\chi}_1^{\pm} ightarrow ~ \widetilde{\tau}_1 \nu_{ au},$
>103	95	¹⁵⁸ HEISTER	036	ALEP	$\widetilde{ au}_1 ightarrow au \widetilde{G})$ R decays, $m_0 > 500$ GeV
>103	95	¹⁵⁹ ACHARD	02		R, MSUGRA
7 102.1	30	¹⁶⁰ GHODBANE	02	THEO	40, 1113 3 31 0 1
> 94.3	95	¹⁶¹ ABREU			$\tilde{\chi}^{\pm} \rightarrow \tau J$
> 93.8	95	¹⁶² ACCIARRI			$ \mathbb{R}, \text{ all } m_0, 0.7 \leq \tan\beta \leq 40 $
>100	95	¹⁶³ BARATE			R decays, $m_0 > 500 \text{ GeV}$
> 91.8	95	¹⁶⁴ ABREU	00V	DLPH	$e^+e^- ightarrow \widetilde{\chi}_1^{\pm} \widetilde{\chi}_1^{\pm} (\widetilde{\chi}_1^{\pm} ightarrow \widetilde{\tau}_1 \nu_{\tau},$
					$\widetilde{ au}_1 ightarrow \ au \widetilde{ ilde{G}})$
		¹⁶⁵ CHO	00 B	THEO	EW analysis
> 76	95	¹⁶⁶ ABBIENDI	99T	OPAL	<i>Ŗ</i> , <i>m</i> ₀ =500 GeV
> 51	95	¹⁶⁷ MALTONI	99 B	THEO	EW analysis, $\Delta m_+ \sim 1$ GeV
> 81.5	95	¹⁶⁸ ABE	98J	CDF	$p\overline{p} \rightarrow \widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0$
		¹⁶⁹ ACKERSTAFF			
> 65.7	95	¹⁷⁰ ACKERSTAFF	98L	OPAL	$\Delta m_+ >$ 3 GeV, $\Delta m_ u >$ 2 GeV
		171 ACKERSTAFF	98V	OPAL	light gluino
		¹⁷² CARENA	97	THEO	$g_{\mu}-2$
		173 KALINOWSKI	97	THEO	$\overset{'}{W} \rightarrow \ \widetilde{\chi}_{1}^{\pm} \widetilde{\chi}_{1}^{0}$
		¹⁷⁴ ABE			
					± -

- 142 ABBIENDI 04H search for charginos and neutralinos in events with acoplanar leptons+jets and multi-jet final states in the 192–209 GeV data, combined with the results on leptonic final states from ABBIENDI 04. The results hold for a scan over the parameter space covering the region 0 < M_2 <5000 GeV, -1000 < μ <1000 GeV and tan β from 1 to 40. This limit supersedes ABBIENDI 00H.
- ABBIENDI 03H used e^+e^- data at $\sqrt{s}=188$ –209 GeV to search for chargino pair production in the case of small Δm_+ They select events with an energetic photon, large $\not\!\!\!E$ and little hadronic or leptonic activity. The bound applies to higgsino-like charginos with zero lifetime and a 100% branching ratio $\widetilde{\chi}_1^\pm \to \widetilde{\chi}_1^0 W^*$. The mass limit for gaugino-like charginos, in case of non-universal gaugino masses, is of 92 GeV for $m_{\widetilde{\nu}}=1000$ GeV and is lowered to 74 GeV for $m_{\widetilde{\nu}}\geq 100$ GeV. Limits in the plane $(m_{\widetilde{\chi}_1^\pm},\Delta m_+)$ are shown in Fig. 7. Exclusion regions are also derived for the AMSB scenario in the $(m_{3/2},\tan\beta)$ plane, see their Fig. 9.
- ABDALLAH 03M searches for the production of charginos using data from $\sqrt{s}=192$ to 208 GeV to investigate topologies with multiple leptons, jets plus leptons, multi-jets, or isolated photons. The first limit holds for $\tan\beta \geq 1$ and is obtained at $\Delta m_+=3$ GeV in the higgsino region. For $\Delta m_+ \geq 10$ (5) GeV and large m_0 , the limit improves to 102.7 (101.7) GeV. For the region of small Δm_+ , all data from $\sqrt{s}=130$ to 208 GeV are used to investigate final states with heavy stable charged particles, decay vertices inside the detector and soft topologies with a photon from initial state radiation. The second limit is obtained in the higgsino region, assuming gaugino mass universality at the GUT scale and $1 < \tan\beta < 50$. For the case of non-universality of gaugino masses, the parameter space is scanned in the domain $1 < \tan\beta < 50$ and, for $\Delta m_+ < 3$ GeV, for values of M_1 , M_2 and μ such that $M_2 \leq 2M_1 \leq 10M_2$ and $|\mu| \geq M_2$. The third limit is obtained in the gaugino region. See Fig. 36 for the dependence of the low Δm_+ limits on Δm_+ . These limits include and update the results of ABREU 00J and ABREU 00T.
- ¹⁴⁵ ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass of charginos is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays), for charginos

- and for sleptons. These limits are valid for values of $M_2 < 1$ TeV, $|\mu| \le 2$ TeV with the $\widetilde{\chi}_1^0$ as LSP. Constraints from the Higgs search in the M_h^{max} scenario assuming $m_t = 174.3$ GeV are included. The quoted limit applies if there is no mixing in the third family or when $m_{\widetilde{\tau}_1} m_{\widetilde{\chi}_1^0} > 6$ GeV. If mixing is included the limit degrades to 90 GeV. See Fig. 43 for the mass limits as a function of $\tan\beta$. These limits update the results of ABREU 00W.
- HEISTER 02J search for chargino production with small Δm_+ in final states with a hard isolated initial state radiation photon and few low-momentum particles, using 189–208 GeV data. This search is sensitive in the intermediate Δm_+ region. Combined with searches for $\not\!E$ topologies and for stable charged particles, the above bound is obtained for m_0 larger than few hundred GeV, $1 < \tan \beta < 300$ and holds for any chargino field contents. For light scalars, the general limit reduces to the one from the Z^0 , but under the assumption of gaugino and sfermion mass unification the above bound is recovered. See Figs. 4–6 for the more general dependence of the limits on Δm_+ . Updates BARATE 98X.
- 147 ACCIARRI 00D data collected at \sqrt{s} =189 GeV. The results hold over the full parameter space defined by 0.7 \leq tan β \leq 60, 0 \leq M_2 \leq 2 TeV, $|\mu|$ \leq 2 TeV m_0 \leq 500 GeV. The results of slepton searches from ACCIARRI 99W are used to help set constraints in the region of small m_0 . See their Figs. 5 for the tan β and M_2 dependence on the limits. See the text for the impact of a large B($\tilde{\chi}^{\pm} \rightarrow \tau \tilde{\nu}_{\tau}$) on the result. The region of small Δm_+ is excluded by the analysis of ACCIARRI 00K. Updates ACCIARRI 98F.
- 148 ACCIARRI 00K searches for the production of charginos with small Δm_+ using data from $\sqrt{s}{=}189$ GeV. They investigate soft final states with a photon from initial state radiation. The results are combined with the limits on prompt decays from ACCIARRI 00D and from heavy stable charged particles from ACCIARRI 99L (see Heavy Charged Lepton Searches). The production and decay branching ratios are evaluated within the MSSM, assuming heavy sfermions. The parameter space is scanned in the domain $1{<}\tan\beta{<}50, 0.3 < M_1/M_2 < 50,$ and $0{<}|\mu| < 2$ TeV. The limit is obtained in the higgsino region and improves to 78.6 GeV for gaugino-like charginos. The limit is unchanged for light scalar quarks. For light $\widetilde{\tau}$ or $\widetilde{\nu}_{\tau}$, the limit is unchanged in the gaugino-like region and is lowered by 0.8 GeV in the higgsino-like case. For light $\widetilde{\mu}$ or $\widetilde{\nu}_{\mu}$, the limit is unchanged in the higgsino-like region and is lowered by 0.9 GeV in the gaugino-like region. No direct mass limits are obtained for light \widetilde{e} or $\widetilde{\nu}_{\rho}$.
- 149 ABAZOV 06D looked in 360 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with three leptons originating from the pair production of charginos and neutralinos, followed by R decays mediated by $LL\overline{E}$ couplings. One coupling is assumed to be dominant at a time. No significant excess was found compared to the background expectation in the $ee\ell$, $\mu\mu\ell$ nor $ee\tau$ ($\ell=e$, μ) final states. Upper limits on the cross-section are extracted in a specific MSUGRA model and a MSSM model without unification of M_1 and M_2 at the GUT scale. A limit is derived on the masses of charginos and neutralinos for both scenarios assuming λ_{ijk} couplings such that the decay length is less than 1 cm, see their Table III and Fig. 4.
- ABAZOV 05A looked in 263 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for diphoton events with large $\not\!\!E_T$. They may originate from the production of $\widetilde{\chi}^\pm$ in pairs or associated to a $\widetilde{\chi}^0_2$, decaying to a $\widetilde{\chi}^0_1$ which itself decays promptly in GMSB to $\widetilde{\chi}^0_1 \to \gamma \, \widetilde{G}$. No significant excess was found at large $\not\!\!E_T$ compared to the background expectation. A limit is derived on the masses of SUSY particles in the GMSB framework for M=2 Λ , N=1, $\tan\beta=15$ and $\mu>0$, see Figure 2. It also excludes $\Lambda<79.6$ TeV. Very similar results are obtained for different choices of parameters, see their Table 2. Supersedes the results of ABBOTT 98.
- 151 ABAZOV 05 U looked in $320~{\rm pb}^{-1}$ of $p\,\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with large E_T , no jets and three leptons (e,μ,τ) of which at least two are e or $\mu.$ No significant excess was found at large E_T compared to the background expectation. A limit is derived on the cross section times branching ratio to 3 leptons, see their Figures 2 and 3. The mass limit assumes gaugino mass universality, three degenerate sleptons and "maximally

- enhanced" leptonic branching fraction, i.e. a decay dominated by a slepton rather than W/Z. If, in addition, squarks are heavy, the limit improves to 132 GeV. Supersedes the results of ABBOTT 98C.
- 152 ACOSTA 05E looked in 202 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.96$ TeV for diphoton events with large $\not\!\!E_T$. They may originate from the production of $\widetilde{\chi}^\pm$ in pairs or associated to a $\widetilde{\chi}^0_2$, decaying to a $\widetilde{\chi}^0_1$ which itself decays promptly in GMSB to $\gamma\,\widetilde{G}$. No events are selected at large $\not\!\!E_T$ compared to the background expectation. A limit is derived on the masses of SUSY particles in the GMSB framework for M=2 $\Lambda,~N=1,~\tan\beta=15$ and $\mu>0$, see Figure 2. It also excludes $\Lambda<69$ TeV. Supersedes the results of ABE 99I.
- ABDALLAH 04H use data from LEP 1 and $\sqrt{s}=192$ –208 GeV. They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region $1 < m_{3/2} < 50$ TeV, $0 < m_0 < 1000$ GeV, $1.5 < \tan \beta < 35$, both signs of μ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z width of 3.2 MeV. The limit is for $m_t=174.3$ GeV (see Table 2 for other m_t values).
- 154 The limit improves to 73 GeV for $\mu~<$ 0.
- ^{155} ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $\overline{U}\overline{D}\overline{D}$ couplings. The results are valid in the ranges $90 < m_0 < 500$ GeV, $0.7 < \tan\beta < 30$, $-200 < \mu < 200$ GeV, $0 < M_2 < 400$ GeV. Supersedes the result of ABREU 01D and ABREU 00U.
- 156 The limit improves to 103 GeV for $LL\overline{E}$ couplings.
- ABDALLAH 03D use data from $\sqrt{s}=183$ –208 GeV. They look for final states with two acoplanar leptons, expected in GMSB when the $\widetilde{\tau}_1$ is the NLSP and assuming a short-lived $\widetilde{\chi}_1^{\pm}$. Limits are obtained in the plane $(\mathsf{m}(\widetilde{\tau}),\mathsf{m}(\widetilde{\chi}_1^{\pm}))$ for different domains of $\mathsf{m}(\widetilde{G})$, after combining these results with the search for slepton pair production from the same paper. The limit above is valid if the $\widetilde{\tau}_1$ is the NLSP for all values of $\mathsf{m}(\widetilde{G})$ provided $\mathsf{m}(\widetilde{\chi}_1^{\pm}) \mathsf{m}(\widetilde{\tau}_1) \geq 0.3$ GeV. For larger $\mathsf{m}(\widetilde{G}) > 100$ eV the limit improves to 102 GeV, see their Fig. 11. In the co-NLSP scenario, the limits are 96 and 102 GeV for all $\mathsf{m}(\widetilde{G})$ and $\mathsf{m}(\widetilde{G}) > 100$ eV, respectively. Supersedes the results of ABREU 01G.
- ¹⁵⁸ HEISTER 03G searches for the production of charginos prompt decays. in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at \sqrt{s} =189–209 GeV. The search is performed for indirect decays, assuming one coupling at a time to be non-zero. The limit holds for tan β =1.41. Excluded regions in the (μ,M_2) plane are shown in their Fig. 3.
- 159 ACHARD 02 searches for the production of sparticles in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit of $\widetilde{\chi}_1^{\pm}$ holds for \overline{UDD} couplings and increases to 103.0 GeV for $LL\overline{E}$ couplings. For L3 limits from $LQ\overline{D}$ couplings, see ACCIARRI 01.
- 160 GHODBANE 02 reanalyzes DELPHI data at \sqrt{s} =189 GeV in the presence of complex phases for the MSSM parameters.
- ¹⁶¹ ABREU 01C looked for τ pairs with $\not \! E$ at \sqrt{s} =183–189 GeV to search for the associated production of charginos, followed by the decay $\widetilde{\chi}^{\pm} \to \tau J$, J being an invisible massless particle. See Fig. 6 for the regions excluded in the (μ, M_2) plane.
- 162 ACCIARRI 01 searches for multi-lepton and/or multi-jet final states from R prompt decays with $LL\overline{E},\,LQ\overline{D},$ or \overline{UDD} couplings at $\sqrt{s}{=}189$ GeV. The search is performed for direct and indirect decays of neutralinos, charginos, and scalar leptons, with the $\widetilde{\chi}^0_1$ or a $\widetilde{\ell}$ as LSP and assuming one coupling to be nonzero at a time. Mass limits are derived using simultaneously the constraints from the neutralino, chargino, and slepton analyses; and the Z^0 width measurements from ACCIARRI 00C in a scan of the parameter space

- assuming MSUGRA with gaugino and scalar mass universality. Updates and supersedes the results from ACCIARRI 991.
- ¹⁶³ BARATE 01B searches for the production of charginos in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings at \sqrt{s} =189–202 GeV. The search is performed for indirect decays, assuming one coupling at a time to be nonzero. Updates BARATE 00H.
- ABREU 00V use data from $\sqrt{s}=183-189$ GeV. They look for final states with two acoplanar leptons, expected in GMSB when the $\widetilde{\tau}_1$ is the NLSP and assuming a short-lived $\widetilde{\chi}_1^{\pm}$. Limits are obtained in the plane $(m_{\widetilde{\tau}}, m_{\widetilde{\chi}_1^{\pm}})$ for different domains of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. The limit above is valid for all values of $m_{\widetilde{C}}$.
- 165 CHO 00B studied constraints on the MSSM spectrum from precision EW observables. Global fits favour charginos with masses at the lower bounds allowed by direct searches. Allowing for variations of the squark and slepton masses does not improve the fits.
- 166 ABBIENDI 99T searches for the production of neutralinos in the case of R-parity violation with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings using data from $\sqrt{s}{=}183$ GeV. They investigate topologies with multiple leptons, jets plus leptons, or multiple jets, assuming one coupling at the time to be non-zero and giving rise to direct or indirect decays. Mixed decays (where one particle has a direct, the other an indirect decay) are also considered for the \overline{UDD} couplings. Upper limits on the cross section are derived which, combined with the constraint from the Z^0 width, allow to exclude regions in the M_2 versus μ plane for any coupling. Limits on the chargino mass are obtained for non-zero $LL\overline{E}$ couplings $> 10^{-5}$ and assuming decays via a W^* .
- 167 MALTONI 99B studied the effect of light chargino-neutralino to the electroweak precision data with a particular focus on the case where they are nearly degenerate ($\Delta m_+ \sim 1$ GeV) which is difficult to exclude from direct collider searches. The quoted limit is for higgsino-like case while the bound improves to 56 GeV for wino-like case. The values of the limits presented here are obtained in an update to MALTONI 99B, as described in MALTONI 00.
- ABE 98J searches for trilepton final states $(\ell=e,\mu)$. Efficiencies are calculated using mass relations in the Minimal Supergravity scenario, exploring the domain of parameter space defined by $1.1 < \tan\beta < 8$, $-1000 < \mu(\text{GeV}) < -200$, and $m_{\widetilde{q}}/m_{\widetilde{g}}=1-2$. In this region $m_{\widetilde{\chi}_1^\pm} \sim m_{\widetilde{\chi}_2^0}$ and $m_{\widetilde{\chi}_1^\pm} \sim 2m_{\widetilde{\chi}_1^0}$. Results are presented in Fig. 1 as upper bounds on $\sigma(p\overline{p}\to\widetilde{\chi}_1^\pm\widetilde{\chi}_2^0)\times \text{B}(3\ell)$. Limits range from 0.8 pb $(m_{\widetilde{\chi}_1^\pm}=50~\text{GeV})$ to 0.23 pb $(m_{\widetilde{\chi}_1^\pm}=100~\text{GeV})$ at 95%CL. The gaugino mass unification hypothesis and the assumed mass relation between squarks and gluinos define the value of the leptonic branching ratios. The quoted result corresponds to the best limit within the selected range of parameters, obtained for $m_{\widetilde{q}}>m_{\widetilde{g}}$, $\tan\beta=2$, and $\mu=-600~\text{GeV}$. Mass limits for different values of $\tan\beta$ and μ are given in Fig. 2.
- ¹⁶⁹ ACKERSTAFF 98K looked for dilepton+ $\not\!\!E_T$ final states at \sqrt{s} =130–172 GeV. Limits on $\sigma(e^+e^- \to \widetilde{\chi}_1^+ \widetilde{\chi}_1^-) \times \mathsf{B}^2(\ell)$, with $\mathsf{B}(\ell) = \mathsf{B}(\chi^+ \to \ell^+ \nu_\ell \chi_1^0)$ ($\mathsf{B}(\ell) = \mathsf{B}(\chi^+ \to \ell^+ \widetilde{\nu}_\ell)$), are given in Fig. 16 (Fig. 17).
- 170 ACKERSTAFF 98L limit is obtained for $0 < M_2 < 1500, \ |\mu| < 500$ and $\tan\beta > 1,$ but remains valid outside this domain. The dependence on the trilinear-coupling parameter A is studied, and found negligible. The limit holds for the smallest value of m_0 consistent with scalar lepton constraints (ACKERSTAFF 97H) and for all values of m_0 where the condition $\Delta m_{\widetilde{\nu}} > 2.0$ GeV is satisfied. $\Delta m_{\nu} > 10$ GeV if $\widetilde{\chi}^{\pm} \rightarrow \ell \widetilde{\nu}_{\ell}$. The limit improves to 84.5 GeV for $m_0 = 1$ TeV. Data taken at $\sqrt{s} = 130 172$ GeV.
- ¹⁷¹ ACKERSTAFF 98V excludes the light gluino with universal gaugino mass where charginos, neutralinos decay as $\widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_2^0 \rightarrow q \overline{q} \widetilde{g}$ from total hadronic cross sections at \sqrt{s} =130–172 GeV. See paper for the case of nonuniversal gaugino mass.

- ¹⁷² CARENA 97 studied the constraints on chargino and sneutrino masses from muon g-2. The bound can be important for large $\tan \beta$.
- ¹⁷³KALINOWSKI 97 studies the constraints on the chargino-neutralino parameter space from limits on $\Gamma(W \to \widetilde{\chi}_1^{\pm} \widetilde{\chi}_1^0)$ achievable at LEP2. This is relevant when $\widetilde{\chi}_1^{\pm}$ is "invisible," i.e., if $\widetilde{\chi}_1^{\pm}$ dominantly decays into $\widetilde{\nu}_{\ell} \ell^{\pm}$ with little energy for the lepton. Small otherwise allowed regions could be excluded.
- ABE 96K looked for trilepton events from chargino-neutralino production. The bound on $m_{\widetilde{\chi}_1^\pm}$ can reach up to 47 GeV for specific choices of parameters. The limits on the combined production cross section times 3-lepton branching ratios range between 1.4 and 0.4 pb, for 45< $m_{\widetilde{\chi}_1^\pm}$ (GeV)<100. See the paper for more details on the parameter dependence of the results.

Long-lived $\tilde{\chi}^{\pm}$ (Chargino) MASS LIMITS

Limits on charginos which leave the detector before decaying.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>102 none 2–93.0		¹⁷⁵ ABBIENDI ¹⁷⁶ ABREU	03L 00T	OPAL DLPH	$m_{\widetilde{\nu}} >$ 500 GeV \widetilde{H}^{\pm} or $m_{\widetilde{\nu}} > m_{\widetilde{\chi}^{\pm}}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

> 83	95	¹⁷⁷ BARATE	97K ALEP
> 28.2	95	ADACHI	90c TOPZ

- ¹⁷⁵ ABBIENDI 03L used e^+e^- data at $\sqrt{s}=130$ –209 GeV to select events with two high momentum tracks with anomalous dE/dx. The excluded cross section is compared to the theoretical expectation as a function of the heavy particle mass in their Fig. 3. The bounds are valid for colorless fermions with lifetime longer than 10^{-6} s. Supersedes the results from ACKERSTAFF 98P.
- 176 ABREU 00T searches for the production of heavy stable charged particles, identified by their ionization or Cherenkov radiation, using data from \sqrt{s} = 130 to 189 GeV. These limits include and update the results of ABREU 98P.
- 177 BARATE 97K uses $e^+\,e^-$ data collected at $\sqrt{s}=130$ –172 GeV. Limit valid for $\tan\beta=\sqrt{2}$ and $m_{\widetilde{\nu}}>100$ GeV. The limit improves to 86 GeV for $m_{\widetilde{\nu}}>250$ GeV.

$\widetilde{ u}$ (Sneutrino) MASS LIMIT

The limits may depend on the number, $N(\widetilde{\nu})$, of sneutrinos assumed to be degenerate in mass. Only $\widetilde{\nu}_L$ (not $\widetilde{\nu}_R$) is assumed to exist. It is possible that $\widetilde{\nu}$ could be the lightest supersymmetric particle (LSP).

We report here, but do not include in the Listings, the limits obtained from the final, but unpublished, fit of the final results obtained by the LEP Collaborations on the invisible width of the Z boson ($\Delta\Gamma_{\rm inv.} <$ 2.0 MeV, LEP 03): $m_{\widetilde{\nu}} >$ 43.7 GeV ($N(\widetilde{\nu})$ =1) and $m_{\widetilde{\nu}} >$ 44.7 GeV ($N(\widetilde{\nu})$ =3) .

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 94	95	¹⁷⁸ ABDALLAH	03м DLPH	
				$m_{\widetilde{e}_R}\!-\!m_{\widetilde{\chi}^0_1}>\!10GeV$
> 84	95		02N ALEP	$\widetilde{ u}_{m{e}}$, any Δm
> 37.1	95		93M L3	$\Gamma(Z o \text{ invisible}); N(\widetilde{\nu})=1$
> 41	95	¹⁸¹ DECAMP	92 ALEP	$\Gamma(Z \rightarrow \text{invisible}); N(\widetilde{\nu})=3$
> 36	95		91F DLPH	$\Gamma(Z o \text{ invisible}); N(\widetilde{\nu})=1$
> 31.2	95	¹⁸² ALEXANDER	91F OPAL	$\Gamma(Z o \text{ invisible}); \ N(\widetilde{\nu}) = 1$

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• • • We do not use the following data for averages, fits, limits, etc. • • •

		¹⁸³ ABAZOV ¹⁸⁴ ABDALLAH	06ı 06c	D0 DLPH	$ \mathcal{R}, \ \lambda'_{211} $ $ \widetilde{\nu}_{\ell}, \ \mathcal{R}, \ (s+t)\text{-channel} $
		¹⁸⁵ ABULENCIA	06м	CDF	$\widetilde{\nu}_{ au}$, R
		186 ABULENCIA	05A	CDF	$p\overline{\overline{p}} ightarrow \ \widetilde{ u} ightarrow \ ee, \mu\mu, ot\!\!/ LQ\overline{\overline{D}}$
		187 ACOSTA	05 R	CDF	$p\overline{p} ightarrow \ \widetilde{ u} ightarrow \ au au$, $ ot\!\!R$, $ ot\!\!LQ\overline{D}$
		188 ABBIENDI	04F	OPAL	$R, \ \widetilde{ u}_{\mathbf{e},\mu, au}$
> 95	95 189	,190 ABDALLAH	04H	DLPH	$AMSB,\ \mu > 0$
> 98	95	¹⁹¹ ABDALLAH	04M	DLPH	$R(LL\overline{E}), \widetilde{\nu}_e, \text{indirect}, \Delta m_0 > 5 \text{ GeV}$
> 85	95	¹⁹¹ ABDALLAH	04M	DLPH	$R(LL\overline{E}), \widetilde{\nu}_{\mu}, \text{indirect}, \Delta m_0 > 5 \text{ GeV}$
> 85	95	¹⁹¹ ABDALLAH	04M	DLPH	$R(LL\overline{E}), \widetilde{\nu}_{\tau}, \text{indirect}, \Delta m_0 > 5 \text{ GeV}$
		¹⁹² ABDALLAH	03F	DLPH	$\widetilde{ u}_{\mu, au}$, R $\mathit{LL}\overline{E}$ decays
		¹⁹³ ACOSTA	03E	CDF	$\widetilde{\nu}$, R , $LQ\overline{D}$ production and $LL\overline{E}$ decays
> 88	95	¹⁹⁴ HEISTER	03 G	ALEP	$\widetilde{\nu}_e$, R decays, $\mu=-200$ GeV,
		104			tan $eta=2$
> 65	95	¹⁹⁴ HEISTER	03 G	ALEP	$\widetilde{ u}_{\mu, au'}$ R decays
		¹⁹⁵ ABAZOV	02H	D0	R, λ'_{211}
> 95	95	¹⁹⁶ ACHARD	02	L3	$\widetilde{\nu}_{e}$, R decays, μ = -200 GeV,
		106			$\tan\!eta\!=\!\!\sqrt{2}$
> 65	95	¹⁹⁶ ACHARD	02	L3	$\widetilde{ u}_{ u, au}$, $ ot\!\!R$ decays
>149	95	¹⁹⁶ ACHARD	02	L3	$\widetilde{ u}$, $ ot\!\!R$ decays, MSUGRA
		¹⁹⁷ HEISTER	02F	ALEP	e $\gamma ightarrow \; \widetilde{ u}_{\mu, au} \ell_{m{k}}$, $ ot\!{R} \; {\it LL} \overline{m{E}}$
none 100-264	95	¹⁹⁸ ABBIENDI	00 R	OPAL	$\widetilde{ u}_{\mu, au}$, R , $(s+t)$ -channel
none 100-200	95	199 ABBIENDI	00 R	OPAL	$\widetilde{ u}_{\mathcal{T}}$, $ ot\!\!R$, s-channel
		200 ABREU	00 S	DLPH	$\widetilde{ u}_\ell$, R , $(s+t)$ -channel
none 50-210	95	²⁰¹ ACCIARRI	00 P	L3	$\widetilde{ u}_{\mu, au}$, R , s-channel
none 50-210	95	²⁰² BARATE	001	ALEP	$\widetilde{ u}_{\mu, au}$, R , $(s+t)$ -channel
none 90-210	95	²⁰³ BARATE	001	ALEP	$\widetilde{\nu}_{\tau}$, R , s-channel
none 100-160	95	²⁰⁴ ABBIENDI	99	OPAL	$\widetilde{\nu}_{m{e}}$, R , t -channel
\neq m $_Z$	95	²⁰⁵ ACCIARRI	97 U	L3	$\widetilde{ u}_{\mathcal{T}}$, R , s-channel
none 125-180	95	²⁰⁵ ACCIARRI	97 U	L3	$\widetilde{ u}_{\mathcal{T}}$, $ ot\!\!R$, s-channel
		²⁰⁶ CARENA	97	THEO	$g_{\mu}-2$
> 46.0	95	²⁰⁷ BUSKULIC	95E	ALEP	$N(\widetilde{\nu})=1, \ \widetilde{\nu} \rightarrow \ \nu \nu \ell \overline{\ell}'$
none 20-25000)	²⁰⁸ BECK	94	COSM	Stable $\widetilde{ u}$, dark matter
<600		²⁰⁹ FALK	94	COSM	$\widetilde{ u}$ LSP, cosmic abundance
none 3–90	90	²¹⁰ SATO	91	KAMI	Stable $\widetilde{ u}_{\mathbf{e}}$ or $\widetilde{ u}_{\mu}$,
none 4 00	90	²¹⁰ SATO	01	IZ A 1 A I	dark matter
none 4–90	90	- 3ATU	91	KAMI	Stable $\widetilde{\nu}_{\tau}$, dark matter

 $^{^{178}}$ ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of $M_2<1$ TeV, $|\mu|\leq 1$ TeV with the $\tilde{\chi}_1^0$ as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of $\tan\beta$. These limits update the results of ABREU 00W.

 $^{^{179}}$ HEISTER 02N derives a bound on $m_{\widetilde{\nu}_e}$ by exploiting the mass relation between the $\widetilde{\nu}_e$ and \widetilde{e} , based on the assumption of universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 and the search described in the \widetilde{e} section. In the MSUGRA framework with

- radiative electroweak symmetry breaking, the limit improves to $m_{\widetilde{\nu}_e} > 130$ GeV, assuming a trilinear coupling $A_0 = 0$ at the GUT scale. See Figs. 5 and 7 for the dependence of the limits on $\tan \beta$.
- $^{180}\,\text{ADRIANI 93M}$ limit from $\Delta\Gamma(Z)\text{(invisible)}{<}$ 16.2 MeV.
- 181 DECAMP 92 limit is from $\Gamma(\text{invisible})/\Gamma(\ell\ell)=5.91\pm0.15$ ($N_{\nu}=2.97\pm0.07$).
- ¹⁸² ALEXANDER 91F limit is for one species of $\tilde{\nu}$ and is derived from $\Gamma(\text{invisible, new})/\Gamma(\ell\ell)$ < 0.38.
- ¹⁸³ ABAZOV 06I looked in 380 pb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with at least 2 muons and 2 jets for s-channel production of $\widetilde{\mu}$ or $\widetilde{\nu}$ and subsequent decay via R couplings $LQ\overline{D}$. The data are in agreement with the SM expectation. They set limits on resonant slepton production and derive exclusion contours on λ'_{211} in the mass plane of $\widetilde{\ell}$ versus $\widetilde{\chi}_1^0$ assuming a MSUGRA model with $\tan\beta=5$, $\mu<0$ and $A_0=0$, see their Fig. 3. For $\lambda'_{211}\geq0.09$ slepton masses up to 358 GeV are excluded. Supersedes the results of ABAZOV 02H.
- ABDALLAH 06C searches for anomalies in the production cross sections and forward-backward asymmetries of the $\ell^+\ell^-(\gamma)$ final states ($\ell=e,\mu,\tau$) from 675 pb $^{-1}$ of e^+e^- data at \sqrt{s} =130–207 GeV. Limits are set on the s- and t-channel exchange of sneutrinos in the presence of R with $\lambda LL\overline{E}$ couplings. For points between the energies at which data were taken, information is obtained from events in which a photon was radiated. Exclusion limits in the $(\lambda,m_{\widetilde{\nu}})$ plane are given in Fig. 16. These limits include and update the results of ABREU 00S.
- ABULENCIA 06M searched in 344 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for an excess of events with oppositely charged $e\mu$ pairs. They might be expected in a SUSY model with R where a sneutrino is produced by $LQ\overline{D}$ couplings and decays via $LL\overline{E}$ couplings, focusing on $\widetilde{\nu}_{\tau}$, hence on the λ'_{311} and λ_{132} constants. No significant excess was found compared to the background expectation. Upper limits on the cross-section times branching ratio are extracted and exclusion regions determined for the $\widetilde{\nu}_{\tau}$ mass as a function of both couplings, see their Fig. 3. As an indication, $\widetilde{\nu}_{\tau}$ masses are excluded up to 300 GeV for $\lambda'_{311} \geq 0.01$ and $\lambda_{132} \geq 0.02$.
- 186 ABULENCIA 05A looked in \sim 200 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for dimuon and dielectron events. They may originate from the R production of a sneutrino decaying to dileptons. No significant excess rate was found compared to the background expectation. A limit is derived on the cross section times branching ratio, B, of $\widetilde{\nu} \rightarrow ee$, $\mu\mu$ of 25 fb at high mass, see their Figure 2. Sneutrino masses are excluded at 95% CL below 680, 620, 460 GeV (ee channel) and 665, 590, 450 GeV ($\mu\mu$ channel) for a λ' coupling and branching ratio such that $\lambda'^2 B = 0.01, 0.005, 0.001$, respectively.
- 187 ACOSTA 05R looked in 195 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for ditau events with one identified hadronic tau decay and one other tau decay. They may originate from the R production of a sneutrino decaying to $\tau\tau$. No significant excess rate was found compared to the background expectation, dominated by Drell-Yan. A limit is derived on the cross section times branching ratio, B, of $\widetilde{\nu} \to \tau\tau$, see their Figure 3. Sneutrino masses below 377 GeV are excluded at 95% CL for a λ' coupling to $d\overline{d}$ and branching ratio such that $\lambda'^2 B=0.01$.
- ABBIENDI 04F use data from $\sqrt{s}=189$ –209 GeV. They derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $LQ\overline{D}$ couplings. The results are valid for $\tan\beta=1.5$, $\mu=-200$ GeV, and a BR for the decay given by CMSSM, assuming no sensitivity to other decays. Limits are quoted for $m_{\widetilde{\chi}0}=60$ GeV and degrade for low-mass $\widetilde{\chi}_1^0$. For $\widetilde{\nu}_e$ the direct (indirect) limits with $LL\overline{E}$ couplings are 89 (95) GeV and with $LQ\overline{D}$ they are 89 (88) GeV. For $\widetilde{\nu}_{\mu,\tau}$ the direct (indirect) limits with $LL\overline{E}$ couplings are 79 (81) GeV and with $LQ\overline{D}$ they are 74 (no limit) GeV. Supersedes the results of ABBIENDI 00.
- ¹⁸⁹ ABDALLAH 04H use data from LEP 1 and $\sqrt{s}=192$ –208 GeV. They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region

- $1 < m_{3/2} <$ 50 TeV, $0 < m_0 <$ 1000 GeV, 1.5 <tan $\beta <$ 35, both signs of μ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z width of 3.2 MeV. The limit is for $m_t = 174.3$ GeV (see Table 2 for other m_t values).
- $^{190}\,\mathrm{The}$ limit improves to 114 GeV for $\mu~<0.$
- 191 ABDALLAH 04M use data from $\sqrt{s}=189$ –208 GeV. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5,\,\Delta m_0>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect decays using the neutralino constraint of 39.5 GeV, also derived in ABDALLAH 04M. For indirect decays the limit on $\widetilde{\nu}_e$ decreases to 96 GeV if the constraint from the neutralino is not used and for direct decays it remains 96 GeV. For indirect decays the limit on $\widetilde{\nu}_\mu$ decreases to 82 GeV if the constraint from the neutralino is not used and to 83 GeV for direct decays. For indirect decays the limit on $\widetilde{\nu}_\tau$ decreases to 82 GeV if the constraint from the neutralino is not used and improves to 91 GeV for direct decays. Supersedes the results of ABREU 00U.
- ¹⁹² ABDALLAH 03F looked for events of the type $e^+e^- \to \widetilde{\nu} \to \widetilde{\chi}^0 \nu$, $\widetilde{\chi}^\pm \ell^\mp$ followed by $\not R$ decays of the $\widetilde{\chi}^0$ via λ_{1j1} (j = 2,3) couplings in the data at $\sqrt{s}=183$ –208 GeV. From a scan over the SUGRA parameters, they derive upper limits on the λ_{1j1} couplings as a function of the sneutrino mass, see their Figs. 5-8.
- ¹⁹³ ACOSTA 03E search for $e\mu$, $e\tau$ and $\mu\tau$ final states, and sets limits on the product of production cross-section and decay branching ratio for a $\tilde{\nu}$ in RPV models (see Fig. 3).
- HEISTER 03G searches for the production of sneutrinos in the case of \not{R} prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The search is performed for direct and indirect decays, assuming one coupling at a time to be non-zero. The limit holds for indirect $\overline{\nu}$ decays via \overline{UDD} couplings and $\Delta m>10$ GeV. Stronger limits are reached for $(\overline{\nu}_e,\overline{\nu}_{\mu,\tau})$ for $LL\overline{E}$ direct (100,90) GeV or indirect (98,89) GeV and for $LQ\overline{D}$ direct (–,79) GeV or indirect (91,78) GeV couplings. For $LL\overline{E}$ indirect decays, use is made of the bound $m(\widetilde{\chi}_1^0)>23$ GeV from BARATE 98S. Supersedes the results from BARATE 01B.
- ABAZOV 02H looked in 94 pb $^{-1}$ of $p\overline{p}$ collisions at \sqrt{s} =1.8 TeV for events with at least 2 muons and 2 jets for s-channel production of $\widetilde{\mu}$ or $\widetilde{\nu}$ and subsequent decay via R couplings $LQ\overline{D}$. A scan over the MSUGRA parameters is performed to exclude regions of the $(m_0, m_{1/2})$ plane, examples being shown in Fig. 2.
- 196 ACHARD 02 searches for the associated production of sneutrinos in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit holds for direct decays via $LL\overline{E}$ couplings. Stronger limits are reached for $(\widetilde{\nu}_e,\widetilde{\nu}_{\mu,\tau})$ for $LL\overline{E}$ indirect (99,78) GeV and for \overline{UDD} direct or indirect (99,70) GeV decays. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit holds for \overline{UDD} couplings and increases to 152.7 GeV for $LL\overline{E}$ couplings.
- HEISTER 02F searched for single sneutrino production via $e\gamma \to \tilde{\nu}_j \ell_k$ mediated by $\not\!\!R$ $LL\overline{E}$ couplings, decaying directly or indirectly via a $\tilde{\chi}_1^0$ and assuming a single coupling to be nonzero at a time. Final states with three leptons and possible $\not\!\!E_T$ due to neutrinos were selected in the 189–209 GeV data. Limits on the couplings λ_{1jk} as function of the sneutrino mass are shown in Figs. 10–14. The couplings λ_{232} and λ_{233} are not accessible and λ_{121} and λ_{131} are measured with better accuracy in sneutrino resonant production. For all tested couplings, except λ_{133} , the limits are significantly improved compared to the low-energy limits.
- ABBIENDI 00R studied the effect of s- and t-channel τ or μ sneutrino exchange in $e^+e^- \rightarrow e^+e^-$ at $\sqrt{s}{=}130{-}189$ GeV, via the R-parity violating coupling $\lambda_{1i1}L_1L_ie_1$ ($i{=}2$ or 3). The limits quoted here hold for $\lambda_{1i1}>0.13$, and supersede the results of ABBIENDI 99. See Fig. 11 for limits on $m_{\widetilde{\nu}}$ versus coupling.

- ABBIENDI 00R studied the effect of s-channel τ sneutrino exchange in $e^+e^- \rightarrow \mu^+\mu^-$ at \sqrt{s} =130–189 GeV, in presence of the R-parity violating couplings $\lambda_{i3i}L_iL_3e_i$ (i=1 and 2), with $\lambda_{131}=\lambda_{232}$. The limits quoted here hold for $\lambda_{131}>$ 0.09, and supersede the results of ABBIENDI 99. See Fig. 12 for limits on $m_{\widetilde{\nu}}$ versus coupling.
- ABREU 00s searches for anomalies in the production cross sections and forward-backward asymmetries of the $\ell^+\ell^-(\gamma)$ final states ($\ell=e,\mu,\tau$) from e^+e^- collisions at \sqrt{s} =130–189 GeV. Limits are set on the s- and t-channel exchange of sneutrinos in the presence of R with $\lambda LL\overline{E}$ couplings. For points between the energies at which data were taken, information is obtained from events in which a photon was radiated. Exclusion limits in the $(\lambda,m_{\widetilde{\nu}})$ plane are given in Fig. 5. These limits include and update the results of ABREU 99A.
- 201 ACCIARRI 00P use the dilepton total cross sections and asymmetries at $\sqrt{s}=m_Z$ and $\sqrt{s}=130-189$ GeV data to set limits on the effect of R $LL\overline{E}$ couplings giving rise to μ or τ sneutrino exchange. See their Fig. 5 for limits on the sneutrino mass versus couplings.
- ²⁰²BARATE 00I studied the effect of s-channel and t-channel τ or μ sneutrino exchange in $e^+e^- \rightarrow e^+e^-$ at $\sqrt{s}=$ 130–183 GeV, via the R-parity violating coupling $\lambda_{1i1}L_1L_ie_1^c$ (i=2 or 3). The limits quoted here hold for $\lambda_{1i1}>$ 0.1. See their Fig. 15 for limits as a function of the coupling.
- ²⁰³ BARATE 00I studied the effect of s-channel τ sneutrino exchange in $e^+e^- \rightarrow \mu^+\mu^-$ at $\sqrt{s} = 130$ –183 GeV, in presence of the *R*-parity violating coupling $\lambda_{i3i}L_iL_3e_i^c$ (i=1 and 2). The limits quoted here hold for $\sqrt{|\lambda_{131}\lambda_{232}|} > 0.2$. See their Fig. 16 for limits as a function of the coupling.
- ²⁰⁴ ABBIENDI 99 studied the effect of *t*-channel electron sneutrino exchange in $e^+e^- \rightarrow \tau^+\tau^-$ at \sqrt{s} =130–183 GeV, in presence of the *R*-parity violating couplings $\lambda_{131}L_1L_3e_1^c$. The limits quoted here hold for $\lambda_{131}>0.6$.
- ²⁰⁵ ACCIARRI 97U studied the effect of the s-channel tau-sneutrino exchange in $e^+e^- \rightarrow e^+e^-$ at $\sqrt{s}=m_Z$ and $\sqrt{s}=130-172$ GeV, via the R-parity violating coupling $\lambda_{131}L_1L_ie_1^c$. The limits quoted here hold for $\lambda_{131}>0.05$. Similar limits were studied in $e^+e^- \rightarrow \mu^+\mu^-$ together with $\lambda_{232}L_2L_3e_2^c$ coupling.
- ²⁰⁶ CARENA 97 studied the constraints on chargino and sneutrino masses from muon g-2. The bound can be important for large $tan\beta$.
- ²⁰⁷ BUSKULIC 95E looked for $Z \to \widetilde{\nu} \overline{\widetilde{\nu}}$, where $\widetilde{\nu} \to \nu \chi_1^0$ and χ_1^0 decays via *R*-parity violating interactions into two leptons and a neutrino.
- ²⁰⁸ BECK 94 limit can be inferred from limit on Dirac neutrino using $\sigma(\tilde{\nu}) = 4\sigma(\nu)$. Also private communication with H.V. Klapdor-Kleingrothaus.
- FALK 94 puts an upper bound on $m_{\widetilde{\nu}}$ when $\widetilde{\nu}$ is LSP by requiring its relic density does not overclose the Universe.
- 210 SATO 91 search for high-energy neutrinos from the sun produced by annihilation of sneutrinos in the sun. Sneutrinos are assumed to be stable and to constitute dark matter in our galaxy. SATO 91 follow the analysis of NG 87, OLIVE 88, and GAISSER 86.

CHARGED SLEPTONS

This section contains limits on charged scalar leptons $(\widetilde{\ell}, \text{ with } \ell = e, \mu, \tau)$. Studies of width and decays of the Z boson (use is made here of $\Delta\Gamma_{\mbox{inv}} < 2.0 \, \mbox{MeV}, \mbox{ LEP 00})$ conclusively rule out $m_{\widetilde{\ell}_R} < 40 \, \mbox{GeV}$ (41

GeV for ℓ_L) , independently of decay modes, for each individual slepton. The limits improve to 43 GeV (43.5 GeV for $\widetilde{\ell}_L$) assuming all 3 flavors to be degenerate. Limits on higher mass sleptons depend on model assumptions and on the mass splitting $\Delta m = m_{\widetilde{\ell}} - m_{\widetilde{\chi}_1^0}$. The mass and composition

of $\widetilde{\chi}^0_1$ may affect the selectron production rate in e^+e^- collisions through

t-channel exchange diagrams. Production rates are also affected by the potentially large mixing angle of the lightest mass eigenstate $\widetilde{\ell}_1 = \widetilde{\ell}_R \sin\theta_\ell + \widetilde{\ell}_L \cos\theta_\ell$. It is generally assumed that only $\widetilde{\tau}$ may have significant mixing. The coupling to the Z vanishes for $\theta_\ell = 0.82$. In the high-energy limit of $e^+ \, e^-$ collisions the interference between γ and Z exchange leads to a minimal cross section for $\theta_\ell = 0.91$, a value which is sometimes used in the following entries relative to data taken at LEP2. When limits on $m_{\widetilde{\ell}_R}$ are quoted, it is understood that limits on $m_{\widetilde{\ell}_\ell}$ are usually at least as strong.

Possibly open decays involving gauginos other than $\widetilde{\chi}^0_1$ will affect the detection efficiencies. Unless otherwise stated, the limits presented here result from the study of $\widetilde{\ell}^+\widetilde{\ell}^-$ production, with production rates and decay properties derived from the MSSM. Limits made obsolete by the recent analyses of e^+e^- collisions at high energies can be found in previous Editions of this Review.

For decays with final state gravitinos (\widetilde{G}) , $m_{\widetilde{G}}$ is assumed to be negligible relative to all other masses.

e (Selectron) MASS LIMIT

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
> 97.5			04	OPAL	$\widetilde{e}_{\mbox{\it R}},\!\Delta m > 11$ GeV, $\left \mu\right > \!\! 100$ GeV, $\tan\beta \!\! = \!\! 1.5$
> 94.4			04	L3	\widetilde{e}_R , $\Delta m > 10$ GeV, $\left \mu \right > 200$ GeV, $\tan \beta > 2$
> 71.3		²¹² ACHARD (04	L3	\widetilde{e}_R , all Δm
none 30-94	95		03м	DLPH	$\Delta m > 15$ GeV, $\widetilde{e}_R^+ \widetilde{e}_R^-$
> 94	95	²¹⁴ ABDALLAH (03м	DLPH	\widetilde{e}_R , $1 \leq \tan \beta \leq 40$, $\Delta m > 10$ GeV
> 95	95	²¹⁵ HEISTER (02E	ALEP	$\Delta m > 15$ GeV, $\widetilde{e}_R^+ \widetilde{e}_R^-$
> 73	95		02N	ALEP	\widetilde{e}_{R} , any Δm
>107	95	²¹⁶ HEISTER (02N	ALEP	\widetilde{e}_L , any Δm
ullet $ullet$ We do	not use	the following data for	avei	rages, fit	ts, limits, etc. • • •
> 89	95		04F	OPAL	R, \widetilde{e}_{l}
> 92	95		04м	DLPH	R , \widetilde{e}_R , indirect, $\Delta m > 5$ GeV
> 93	95		03G	ALEP	\widetilde{e}_{R} , R decays, μ = -200 GeV, $\tan \beta$ = 2
> 69	95	²²⁰ ACHARD (02	L3	\widetilde{e}_R , R decays, μ = -200 GeV, $\tan \beta = \sqrt{2}$
> 92	95	²²¹ BARATE (01	ALEP	$\Delta m > 10$ GeV, $\widetilde{e}_R^+ \widetilde{e}_R^-$
> 77	95	²²² ABBIENDI (001	OPAL	$\Delta m > 5 \text{ GeV}, \ \widetilde{e}_R^+ \widetilde{e}_R^-$
> 83	95	²²³ ABREU (0 0 U	DLPH	Superseded by ABDALLAH 04M
> 67	95	224	00∨	DLPH	$\widetilde{e}_R \widetilde{e}_R (\widetilde{e}_R \rightarrow e \widetilde{G}), m_{\widetilde{G}} > 10 \text{ eV}$
> 85	95	²²⁵ BARATE (00G	ALEP	$\widetilde{\ell}_{R} ightarrow \ell \widetilde{G}$, any $ au(\widetilde{\ell}_{R})$
> 29.5	95	226	991	L3	\widetilde{e}_R , R , $\tan \beta \geq 2$
> 56	95	²²⁷ ACCIARRI 9	98F	L3	$\Delta m >$ 5 GeV, $\widetilde{e}_R^+ \widetilde{e}_R^-$, $ an eta \geq 1.41$
> 77	95	²²⁸ BARATE	98K	ALEP	Any Δm , $\widetilde{e}_{R}^{+}\widetilde{e}_{R}^{-}$, $\widetilde{e}_{R}^{-} \rightarrow e\gamma \widetilde{G}$
> 77	95		98	ZEUS	$m_{\widetilde{q}} = m_{\widetilde{e}}, \ m(\widetilde{\chi}_1^0) = 40 \text{ GeV}$
> 63	95	²³⁰ AID	96 C	H1	$m_{\widetilde{q}} = m_{\widetilde{e}}, m_{\widetilde{\chi}_1^0} = 35 \text{ GeV}$

- ²¹¹ ABBIENDI 04 search for $\widetilde{e}_R\widetilde{e}_R$ production in acoplanar di-electron final states in the 183–208 GeV data. See Fig. 13 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$ and for the limit at $\tan\beta$ =35 This limit supersedes ABBIENDI 00G.
- ACHARD 04 search for $\widetilde{e}_R\widetilde{e}_L$ and $\widetilde{e}_R\widetilde{e}_R$ production in single- and acoplanar di-electron final states in the 192–209 GeV data. Absolute limits on $m_{\widetilde{e}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 , $1 \le \tan\beta \le 60$ and $-2 \le \mu \le 2$ TeV. See Fig. 4 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. This limit supersedes ACCIARRI 99W.
- ²¹³ ABDALLAH 03M looked for acoplanar dielectron $+\cancel{E}$ final states at $\sqrt{s}=189$ –208 GeV. The limit assumes $\mu=-200$ GeV and $\tan\beta=1.5$ in the calculation of the production cross section and B($\widetilde{e} \rightarrow e \widetilde{\chi}_1^0$). See Fig. 15 for limits in the $(m_{\widetilde{e}_R}, m_{\widetilde{\chi}_1^0})$ plane. These limits include and update the results of ABREU 01
- ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of $M_2 < 1$ TeV, $|\mu| \le 1$ TeV with the $\widetilde{\chi}_1^0$ as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of $\tan\beta$. These limits update the results of ABREU 00W.
- ²¹⁵ HEISTER 02E looked for acoplanar dielectron $+ \not\!\!\!E_T$ final states from e^+e^- interactions between 183 and 209 GeV. The mass limit assumes $\mu < -200$ GeV and $\tan\beta = 2$ for the production cross section and B($\tilde{e} \rightarrow e \tilde{\chi}_1^0$)=1. See their Fig. 4 for the dependence of the limit on Δm . These limits include and update the results of BARATE 01.
- 216 HEISTER 02N search for $\widetilde{e}_R\widetilde{e}_L$ and $\widetilde{e}_R\widetilde{e}_R$ production in single- and acoplanar di-electron final states in the 183–208 GeV data. Absolute limits on $m_{\widetilde{e}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 , $1 \leq \tan\beta \leq 50$ and $-10 \leq \mu \leq 10$ TeV. The region of small $|\mu|$, where cascade decays are important, is covered by a search for $\widetilde{\chi}_1^0\widetilde{\chi}_3^0$ in final states with leptons and possibly photons. Limits on $m_{\widetilde{e}_L}$ are derived by exploiting the mass relation between the \widetilde{e}_L and \widetilde{e}_R , based on universal m_0 and $m_{1/2}$. When the constraint from the mass limit of the lightest Higgs from HEISTER 02 is included, the bounds improve to $m_{\widetilde{e}_R} > 77(75)$ GeV and $m_{\widetilde{e}_L} > 115(115)$ GeV for a top mass of 175(180) GeV. In the MSUGRA framework with radiative electroweak symmetry breaking, the limits improve further to $m_{\widetilde{e}_R} > 95$ GeV and $m_{\widetilde{e}_L} > 152$ GeV, assuming a trilinear coupling $A_0 = 0$ at the GUT scale. See Figs. 4, 5, 7 for the dependence of the limits on $\tan\beta$.
- ABBIENDI 04F use data from $\sqrt{s}=189-209$ GeV. They derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $LQ\overline{D}$ couplings. The results are valid for $\tan\beta=1.5,~\mu=-200$ GeV, with, in addition, $\Delta m>5$ GeV for indirect decays via $LQ\overline{D}$. The limit quoted applies to direct decays via $LL\overline{E}$ or $LQ\overline{D}$ couplings. For indirect decays, the limits on the \widetilde{e}_R mass are respectively 99 and 92 GeV for $LL\overline{E}$ and $LQ\overline{D}$ couplings and $m_{\widetilde{\chi}0}=10$ GeV and degrade slightly for larger $\widetilde{\chi}_1^0$ mass. Supersedes the results of ABBIENDI 00.
- ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or \overline{UDD} couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5$, $\Delta m>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect \overline{UDD} decays using the neutralino constraint of 39.5 GeV for $LL\overline{E}$ and of 38.0 GeV for \overline{UDD} couplings, also derived in ABDALLAH 04M. For indirect decays via $LL\overline{E}$ the limit improves to 95 GeV if the constraint from the neutralino is used and to 94 GeV if it is not used. For indirect decays via \overline{UDD} couplings it remains unchanged when the neutralino constraint is not used. Supersedes the result of ABREU 00U.

- 219 HEISTER 03G searches for the production of selectrons in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The search is performed for direct and indirect decays, assuming one coupling at a time to be non-zero. The limit holds for indirect decays mediated by $LQ\overline{D}$ couplings with $\Delta m>10$ GeV. Limits are also given for $LL\overline{E}$ direct $(m_{\widetilde{e},R}>96$ GeV) and indirect decays $(m_{\widetilde{e},R}>96$ GeV for $m(\widetilde{\chi}_1^0)>23$ GeV from BARATE 98S) and for \overline{UDD} indirect decays $(m_{\widetilde{e},R}>94$ GeV with $\Delta m>10$ GeV). Supersedes the results from BARATE 01B.
- 220 ACHARD 02 searches for the production of selectrons in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit holds for direct decays via $LL\overline{E}$ couplings. Stronger limits are reached for $LL\overline{E}$ indirect (79 GeV) and for \overline{UDD} direct or indirect (96 GeV) decays.
- BARATE 01 looked for acoplanar dielectron $+ \not\!\! E_T$ final states at 189 to 202 GeV. The limit assumes $\mu = -200$ GeV and $\tan\beta = 2$ for the production cross section and 100% branching ratio for $\stackrel{.}{e} \rightarrow e \stackrel{.}{\chi}^0_1$. See their Fig. 1 for the dependence of the limit on Δm . These limits include and update the results of BARATE 99Q.
- ^222 ABBIENDI 00J looked for acoplanar dielectron + E_T final states at \sqrt{s} = 161–183 GeV. The limit assumes $\mu < -100$ GeV and $\tan\beta$ =1.5 for the production cross section and decay branching ratios, evaluated within the MSSM, and zero efficiency for decays other than $\widetilde{e} \rightarrow e \widetilde{\chi}_1^0$. See their Fig. 12 for the dependence of the limit on Δm and $\tan\beta$.
- ABREU 000 studies decays induced by *R*-parity violating $LL\overline{E}$ couplings, using data from \sqrt{s} =189 GeV. They investigate topologies with multiple leptons, assuming one coupling at the time to be nonzero and giving rise to indirect decays. The limits assume a neutralino mass limit of 30 GeV, also derived in ABREU 000. Updates ABREU 001.
- ABREU 00V use data from \sqrt{s} = 130–189 GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as a function of $m_{\widetilde{G}}$, from a scan of the GMSB parameters space, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. For limits at different $m_{\widetilde{G}}$, see their Fig. 12.
- ²²⁵BARATE 00G combines the search for acoplanar dileptons, leptons with large impact parameters, kinks, and stable heavy-charged tracks, assuming 3 flavors of degenerate sleptons, produced in the schannel. Data collected at \sqrt{s} =189 GeV.
- ^226 ACCIARRI 99I establish indirect limits on $m_{\widetilde{e}_R}$ from the regions excluded in the M_2 versus m_0 plane by their chargino and neutralino searches at \sqrt{s} =130–183 GeV. The situations where the $\widetilde{\chi}_1^0$ is the LSP (indirect decays) and where a $\widetilde{\ell}$ is the LSP (direct decays) were both considered. The weakest limit, quoted above, comes from direct decays with \overline{UDD} couplings; $LL\overline{E}$ couplings or indirect decays lead to a stronger limit.
- ²²⁷ ACCIARRI 98F looked for acoplanar dielectron+ $\not\!\!E_T$ final states at \sqrt{s} =130–172 GeV. The limit assumes μ =-200 GeV, and zero efficiency for decays other than $\tilde{e}_R \to e \tilde{\chi}_1^0$. See their Fig. 6 for the dependence of the limit on Δm .
- BARATE 98K looked for $e^+e^-\gamma\gamma+\cancel{E}$ final states at $\sqrt{s}=$ 161–184 GeV. The limit assumes $\mu=-200$ GeV and $\tan\beta=2$ for the evaluation of the production cross section. See Fig. 4 for limits on the $(m_{\widetilde{e}_R},m_{\widetilde{\chi}^0_1})$ plane and for the effect of cascade decays.
- ²²⁹ BREITWEG 98 used positron+jet events with missing energy and momentum to look for $e^+ q \to \widetilde{e} \widetilde{q}$ via gaugino-like neutralino exchange with decays into $(e \widetilde{\chi}_1^0)(q \widetilde{\chi}_1^0)$. See paper for dependences in $m(\widetilde{q})$, $m(\widetilde{\chi}_1^0)$.
- ²³⁰ AID 96C used positron+jet events with missing energy and momentum to look for $e^+ q \rightarrow \widetilde{e} \, \widetilde{q} \,$ via neutralino exchange with decays into $(e \, \widetilde{\chi}_1^0)(q \, \widetilde{\chi}_1^0)$. See the paper for dependences on $m_{\widetilde{q}}$, $m_{\widetilde{\chi}_1^0}$.

$\widetilde{\mu}$ (Smuon) MASS LIMIT

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>91.0	95	²³¹ ABBIENDI	04	OPAL	$\Delta m >$ 3 GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$,
>86.7	95	²³² ACHARD	04	L3	$ \mu > 100$ GeV, $\tan \beta = 1.5$ $\Delta m > 10$ GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$
none 30–88	95	233 ABDALLAH		DLPH	$ \mu > 200$ GeV, $\tan \beta \ge 2$ $\Delta m > 5$ GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$
>94	95	²³⁴ ABDALLAH	03м	DLPH	$\widetilde{\mu}_{R,1} \leq aneta \leq a0, \ \Delta m > 10 \; {\sf GeV}$
>88	95	²³⁵ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

		²³⁶ ABAZOV	06ı	D0	R, λ'_{211}
>74	95	²³⁷ ABBIENDI	04F	OPAL	$R, \widetilde{\mu}_I$
>87	95	²³⁸ ABDALLAH	04M	DLPH	R , $\widetilde{\mu}_R$, indirect, $\Delta m > 5$ GeV
>81	95	²³⁹ HEISTER	03 G	ALEP	$\widetilde{\mu}_L$, \mathcal{R} decays
		²⁴⁰ ABAZOV	02н	D0	$\mathbb{R}, \lambda'_{211}$
>61	95	²⁴¹ ACHARD	02	L3	$\widetilde{\mu}_{R}$, R decays
>85	95	²⁴² BARATE	01		$\Delta m > 10$ GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$
>65	95	²⁴³ ABBIENDI	001	OPAL	$\Delta m > 2 \text{ GeV}, \ \widetilde{\mu}_R^+ \widetilde{\widetilde{\mu}_R}^-$
>80	95	²⁴⁴ ABREU	00V	DLPH	$\widetilde{\mu}_R \widetilde{\mu}_R \ (\widetilde{\mu}_R \to \mu \widetilde{G}), \ m_{\widetilde{G}} > 8$
>77	95	²⁴⁵ BARATE	98K		eV Any Δm , $\widetilde{\mu}_R^+\widetilde{\mu}_R^-$, $\widetilde{\mu}_R \to \mu\gamma\widetilde{G}$

- 231 ABBIENDI 04 search for $\widetilde{\mu}_R\widetilde{\mu}_R$ production in acoplanar di-muon final states in the 183–208 GeV data. See Fig. 14 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$ and for the limit at $\tan\beta$ =35. Under the assumption of 100% branching ratio for $\widetilde{\mu}_R \to \mu \ \widetilde{\chi}_1^0$, the limit improves to 94.0 GeV for $\Delta m >$ 4 GeV. See Fig. 11 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$ at several values of the branching ratio. This limit supersedes ABBIENDI 00G.
- ACHARD 04 search for $\widetilde{\mu}_R\widetilde{\mu}_R$ production in acoplanar di-muon final states in the 192–209 GeV data. Limits on $m_{\widetilde{\mu}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 , $1 \leq \tan\beta \leq 60$ and $-2 \leq \mu \leq 2$ TeV. See Fig. 4 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. This limit supersedes ACCIARRI 99W.
- ²³³ ABDALLAH 03M looked for acoplanar dimuon $+\cancel{E}$ final states at $\sqrt{s}=189$ –208 GeV. The limit assumes B($\widetilde{\mu}\to \mu\widetilde{\chi}_1^0$) = 100%. See Fig. 16 for limits on the $(m_{\widetilde{\mu}_R}, m_{\widetilde{\chi}_1^0})$ plane. These limits include and update the results of ABREU 01.
- ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of $M_2 < 1$ TeV, $|\mu| \le 1$ TeV with the $\widetilde{\chi}_1^0$ as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of $\tan\beta$. These limits update the results of ABREU 00W.
- ²³⁵ HEISTER 02E looked for acoplanar dimuon $+ \not\!\!E_T$ final states from e^+e^- interactions between 183 and 209 GeV. The mass limit assumes B($\widetilde{\mu} \to \mu \widetilde{\chi}_1^0$)=1. See their Fig. 4 for the dependence of the limit on Δm . These limits include and update the results of BARATE 01.

- 236 ABAZOV 06I looked in 380 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with at least 2 muons and 2 jets for s-channel production of $\widetilde{\mu}$ or $\widetilde{\nu}$ and subsequent decay via R couplings $LQ\overline{D}$. The data are in agreement with the SM expectation. They set limits on resonant slepton production and derive exclusion contours on λ'_{211} in the mass plane of $\widetilde{\ell}$ versus $\widetilde{\chi}_1^0$ assuming a MSUGRA model with $\tan\beta=5$, $\mu<0$ and $A_0=0$, see their Fig. 3. For $\lambda'_{211}\geq0.09$ slepton masses up to 358 GeV are excluded. Supersedes the results of ABAZOV 02H.
- ABBIENDI 04F use data from $\sqrt{s}=189$ –209 GeV. They derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $LQ\overline{D}$ couplings. The results are valid for $\tan\beta=1.5,\ \mu=-200$ GeV, with, in addition, $\Delta m>5$ GeV for indirect decays via $LQ\overline{D}$. The limit quoted applies to direct decays with $LL\overline{E}$ couplings and improves to 75 GeV for $LQ\overline{D}$ couplings. The limits on the $\widetilde{\mu}_R$ mass for indirect decays are respectively 94 and 87 GeV for $LL\overline{E}$ and $LQ\overline{D}$ couplings and $m_{\widetilde{\chi}0}=10$ GeV. Supersedes the results of ABBIENDI 00.
- ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or \overline{UDD} couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5$, $\Delta m>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect \overline{UDD} decays using the neutralino constraint of 39.5 GeV for $LL\overline{E}$ and of 38.0 GeV for \overline{UDD} couplings, also derived in ABDALLAH 04M. For indirect decays via $LL\overline{E}$ the limit improves to 90 GeV if the constraint from the neutralino is used and remains at 87 GeV if it is not used. For indirect decays via \overline{UDD} couplings it degrades to 85 GeV when the neutralino constraint is not used. Supersedes the result of ABREU 000.
- HEISTER 03G searches for the production of smuons in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The search is performed for direct and indirect decays, assuming one coupling at a time to be non-zero. The limit holds for direct decays mediated by R $LQ\overline{D}$ couplings and improves to 90 GeV for indirect decays (for $\Delta m>10$ GeV). Limits are also given for $LL\overline{E}$ direct ($m_{\widetilde{\mu}R}>87$ GeV) and indirect decays ($m_{\widetilde{\mu}R}>96$ GeV for $m(\widetilde{\chi}_1^0)>23$ GeV from BARATE 98S) and for \overline{UDD} indirect decays ($m_{\widetilde{\mu}R}>85$ GeV for $\Delta m>10$ GeV). Supersedes the results from BARATE 01B.
- ^240 ABAZOV 02H looked in 94 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for events with at least 2 muons and 2 jets for s-channel production of $\widetilde{\mu}$ or $\widetilde{\nu}$ and subsequent decay via R couplings $LQ\overline{D}$. A scan over the MSUGRA parameters is performed to exclude regions of the $(m_0,m_{1/2})$ plane, examples being shown in Fig. 2.
- ²⁴¹ ACHARD 02 searches for the production of smuons in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit holds for direct decays via $LL\overline{E}$ couplings. Stronger limits are reached for $LL\overline{E}$ indirect (87 GeV) and for \overline{UDD} direct or indirect (86 GeV) decays.
- ²⁴²BARATE 01 looked for acoplanar dimuon $+ \not\!\!E_T$ final states at 189 to 202 GeV. The limit assumes 100% branching ratio for $\widetilde{\mu} \to \mu \widetilde{\chi}_1^0$. See their Fig. 1 for the dependence of the limit on Δm . These limits include and update the results of BARATE 99Q.
- ²⁴³ ABBIENDI 00J looked for acoplanar dimuon $+ \not\!\! E_T$ final states at $\sqrt{s} = 161$ –183 GeV. The limit assumes B($\widetilde{\mu} \to \mu \widetilde{\chi}_1^0$)=1. Using decay branching ratios derived from the MSSM, a lower limit of 65 GeV is obtained for $\mu < -100$ GeV and $\tan \beta = 1.5$. See their Figs. 10 and 13 for the dependence of the limit on the branching ratio and on Δm .
- ABREU 00V use data from $\sqrt{s}=130-189$ GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. For limits at different $m_{\widetilde{G}}$, see their Fig. 12.
- ²⁴⁵ BARATE 98K looked for $\mu^+\mu^-\gamma\gamma+\cancel{E}$ final states at $\sqrt{s}=$ 161–184 GeV. See Fig. 4 for limits on the $(m_{\widetilde{\mu}_R},m_{\widetilde{\chi}_1^0})$ plane and for the effect of cascade decays.

$\widetilde{ au}$ (Stau) MASS LIMIT

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT	
>85.2	95	²⁴⁶ ABBIENDI	04	OPAL	$\Delta m~>$ 6 GeV, $ heta_{ au}{=}\pi/2$,	
>78.3	95	²⁴⁷ ACHARD	04	L3	$\begin{array}{l} \mu > 100 \; \mathrm{GeV}, \; \tan\!\beta = \! 1.5 \\ \Delta m > 15 \; \mathrm{GeV}, \; \theta_{\mathcal{T}} = \! \pi/2, \\ \mu > \! 200 \; \mathrm{GeV}, \; \tan\!\beta \; \geq \\ 2 \end{array}$	
>81.9	95	²⁴⁸ ABDALLAH	03м	DLPH	Δm $>$ 15 GeV, all $ heta_{ au}$	
none $m_{ au}-$ 26.3	95	²⁴⁸ ABDALLAH			$\Delta m > m_{ au}$, all $ heta_{ au}$	
>79	95	²⁴⁹ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $ heta_{ au} = \pi/2$	
>76	95	²⁴⁹ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $ heta_{ au} {=} 0.91$	
• • We do not use the following data for averages, fits, limits, etc. • • •						
		250			~	

>87.4	95	²⁵⁰ ABBIENDI	06 B	OPAL	$\widetilde{ au}_{m{R}} ightarrow \ au \widetilde{m{G}}$, all $ au(\widetilde{ au}_{m{R}})$
>74	95	²⁵¹ ABBIENDI	04F	OPAL	$R, \widetilde{\tau}_{l}$
>68	95 ²⁵²	^{2,253} ABDALLAH	04H	DLPH	AMSB, $\mu > 0$
>90	95	²⁵⁴ ABDALLAH	04M	DLPH	R , $\widetilde{\tau}_R$, indirect, $\Delta m >$ 5 GeV
>82.5		²⁵⁵ ABDALLAH	03 D	DLPH	$\widetilde{ au}_{m{R}} ightarrow \ au \widetilde{m{G}}$, all $ au(\widetilde{ au}_{m{R}})$
>70	95	²⁵⁶ HEISTER	03 G	ALEP	$\widetilde{ au}_R$, K decay
>61	95	²⁵⁷ ACHARD		L3	$\widetilde{ au}_{R}$, R decays
>77	95	²⁵⁸ HEISTER	02 R	ALEP	$ au_1$, any lifetime
>70	95	²⁵⁹ BARATE	01	ALEP	$\Delta m > 10$ GeV, $ heta_{ au} {=} \pi/2$
>68	95	²⁵⁹ BARATE	01	ALEP	$\Delta m > 10$ GeV, $ heta_{ au}^{\cdot} {=} 0.91$
>64	95	²⁶⁰ ABBIENDI	001	OPAL	$\Delta m > 10$ GeV, $\widetilde{\tau}_R^+ \widetilde{\tau}_R^-$
>84	95	²⁶¹ ABREU	00V	DLPH	$\widetilde{\ell}_R \widetilde{\ell}_R (\widetilde{\ell}_R \to \ell \widetilde{G}), m_{\widetilde{G}} > 9$
>73	95	²⁶² ABREU	00v	DLPH	$\stackrel{eV}{\widetilde{\tau}_1} \stackrel{eV}{\widetilde{\tau}_1} (\widetilde{\tau}_1 \to \ \tau \widetilde{G}), all \tau (\widetilde{\tau}_1)$
>52	95	²⁶³ BARATE			Any Δm , $\theta_{\tau} = \pi/2$, $\widetilde{\tau}_{R} \rightarrow$
•					$ au\gamma\widetilde{G}$
					'

²⁴⁶ ABBIENDI 04 search for $\widetilde{\tau}\widetilde{\tau}$ production in acoplanar di-tau final states in the 183–208 GeV data. See Fig. 15 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$ and for

the limit at $\tan\beta$ =35. Under the assumption of 100% branching ratio for $\widetilde{\tau}_R \to \tau \ \widetilde{\chi}_1^0$, the limit improves to 89.8 GeV for $\Delta m >$ 8 GeV. See Fig. 12 for the dependence of the limits on $\mathbf{m}_{\widetilde{\chi}_1^0}$ at several values of the branching ratio and for their dependence on θ_{τ} .

This limit supersedes ABBIENDI 00G.

 247 ACHARD 04 search for $\widetilde{\tau}\widetilde{\tau}$ production in acoplanar di-tau final states in the 192–209 GeV data. Limits on $m_{\widetilde{\tau}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and $m_0,~1~\leq \tan\beta \leq 60$ and $-2 \leq \mu \leq ~2$ TeV. See Fig. 4 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$.

²⁴⁸ ABDALLAH 03M looked for acoplanar ditaus $+\cancel{E}$ final states at $\sqrt{s}=130$ –208 GeV. A dedicated search was made for low mass $\widetilde{\tau}$ s decoupling from the Z^0 . The limit assumes B($\widetilde{\tau} \to \tau \widetilde{\chi}^0_1$) = 100%. See Fig. 20 for limits on the $(m_{\widetilde{\tau}}, m_{\widetilde{\chi}^0_1})$ plane and as function

of the $\widetilde{\chi}_1^0$ mass and of the branching ratio. The limit in the low-mass region improves to 29.6 and 31.1 GeV for $\widetilde{\tau}_R$ and $\widetilde{\tau}_L$, respectively, at $\Delta m > m_{\tau}$. The limit in the high-mass region improves to 84.7 GeV for $\widetilde{\tau}_R$ and $\Delta m > 15$ GeV. These limits include and update the results of ABREU 01.

HEISTER 02E looked for acoplanar ditau $+ \not\!\! E_T$ final states from e^+e^- interactions between 183 and 209 GeV. The mass limit assumes $\mathsf{B}(\widetilde{\tau} \to \tau \widetilde{\chi}_1^0) = 1$. See their Fig. 4 for the dependence of the limit on Δm . These limits include and update the results of BARATE 01.

- $250\,\mathrm{ABBIENDI}$ 06B use 600 pb $^{-1}$ of data from $\sqrt{s}=189$ –209 GeV. They look for events from pair-produced staus in a GMSB scenario with $\widetilde{\tau}$ NLSP including prompt $\widetilde{\tau}$ decays to ditaus $+\not\!\! E$ final states, large impact parameters, kinked tracks and heavy stable charged particles. Limits on the cross-section are computed as a function of $\mathrm{m}(\widetilde{\tau})$ and the lifetime, see their Fig. 7. The limit is compared to the $\sigma \cdot BR^2$ from a scan over the GMSB parameter space.
- ABBIENDI 04F use data from $\sqrt{s}=189$ –209 GeV. They derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $LQ\overline{D}$ couplings. The results are valid for $\tan\beta=1.5,\ \mu=-200$ GeV, with, in addition, $\Delta m>5$ GeV for indirect decays via $LQ\overline{D}$. The limit quoted applies to direct decays with $LL\overline{E}$ couplings and improves to 75 GeV for $LQ\overline{D}$ couplings. The limit on the $\widetilde{\tau}_R$ mass for indirect decays is 92 GeV for $LL\overline{E}$ couplings at $m_{\widetilde{\chi}0}=10$ GeV and no exclusion is obtained for $LQ\overline{D}$ couplings. Supersedes the results of ABBIENDI 00.
- ²⁵² ABDALLAH 04H use data from LEP 1 and $\sqrt{s}=192$ –208 GeV. They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region $1 < m_3/2 < 50$ TeV, $0 < m_0 < 1000$ GeV, $1.5 < \tan \beta < 35$, both signs of μ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z width of 3.2 MeV. The limit is for $m_t=174.3$ GeV (see Table 2 for other m_t values).
- 253 The limit improves to 75 GeV for $\mu~<$ 0.
- 254 ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5,~\Delta m~>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect decays using the neutralino constraint of 39.5 GeV, also derived in ABDALLAH 04M. For indirect decays via $LL\overline{E}$ the limit decreases to 86 GeV if the constraint from the neutralino is not used. Supersedes the result of ABREU 00U.
- ²⁵⁵ ABDALLAH 03D use data from $\sqrt{s}=130$ –208 GeV to search for tracks with large impact parameter or visible decay vertices and for heavy charged stable particles. Limits are obtained as function of m(\widetilde{G}), after combining these results with the search for slepton pair production in the SUGRA framework from ABDALLAH 03M to cover prompt decays. The above limit is reached for the stau decaying promptly, m(\widetilde{G}) < 6 eV, and is computed for stau mixing yielding the minimal cross section. Stronger limits are obtained for longer lifetimes, See their Fig. 9. Supersedes the results of ABREU 01G.
- ²⁵⁶ HEISTER 03G searches for the production of stau in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The search is performed for direct and indirect decays, assuming one coupling at a time to be non-zero. The limit holds for indirect decays mediated by R \overline{UDD} couplings with $\Delta m>10$ GeV. Limits are also given for $LL\overline{E}$ direct ($m_{\widetilde{T}_R}>87$ GeV) and indirect decays ($m_{\widetilde{T}_R}>95$ GeV for $m(\widetilde{\chi}_1^0)>23$ GeV from BARATE 98S) and for $LQ\overline{D}$ indirect decays ($m_{\widetilde{T}_R}>76$ GeV). Supersedes the results from BARATE 01B.
- ²⁵⁷ ACHARD 02 searches for the production of staus in the case of $\not R$ prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit holds for direct decays via $LL\overline{E}$ couplings. Stronger limits are reached for $LL\overline{E}$ indirect (86 GeV) and for \overline{UDD} direct or indirect (75 GeV) decays.
- 258 HEISTER 02R search for signals of GMSB in the 189–209 GeV data. For the $\widetilde{\chi}^0_1$ NLSP scenario, they looked for topologies consisting of $\gamma\gamma E\!\!\!\!/$ or a single γ not pointing to the interaction vertex. For the $\widetilde{\ell}$ NLSP case, the topologies consist of $\ell\ell E\!\!\!\!/$, including leptons with large impact parameters, kinks, or stable particles. Limits are derived from a scan over the GMSB parameters (see their Table 5 for the ranges). The limit remains valid whichever is the NLSP. The absolute mass bound on the $\widetilde{\chi}^0_1$ for any lifetime includes indirect limits from the slepton search HEISTER 02E preformed within the MSUGRA framework. A bound for any NLSP and any lifetime of 77 GeV has also been derived by using the constraints from the neutral Higgs search in HEISTER 02. In the co-NLSP

- scenario, limits $m_{\widetilde{e}_R} >$ 83 GeV (neglecting *t*-channel exchange) and $m_{\widetilde{\mu}_R} >$ 88 GeV are obtained independent of the lifetime. Supersedes the results from BARATE 00G.
- ²⁵⁹BARATE 01 looked for acoplanar ditau $+ \not\!\! E_T$ final states at 189 to 202 GeV. A slight excess (with 1.2% probability) of events is observed relative to the expected SM background. The limit assumes 100% branching ratio for $\tau \to \tau \tilde{\chi}_1^0$. See their Fig. 1 for the dependence of the limit on Δm . These limits include and update the results of BARATE 99Q.
- ABBIENDI 00J looked for acoplanar ditau $+ \not\!\! E_T$ final states at $\sqrt{s} = 161$ –183 GeV. The limit assumes B($\widetilde{\tau} \to \tau \widetilde{\chi}_1^0$)=1. Using decay branching ratios derived from the MSSM, a lower limit of 60 GeV at $\Delta m > 9$ GeV is obtained for $\mu < -100$ GeV and $\tan \beta = 1.5$. See their Figs. 11 and 14 for the dependence of the limit on the branching ratio and on Δm .
- ABREU 00V use data from $\sqrt{s}=130-189$ GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. The above limit assumes the degeneracy of stau and smuon. For limits at different $m_{\widetilde{G}}$, see their Fig. 12.
- ABREU 00V use data from $\sqrt{s}=130$ –189 GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. The above limit is reached for the stau mixing yielding the minimal cross section and decaying promptly. Stronger limits are obtained for longer lifetimes or for $\widetilde{\tau}_R$; see their Fig. 11. For $10 \leq m_{\widetilde{G}} \leq 310\,\mathrm{eV}$, the whole range $2 \leq m_{\widetilde{\tau}_1} \leq 80\,\mathrm{GeV}$ is excluded. Supersedes the results of ABREU 99C and ABREU 99F.
- $^{263}\,\text{BARATE}$ 98K looked for $\tau^+\,\tau^-\,\gamma\gamma+\cancel{E}$ final states at $\sqrt{s}{=}$ 161–184 GeV. See Fig. 4 for limits on the $(m_{\widetilde{\tau}_R},m_{\widetilde{\chi}_1^0})$ plane and for the effect of cascade decays.

Degenerate Charged Sleptons

Unless stated otherwise in the comment lines or in the footnotes, the following limits assume 3 families of degenerate charged sleptons.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>93	95	²⁶⁴ BARATE	01	ALEP	$\Delta m > 10$ GeV, $\widetilde{\ell}_R^+ \widetilde{\ell}_R^-$
>70	95	²⁶⁴ BARATE	01	ALEP	all Δm , $\widetilde{\ell}_R^+\widetilde{\ell}_R^-$
• • • We do not	use the follow	ing data for average	s, fits,	limits, e	etc. • • •
>91.9	95	²⁶⁵ ABBIENDI	06 B	OPAL	$\widetilde{\ell}_R o \ \ell \widetilde{G}$, all $\ell (\widetilde{\ell}_R)$
>88		²⁶⁶ ABDALLAH	03 D	DLPH	$\widetilde{\ell}_R \to \ell \widetilde{G}, all \ell(\widetilde{\ell}_R)$
>82.7	95	²⁶⁷ ACHARD	02	L3	$\widetilde{\ell}_{R}$, R decays,
>83	95	²⁶⁸ ABBIENDI	01	OPAL	$e^+e^- ightarrow\widetilde{\ell}_1\widetilde{\ell}_1, \ GMSB,\ tan\beta\!\!=\!\!2$
		²⁶⁹ ABREU	01	DLPH	$\widetilde{\ell} \rightarrow \ell \widetilde{\chi}_{2}^{0}, \widetilde{\chi}_{2}^{0} \rightarrow \gamma \widetilde{\chi}_{1}^{0},$
>68.8	95	²⁷⁰ ACCIARRI	01	L3	$\ell = e, \mu$ $\widetilde{\ell}_{R}, \mathcal{R}, 0.7 \le \tan\beta \le 40$
>84		1,272 ABREU	00V	DLPH	$\widetilde{\ell}_{R}\widetilde{\ell}_{R}\widetilde{\ell}_{R}(\widetilde{\ell}_{R} \to \ell\widetilde{G}),$ $m_{\widetilde{G}} > 9 \text{ eV}$

- 264 BARATE 01 looked for acoplanar dilepton $+ \not\!\!E_T$ and single electron (for $\tilde{e}_R \, \tilde{e}_L)$ final states at 189 to 202 GeV. The limit assumes $\mu{=}{-}200$ GeV and $\tan\beta{=}2$ for the production cross section and decay branching ratios, evaluated within the MSSM, and zero efficiency for decays other than $\tilde{\ell} \to \ell \, \tilde{\chi}_1^0$. The slepton masses are determined from the GUT relations without stau mixing. See their Fig. 1 for the dependence of the limit on Δm .
- ²⁶⁵ ABBIENDI 06B use 600 pb⁻¹ of data from $\sqrt{s}=189$ –209 GeV. They look for events from pair-produced staus in a GMSB scenario with ℓ co-NLSP including prompt ℓ decays to dileptons + ℓ final states, large impact parameters, kinked tracks and heavy stable charged particles. Limits on the cross-section are computed as a function of m(ℓ) and the lifetime, see their Fig. 7. The limit is compared to the $\sigma \cdot BR^2$ from a scan over the GMSB parameter space. The highest mass limit is reached for μ_R , from which the quoted mass limit is derived by subtracting m_{τ} .
- 266 ABDALLAH 03D use data from $\sqrt{s}=130\text{--}208$ GeV to search for tracks with large impact parameter or visible decay vertices and for heavy charged stable particles. Limits are obtained as function of $\mathrm{m}(\widetilde{G})$, after combining these results with the search for slepton pair production in the SUGRA framework from ABDALLAH 03M to cover prompt decays. The above limit is reached for prompt decays and assumes the degeneracy of the sleptons. For limits at different $\mathrm{m}(\widetilde{G})$, see their Fig. 9. Supersedes the results of ABREU 01G.
- 267 ACHARD 02 searches for the production of sparticles in the case of R prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at $\sqrt{s}{=}189{-}208$ GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale and no mixing in the slepton sector, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit holds for $LL\overline{E}$ couplings and increases to 88.7 GeV for \overline{UDD} couplings. For L3 limits from $LQ\overline{D}$ couplings, see ACCIARRI 01.
- ABBIENDI 01 looked for final states with $\gamma\gamma E$, $\ell\ell E$, with possibly additional activity and four leptons + E to search for prompt decays of $\widetilde{\chi}_1^0$ or $\widetilde{\ell}_1$ in GMSB. They derive limits in the plane $(m_{\widetilde{\chi}_1^0}, m_{\widetilde{\tau}_1})$, see Fig. 6, allowing either the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}_1$ to be the NLSP. Two scenarios are considered: $\tan\beta{=}2$ with the 3 sleptons degenerate in mass and $\tan\beta{=}20$ where the $\widetilde{\tau}_1$ is lighter than the other sleptons. Data taken at $\sqrt{s}{=}189~{\rm GeV}$. For $\tan\beta{=}20$, the obtained limits are $m_{\widetilde{\tau}_1}>69~{\rm GeV}$ and $m_{\widetilde{e}_1,\widetilde{\mu}_1}>88~{\rm GeV}$.
- ²⁶⁹ ABREU 01 looked for acoplanar dilepton + diphoton + $\not\!\!E$ final states from $\widetilde{\ell}$ cascade decays at \sqrt{s} =130–189 GeV. See Fig. 9 for limits on the (μ,M_2) plane for $m_{\widetilde{\ell}}$ =80 GeV, $\tan\beta$ =1.0, and assuming degeneracy of $\widetilde{\mu}$ and \widetilde{e} .
- ACCIARRI 01 searches for multi-lepton and/or multi-jet final states from R prompt decays with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings at $\sqrt{s}{=}189$ GeV. The search is performed for direct and indirect decays of neutralinos, charginos, and scalar leptons, with the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}$ as LSP and assuming one coupling to be nonzero at a time. Mass limits are derived using simultaneously the constraints from the neutralino, chargino, and slepton analyses; and the Z^0 width measurements from ACCIARRI 00C in a scan of the parameter space assuming MSUGRA with gaugino and scalar mass universality. Updates and supersedes the results from ACCIARRI 99I.
- ABREU 00V use data from $\sqrt{s}=130-189$ GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. For limits at different $m_{\widetilde{G}}$, see their Fig. 12.
- 272 The above limit assumes the degeneracy of stau and smuon.

Long-lived $\widetilde{\ell}$ (Slepton) MASS LIMIT

Limits on scalar leptons which leave detector before decaying. Limits from Z decays are independent of lepton flavor. Limits from continuum e^+e^- annihilation are also independent of flavor for smuons and staus. Selectron limits from e^+e^- collisions in the continuum depend on MSSM parameters because of the additional neutralino exchange contribution.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>98	95	²⁷³ ABBIENDI	03L	OPAL	$\widetilde{\mu}_{R}$, $\widetilde{ au}_{R}$
none 2-87.5	95	²⁷⁴ ABREU	00Q	DLPH	$\widetilde{\mu}_{R}$, $\widetilde{\tau}_{R}$
>81.2	95	²⁷⁵ ACCIARRI	99н	L3	$\widetilde{\mu}_{R}$, $\widetilde{\tau}_{R}$
>81	95	²⁷⁶ BARATE	98K	ALEP	$\widetilde{\mu}_{R}$, $\widetilde{\tau}_{R}$

- 273 ABBIENDI 03L used e^+e^- data at $\sqrt{s}=130$ –209 GeV to select events with two high momentum tracks with anomalous dE/dx. The excluded cross section is compared to the theoretical expectation as a function of the heavy particle mass in their Fig. 3. The limit improves to 98.5 GeV for $\widetilde{\mu}_L$ and $\widetilde{\tau}_L$. The bounds are valid for colorless spin 0 particles with lifetimes longer than 10^{-6} s. Supersedes the results from ACKERSTAFF 98P.
- ABREU 00Q searches for the production of pairs of heavy, charged stable particles in $e^+\,e^-$ annihilation at $\sqrt{s}{=}$ 130–189 GeV. The upper bound improves to 88 GeV for $\widetilde{\mu}_L$, $\widetilde{\tau}_L$. These limits include and update the results of ABREU 98P.
- ²⁷⁵ ACCIARRI 99H searched for production of pairs of back-to-back heavy charged particles at \sqrt{s} =130–183 GeV. The upper bound improves to 82.2 GeV for $\widetilde{\mu}_I$, $\widetilde{\tau}_I$.
- ²⁷⁶ The BARATE 98K mass limit improves to 82 GeV for $\widetilde{\mu}_L,\widetilde{\tau}_L$. Data collected at \sqrt{s} =161–184 GeV.

q (Squark) MASS LIMIT

degenerate, $m_{\widetilde{\alpha}}$ <45 GeV.

HTTP://PDG.LBL.GOV

For $m_{\widetilde{q}} >$ 60–70 GeV, it is expected that squarks would undergo a cascade decay via a number of neutralinos and/or charginos rather than undergo a direct decay to photinos as assumed by some papers. Limits obtained when direct decay is assumed are usually higher than limits when cascade decays are included.

Limits from e^+e^- collisions depend on the mixing angle of the lightest mass eigenstate $\widetilde{q}_1 = \widetilde{q}_R \sin\theta_q + \widetilde{q}_L \cos\theta_q$. It is usually assumed that only the sbottom and stop squarks have non-trivial mixing angles (see the stop and sbottom sections). Here, unless otherwise noted, squarks are always taken to be either left/right degenerate, or purely of left or right type. Data from Z decays have set squark mass limits above 40 GeV, in the case of $\widetilde{q} \to q\widetilde{\chi}_1$ decays if $\Delta m = m_{\widetilde{q}} - m_{\widetilde{\chi}_1^0} \gtrsim 5$ GeV. For smaller values of Δm , current constraints on the invisible width of the Z ($\Delta \Gamma_{\rm inv} < 2.0$ MeV, LEP 00) exclude $m_{\widetilde{u}_L,R} <$ 44 GeV, $m_{\widetilde{d}_R} <$ 33 GeV, $m_{\widetilde{d}_L} <$ 44 GeV and, assuming all squarks

Limits made obsolete by the most recent analyses of e^+e^- , $p\overline{p}$, and ep collisions can be found in previous Editions of this *Review*.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>325	95	²⁷⁷ ABAZOV	06 C	D0	jets+ E_T ,tan eta =3, μ <0, A_0 =0, any $m_{\widetilde{g}}$
> 99.5		²⁷⁸ ACHARD	04	L3	$\Delta m > 10$ GeV, $e^+e^- \rightarrow \widetilde{q}_{L,R} \overline{\widetilde{q}}_{L,R}$
> 97		²⁷⁸ ACHARD	04	L3	$\Delta m > 10 \text{ GeV}, e^+e^- \rightarrow \widetilde{q}_R \overline{\widetilde{q}}_R$

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>138	95	²⁷⁹ ABBOTT	01 D	D0	$\ell\ell+{ m jets}+ ot\!$
>255	95	²⁷⁹ ABBOTT	01 D	D0	$\tan \beta = 2$, $m_{\widetilde{g}} = m_{\widetilde{q}}$, $\mu < 0$, $A_0 = 0$, $\ell \ell + \text{jets} + \cancel{E}_T$
> 97	95	²⁸⁰ BARATE	01	ALEP	$e^+e^- \rightarrow \widetilde{q}\overline{\widetilde{q}}, \Delta m > 6 \text{ GeV}$
>250	95	²⁸¹ ABBOTT	99L	D0	$\tan \beta = 2$, $\mu < 0$, $A = 0$, $\text{jets} + \cancel{E}_T$
>224	95	²⁸² ABE	96D	CDF	$m_{\widetilde{g}} \leq m_{\widetilde{q}}$; with cascade
/ .	30	,	000	02.	decays, $\ell\ell+$ jets $+ ot\!$
• • • We do not	use the	following data for a	verage	es, fits, l	-
	95	²⁸³ CHEKANOV		ZEUS	$\widetilde{q} \rightarrow \mu q, R, LQ\overline{D}, \lambda=0.3$
>273	95 95	²⁸³ CHEKANOV		ZEUS	$\widetilde{q} \rightarrow \mu q, \not k, LQD, \lambda=0.3$ $\widetilde{q} \rightarrow \tau q, \not R, LQ\overline{D}, \lambda=0.3$
>270	95	²⁸⁴ AKTAS			
>275		²⁸⁴ AKTAS	04D	H1	$e^{\pm} p \rightarrow \widetilde{U}_{L}, R, LQ\overline{D}$
>280		285	04 D	H1	$e^{\pm} p \rightarrow \widetilde{D}_{R}, R, LQ\overline{D}$
		²⁸⁵ ADLOFF	03	H1	$e^{\pm} p \rightarrow \widetilde{q}, R, LQ\overline{D}$
>276	95	²⁸⁶ CHEKANOV	03 B	ZEUS	$\widetilde{d} \rightarrow e^- u, \nu d, R, LQ\overline{D}, \lambda > 0.1$
>260	95	286 CHEKANOV	03 B	ZEUS	$\widetilde{u} \rightarrow e^+ d, R, LQ\overline{D}, \lambda > 0.1$
> 82.5	95	²⁸⁷ HEISTER	03 G	ALEP	\widetilde{u}_{R} , R decay
> 77	95	²⁸⁷ HEISTER	03 G	ALEP	d_{R} , R decay
>240	95	²⁸⁸ ABAZOV	02F	D0	\widetilde{q} , $\Re \lambda'_{2jk}$ indirect decays,
					$\tan\!eta=2$, any $m_{\widetilde{g}}$
>265	95	²⁸⁸ ABAZOV	02F	D0	\tilde{q} , $R \lambda'_{2ik}$ indirect decays,
/200	33	71B/120 V	021	В	
		280			$\tan \beta = 2$, $m_{\widetilde{q}} = m_{\widetilde{g}}$
00.101		²⁸⁹ ABAZOV	02G		$p\overline{p} \to \widetilde{g}\widetilde{g},\widetilde{g}\widetilde{q}$
none 80–121	95	²⁹⁰ ABBIENDI	02	OPAL	$e\gamma \rightarrow \widetilde{u}_L$, $RLQ\overline{D}$, $\lambda=0.3$
none 80–158	95	²⁹⁰ ABBIENDI	02	OPAL	$e\gamma \rightarrow d_R$, $R LQ\overline{D}$, λ =0.3
none 80–185	95	²⁹¹ ABBIENDI	02 B	OPAL	$e\gamma \rightarrow \widetilde{u}_L$, $R LQ\overline{D}$, $\lambda=0.3$
none 80–196	95	²⁹¹ ABBIENDI	02 B	OPAL	$e\gamma \rightarrow d_R$, $RLQ\overline{D}$, λ =0.3
> 79	95	²⁹² ACHARD	02	L3	$\widetilde{\widetilde{u}}_{R}$, R decays
> 55	95	²⁹² ACHARD	02	L3	d_R , R decays
>263	95	²⁹³ CHEKANOV	02	ZEUS	
>258	95	²⁹³ CHEKANOV	02	ZEUS	$\widetilde{u}_{L} \rightarrow \tau q$, R , $LQ\overline{D}$, λ =0.3
> 82	95	²⁹⁴ BARATE	01 B	ALEP	\widetilde{u}_{R} , R decays
> 68	95	²⁹⁴ BARATE	01 B	ALEP	d_R , R decays
none 150–204	95	²⁹⁵ BREITWEG	01	ZEUS	$e^{+} p \rightarrow \widetilde{d}_{R}, R LQ\overline{D}, \lambda=0.3$
>200	95	²⁹⁶ АВВОТТ	00 C	D0	\widetilde{u}_L , \mathcal{R} , λ'_{2jk} decays
>180	95	²⁹⁶ АВВОТТ	00 C	D0	\tilde{d}_R , R , λ'_{2ik} decays
>390	95	²⁹⁷ ACCIARRI	00P	L3	$e^+e^- \rightarrow q\overline{q}$, R , $\lambda=0.3$
>148	95	²⁹⁸ AFFOLDER		CDF	~
		²⁹⁹ BARATE			
>200	95		001	ALEP	
none 150-269	95	300 BREITWEG			$e^+ p \rightarrow \widetilde{u}_L$, R , $LQ\overline{D}$, $\lambda=0.3$
>240	95	³⁰¹ ABBOTT	99	D0	$\widetilde{q} \rightarrow \widetilde{\chi}_2^0 X \rightarrow \widetilde{\chi}_1^0 \gamma X$
					$m_{\widetilde{\chi}_2^0} - m_{\widetilde{\chi}_1^0} > 20 \text{ GeV}$
>320	95	³⁰¹ ABBOTT	99	D0	$\widetilde{q} \rightarrow \widetilde{\chi}_1^0 X \rightarrow \widetilde{G} \gamma X$
>243	95	302 ABBOTT		D0	any $m_{\widetilde{g}}$, R , $ an \beta = 2$, $\mu < 0$
>200	95	303 ABE		CDF	$p\overline{p} \rightarrow \widetilde{q}\widetilde{q}, R$
/200	30	ADL	JJIVI	CDI	$pp \rightarrow qq, p$

none 80-134	95	³⁰⁴ ABREU	99 G	DLPH	$e\gamma \rightarrow \widetilde{u}_L$, $RLQ\overline{D}$, $\lambda=0.3$
none 80-161	95	³⁰⁴ ABREU			$e\gamma \rightarrow \widetilde{d}_{R}$, $R LQ\overline{D}$, λ =0.3
>225	95	³⁰⁵ ABBOTT	98E	D0	\widetilde{u}_L , \mathcal{R} , λ_{1ik}^{\prime} decays
>204	95	³⁰⁵ ABBOTT	98E	D0	\widetilde{d}_R , R , λ'_{1ik} decays
> 79	95	³⁰⁵ ABBOTT	98E	D0	\tilde{d}_L , \mathcal{R} , λ'_{ijk} decays
>202	95	³⁰⁶ ABE	985	CDF	\widetilde{u}_L , $\Re \lambda'_{2ik}$ decays
>160	95	³⁰⁶ ABE	98 S	CDF	\tilde{d}_R , $\Re \lambda_{2ik}'$ decays
>140	95	³⁰⁷ ACKERSTAFF	98V	OPAL	$e^+e^- \rightarrow q\overline{q}$, R , λ =0.3
> 77	95	³⁰⁸ BREITWEG	98	ZEUS	$m_{\widetilde{q}} = m_{\widetilde{e}}, \ m(\widetilde{\chi}_1^0) = 40 \text{ GeV}$
		309 DATTA	97	THEO	$\widetilde{\nu}$'s lighter than $\widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_2^0$
>216	95	³¹⁰ DERRICK	97	ZEUS	$ep ightarrow \ \widetilde{q}, \ \widetilde{q} ightarrow \ \mu j \ { m or} \ au j, R$
none 130-573	95	³¹¹ HEWETT	97	THEO	$q\widetilde{g} \rightarrow \widetilde{q}, \widetilde{q} \rightarrow q\widetilde{g}, \text{ with a}$
none 190-650	95	312 TEREKHOV	97	THEO	light gluino $qg \rightarrow \widetilde{q}\widetilde{g}, \widetilde{q} \rightarrow q\widetilde{g}, \text{ with a}$
> 63	95	313 AID	96 C	H1	light gluino $m_{\widetilde{q}} = m_{\widetilde{e}}, m_{\widetilde{\chi}_1^0} = 35 \text{ GeV}$
none 330–400	95	314 TEREKHOV	96	THEO	$ug \rightarrow \widetilde{u}\widetilde{g}, \widetilde{u} \rightarrow u\widetilde{g}$ with a
>176	95	315 ABACHI	95 C	D0	light gluino Any $m_{\widetilde{g}}$ <300 GeV; with cas-
		316 ABE	95T	CDF	cade decays $\widetilde{q} \to \widetilde{\chi}_2^0 \to \widetilde{\chi}_1^0 \gamma$
> 90	90	317 ABE	92L	CDF	Any $m_{\widetilde{g}}$ <410 GeV; with
,					cascade decay
>100		³¹⁸ ROY	92	RVUE	$p\overline{p} \rightarrow \widetilde{q}\widetilde{q}; \not R$
		³¹⁹ NOJIRI	91	COSM	

 277 ABAZOV 06C looked in 310 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with acoplanar jets or multijets with large E_T . No significant excess was found compared to the background expectation. A limit is derived on the masses of squarks and gluinos for specific MSUGRA parameter values, see Figure 3. Similar results would be obtained for a large class of parameter sets. Supersedes the results of ABBOTT 99L.

ACHARD 04 search for the production of $\widetilde{q}\widetilde{q}$ of the first two generations in acoplanar di-jet final states in the 192–209 GeV data. Degeneracy of the squark masses is assumed either for both left and right squarks or for right squarks only, as well as $B(\widetilde{q} \to q \widetilde{\chi}_1^0) = 1$ See Fig. 7 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. This limit supersedes ACCIARRI 99v.

ABBOTT 01D looked in $\sim 108~{\rm pb}^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for events with $e\,e,$ $\mu\,\mu,$ or $e\,\mu$ accompanied by at least 2 jets and E_T . Excluded regions are obtained in the MSUGRA framework from a scan over the parameters $0{<}m_0$ ${<}300~{\rm GeV},\,10{<}m_{1/2}$ ${<}110~{\rm GeV},\,$ and 1.2 ${<}{\rm tan}\beta$ ${<}10.$

²⁸⁰ BARATE 01 looked for acoplanar dijets $+ \not\!\!E_T$ final states at 189 to 202 GeV. The limit assumes B($\widetilde{q} \to q \widetilde{\chi}_1^0$)=1, with $\Delta m = m_{\widetilde{q}} - m_{\widetilde{\chi}_1^0}$. It applies to $\tan \beta = 4$, $\mu = -400$ GeV. See their Fig. 2 for the exclusion in the $(m_{\widetilde{q}}, m_{\widetilde{g}})$ plane. These limits include and update

the results of BARATE 99Q. 281 ABBOTT 99L consider events with three or more jets and large E_T . Spectra and decay rates are evaluated in the framework of minimal Supergravity, assuming five flavors of degenerate squarks, and scanning the space of the universal gaugino $(m_{1/2})$ and scalar (m_0) masses. See their Figs. 2–3 for the dependence of the limit on the relative value of

 $m_{\widetilde{q}}$ and $m_{\widetilde{g}}$.

²⁸² ABE 96D searched for production of gluinos and five degenerate squarks in final states containing a pair of leptons, two jets, and missing E_T . The two leptons arise from the

- semileptonic decays of charginos produced in the cascade decays. The limit is derived for fixed $\tan\beta=4.0$, $\mu=-400$ GeV, and $m_{H^+}=500$ GeV, and with the cascade decays of the squarks and gluinos calculated within the framework of the Minimal Supergravity scenario.
- CHEKANOV 05A search for lepton flavor violating processes $e^{\pm} p \rightarrow \ell X$, where $\ell = \mu$ or τ with high p_T , in 130 pb $^{-1}$ at 300 and 318 GeV. Such final states may originate from LQD couplings with simultaneously non-zero λ'_{1jk} and λ'_{ijk} (i=2 or 3). The quoted mass bounds hold for a u-type squark, assume a λ' of electromagnetic strength and contributions from only direct squark decays. For d-type squarks the bounds are strengthened to 278 and 275 GeV for the μ and τ final states, respectively. Supersedes the results of CHEKANOV 02.
- AKTAS 04D looked in 77.8 pb^{-1} of $e^{\pm}p$ collisions at $\sqrt{s}=319$ GeV for resonant production of \widetilde{q} by R-parity violating $LQ\overline{D}$ couplings assuming that one of the λ' couplings dominates over all others. They consider final states with or without leptons and/or jets and/or p_T' resulting from direct and indirect decays. They combine the channels to derive limits on λ'_{1j1} and λ'_{11k} as a function of the squark mass, see their Figs. 8 and 9, from a scan over the parameters $70 < M_2 < 350$ GeV, $-300 < \mu < 300$ GeV, $\tan\beta = 6$, for a fixed mass of 90 GeV for degenerate sleptons and an LSP mass > 30 GeV. The quoted limits refer to $\lambda' = 0.3$, with U=u,c,t and D=d,s,b. Supersedes the results of ADLOFF 01B.
- ADLOFF 03 looked for the s-channel production of squarks via R $LQ\overline{D}$ couplings in 117.2 pb^{-1} of e^+p data at $\sqrt{s}=301$ and 319 GeV and of e^-p data at $\sqrt{s}=319$ GeV. The comparison of the data with the SM differential cross section allows limits to be set on couplings for processes mediated through contact interactions. They obtain lower bounds on the value of $m_{\widetilde{q}}/\lambda'$ of 710 GeV for the process $e^+\bar{u}\to\widetilde{d}^k$ (and charge conjugate), mediated by $\lambda'_{1\,i\,i}$, and of 430 GeV for the process $e^+d\to\widetilde{u}^j$ (and charge conjugate), mediated by $\lambda'_{1\,i\,i}$.
- 286 CHEKANOV 03B used 131.5 pb^{-1} of e^+p and e^-p data taken at 300 and 318 GeV to look for narrow resonances in the eq or νq final states. Such final states may originate from $LQ\overline{D}$ couplings with non-zero λ'_{1j1} (leading to \widetilde{u}_j) or λ'_{11k} (leading to \widetilde{d}_k). See their Fig. 8 and explanations in the text for limits. The quoted mass bound assumes that only direct squark decays contribute.
- ²⁸⁷ HEISTER 03G searches for the production of squarks in the case of R prompt decays with \overline{UDD} direct couplings at at $\sqrt{s}=189$ –209 GeV.
- 288 ABAZOV 02F looked in 77.5 pb $^{-1}$ of $p\overline{p}$ collisions at 1.8 TeV for events with $\geq 2\mu + \geq$ 4jets, originating from associated production of squarks followed by an indirect R decay (of the $\widetilde{\chi}_1^0$) via $LQ\overline{D}$ couplings of the type λ'_{2jk} where j=1,2 and k=1,2,3. Bounds are obtained in the MSUGRA scenario by a scan in the range $0 \leq M_0 \leq 400$ GeV, $60 \leq m_{1/2} \leq 120$ GeV for fixed values $A_0=0$, $\mu<0$, and $\tan\beta=2$ or 6. The bounds are weaker for $\tan\beta=6$. See Figs. 2,3 for the exclusion contours in $m_{1/2}$ versus m_0 for $\tan\beta=2$ and 6, respectively.
- ABAZOV 02G search for associated production of gluinos and squarks in 92.7 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV, using events with one electron, \geq 4 jets, and large E_T . The results are compared to a MSUGRA scenario with μ <0, $A_0{=}0$, and $\tan\beta{=}3$ and allow to exclude a region of the $(m_0,m_{1/2})$ shown in Fig. 11.
- ABBIENDI 02 looked for events with an electron or neutrino and a jet in e^+e^- at 189 GeV. Squarks (or leptoquarks) could originate from a $LQ\overline{D}$ coupling of an electron with a quark from the fluctuation of a virtual photon. Limits on the couplings $\lambda'_{1j\,k}$ as a function of the squark mass are shown in Figs. 8–9, assuming that only direct squark decays contribute.

- ABBIENDI 02B looked for events with an electron or neutrino and a jet in e^+e^- at 189–209 GeV. Squarks (or leptoquarks) could originate from a $LQ\overline{D}$ coupling of an electron with a quark from the fluctuation of a virtual photon. Limits on the couplings λ'_{1jk} as a function of the squark mass are shown in Fig. 4, assuming that only direct squark decays contribute. The quoted limits are read off from Fig. 4. Supersedes the results of ABBIENDI 02.
- ACHARD 02 searches for the production of squarks in the case of R prompt decays with \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit holds for indirect decays. Stronger limits are reached for $(\widetilde{u}_R, \widetilde{d}_R)$ direct (80,56) GeV and $(\widetilde{u}_L, \widetilde{d}_L)$ direct or indirect (87,86) GeV decays.
- ²⁹³ CHEKANOV 02 search for lepton flavor violating processes $e^+p \to \ell X$, where $\ell=\mu$ or τ with high p_T , in 47.7 pb $^{-1}$ of e^+p collisions at 300 GeV. Such final states may originate from $LQ\overline{D}$ couplings with simultaneously nonzero $\lambda'_{1j\,k}$ and $\lambda'_{ij\,k}$ (i=2 or 3). The quoted mass bound assumes that only direct squark decays contribute.
- 294 BARATE 01B searches for the production of squarks in the case of $R\!\!\!\!/$ prompt decays with $LL\overline{E}$ indirect or \overline{UDD} direct couplings at $\sqrt{s}{=}189{-}202$ GeV. The limit holds for direct decays mediated by $R\!\!\!\!\!\!\!\!/$ \overline{UDD} couplings. Limits are also given for $LL\overline{E}$ indirect decays $(m_{\widetilde{U}_R}>90$ GeV and $m_{\widetilde{d}_R}>89$ GeV). Supersedes the results from BARATE 00H.
- BREITWEG 01 searches for squark production in 47.7 pb $^{-1}$ of e^+p collisions, mediated by R couplings $LQ\overline{D}$ and leading to final states with $\widetilde{\nu}$ and ≥ 1 jet, complementing the e^+X final states of BREITWEG 00E. Limits are derived on $\lambda'\sqrt{\beta}$, where β is the branching fraction of the squarks into $e^+q+\overline{\nu}q$, as function of the squark mass, see their Fig. 15. The quoted mass limit assumes that only direct squark decays contribute.
- ABBOTT 00C searched in $\sim 94~{\rm pb}^{-1}$ of $p\overline{p}$ collisions for events with $\mu\mu+{\rm jets}$, originating from associated production of leptoquarks. The results can be interpreted as limits on production of squarks followed by direct R decay via $\lambda'_{2j\,k}L_2Q_jd_k^c$ couplings. Bounds are obtained on the cross section for branching ratios of 1 and of 1/2, see their Fig. 4. The former yields the limit on the \widetilde{u}_L . The latter is combined with the bound of ABBOTT 99J from the $\mu\nu+{\rm jets}$ channel and of ABBOTT 98E and ABBOTT 98J from the $\nu\nu+{\rm jets}$ channel to yield the limit on \widetilde{d}_R .
- ²⁹⁷ ACCIARRI 00P studied the effect on hadronic cross sections of *t*-channel down-type squark exchange via *R*-parity violating coupling $\lambda_{1jk}^{'}L_1Q_jd_k^c$. The limit here refers to the case $j{=}1,2$, and holds for $\lambda_{1jk}^{'}{=}0.3$. Data collected at $\sqrt{s}{=}130{-}189$ GeV, superseding the results of ACCIARRI 98J.
- AFFOLDER 00K searched in $\sim 88\,\mathrm{pb}^{-1}$ of $p\overline{p}$ collisions for events with 2–3 jets, at least one being b-tagged, large E_T and no high p_T leptons. Such $\nu\nu+b$ -jets events would originate from associated production of squarks followed by direct R decay via $\lambda'_{ij3}L_iQ_jd_3^c$ couplings. Bounds are obtained on the production cross section assuming zero branching ratio to charged leptons.
- BARATE 001 studied the effect on hadronic cross sections and charge asymmetries of t-channel down-type squark exchange via R-parity violating coupling $\lambda'_{1jk}L_1Q_jd_k^c$. The limit here refers to the case j=1,2, and holds for λ'_{1jk} =0.3. A 50 GeV limit is found for up-type squarks with k=3. Data collected at \sqrt{s} = 130–183 GeV.
- 300 BREITWEG 00E searches for squark exchange in e^+p collisions, mediated by R couplings $LQ\overline{D}$ and leading to final states with an identified e^+ and ≥ 1 jet. The limit applies to up-type squarks of all generations, and assumes $B(\widetilde{q} \to q \, e) = 1$.
- 301 ABBOTT 99 searched for $\gamma \not\!\! E_T + \geq 2$ jet final states, and set limits on $\sigma(p\overline{p} \to q + X) \cdot B(\widetilde{q} \to \gamma \not\!\! E_T X)$. The quoted limits correspond to $m_{\widetilde{g}} \geq m_{\widetilde{q}}$, with $B(\widetilde{\chi}_2^0 \to q + X) \cdot B(\widetilde{q} \to q + X) \cdot B(\widetilde{q} \to q + X)$.

- $\widetilde{\chi}_1^0\gamma)=1$ and B($\widetilde{\chi}_1^0\to\widetilde{G}\gamma)=1$, respectively. They improve to 310 GeV (360 GeV in the case of $\gamma\widetilde{G}$ decay) for $m_{\widetilde{g}}=m_{\widetilde{g}}$.
- 302 ABBOTT 99K uses events with an electron pair and four jets to search for the decay of the $\widetilde{\chi}_1^0$ LSP via R $LQ\overline{D}$ couplings. The particle spectrum and decay branching ratios are taken in the framework of minimal supergravity. An excluded region at 95% CL is obtained in the $(m_0,m_{1/2})$ plane under the assumption that A_0 =0, μ <0, $\tan\beta$ =2 and any one of the couplings $\lambda_{1jk}'>10^{-3}$ (j=1,2 and k=1,2,3) and from which the above limit is computed. For equal mass squarks and gluinos, the corresponding limit is 277 GeV. The results are essentially independent of A_0 , but the limit deteriorates rapidly with increasing $\tan\beta$ or μ >0.
- 303 ABE 99M looked in 107 pb $^{-1}$ of $p\overline{p}$ collisions at \sqrt{s} =1.8 TeV for events with like sign dielectrons and two or more jets from the sequential decays $\widetilde{q} \to q\widetilde{\chi}_1^0$ and $\widetilde{\chi}_1^0 \to eq\overline{q}'$, assuming R coupling $L_1Q_jD_k^c$, with j=2,3 and k=1,2,3. They assume five degenerate squark flavors, B($\widetilde{q} \to q\widetilde{\chi}_1^0$)=1, B($\widetilde{\chi}_1^0 \to eq\overline{q}'$)=0.25 for both e^+ and e^- , and $m_{\widetilde{g}} \ge 200$ GeV. The limit is obtained for $m_{\widetilde{\chi}_1^0} \ge m_{\widetilde{q}}/2$ and improves for heavier gluinos or heavier χ_1^0 .
- ABREU 99G looked for events with an electron or neutrino and a jet in e^+e^- at 183 GeV. Squarks (or leptoquarks) could originate from a $LQ\overline{D}$ coupling of an electron with a quark from the fluctuation of a virtual photon. Limits on the couplings λ'_{1jk} as a function of the squark mass are shown in Fig. 4, assuming that only direct squark decays contribute.
- 305 ABBOTT 98E searched in $\sim 115~{\rm pb}^{-1}$ of $p\overline{p}$ collisions for events with $e\nu+{\rm jets}$, originating from associated production of squarks followed by direct R decay via $\lambda'_{1j\,k}L_1Q_jd^c_k$ couplings. Bounds are obtained by combining these results with the previous bound of ABBOTT 97B from the $ee+{\rm jets}$ channel and with a reinterpretation of ABACHI 96B $\nu\nu+{\rm jets}$ channel.
- 306 ABE 98S looked in $\sim 110\,\mathrm{pb}^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for events with $\mu\mu+\mathrm{jets}$ originating from associated production of squarks followed by direct R decay via $\lambda'_{2j\,k}L_2Q_jd^c_k$ couplings. Bounds are obtained on the production cross section times the square of the branching ratio, see Fig. 2. Mass limits result from the comparison with theoretical cross sections and branching ratio equal to 1 for \widetilde{u}_L and 1/2 for \widetilde{d}_R .
- 307 ACKERSTAFF 98V and ACCIARRI 98J studied the interference of t-channel squark (d_R) exchange via R-parity violating $\lambda_{1jk}' L_1 Q_j d_k^c$ coupling in $e^+e^- \to q \overline{q}$. The limit is for $\lambda_{1jk}' = 0.3$. See paper for related limits on \widetilde{u}_L exchange. Data collected at $\sqrt{s} = 130-172$ GeV.
- 308 BREITWEG 98 used positron+jet events with missing energy and momentum to look for $e^+ q \to \widetilde{e} \widetilde{q}$ via gaugino-like neutralino exchange with decays into $(e \widetilde{\chi}_1^0)(q \widetilde{\chi}_1^0)$. See paper for dependences in $m_{\widetilde{e}}$, $m_{\widetilde{\chi}_1^0}$.
- 309 DATTA 97 argues that the squark mass bound by ABACHI 95C can be weakened by 10–20 GeV if one relaxes the assumption of the universal scalar mass at the GUT-scale so that the $\widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_2^0$ in the squark cascade decays have dominant and invisible decays to $\widetilde{\chi}_1^0$
- 310 DERRICK 97 looked for lepton-number violating final states via *R*-parity violating couplings $\lambda'_{ijk}L_iQ_jd_k$. When $\lambda'_{11k}\lambda'_{ijk}\neq 0$, the process $eu\to \widetilde{d}_k^*\to \ell_iu_j$ is possible. When $\lambda'_{1j1}\lambda'_{ijk}\neq 0$, the process $e\overline{d}\to \widetilde{u}_j^*\to \ell_i\overline{d}_k$ is possible. 100% branching fraction $\widetilde{q}\to\ell j$ is assumed. The limit quoted here corresponds to $\widetilde{t}\to\tau q$ decay, with $\lambda'=0.3$. For different channels, limits are slightly better. See Table 6 in their paper.

- ³¹¹ HEWETT 97 reanalyzed the limits on possible resonances in di-jet mode $(\tilde{q} \rightarrow q\tilde{g})$ from ALITTI 93 quoted in "Limits for Excited q (q^*) from Single Production," ABE 96 in "SCALE LIMITS for Contact Interactions: $\Lambda(qqqq)$," and unpublished CDF, DØ bounds. The bound applies to the gluino mass of 5 GeV, and improves for lighter gluino. The analysis has gluinos in parton distribution function.
- 312 TEREKHOV 97 improved the analysis of TEREKHOV 96 by including di-jet angular distributions in the analysis.
- 313 AID 96C used positron+jet events with missing energy and momentum to look for $e^+ q \to \tilde{e} \, \tilde{q}$ via neutralino exchange with decays into $(e \, \tilde{\chi}^0_1)(q \, \tilde{\chi}^0_1)$. See the paper for dependences on $m_{\widetilde{e}}$, $m_{\widetilde{\chi}^0_1}$.
- ³¹⁴ TEREKHOV 96 reanalyzed the limits on possible resonances in di-jet mode $(\widetilde{u} \rightarrow u\widetilde{g})$ from ABE 95N quoted in "MASS LIMITS for g_A (axigluon)." The bound applies only to the case with a light gluino.
- 315 ABACHI 95C assume five degenerate squark flavors with $m_{\widetilde{q}_L} = m_{\widetilde{q}_R}$. Sleptons are assumed to be heavier than squarks. The limits are derived for fixed $\tan\beta = 2.0~\mu = -250~{\rm GeV}$, and $m_{H^+} = 500~{\rm GeV}$, and with the cascade decays of the squarks and gluinos calculated within the framework of the Minimal Supergravity scenario. The bounds are weakly sensitive to the three fixed parameters for a large fraction of parameter space. No limit is given for $m_{\rm gluino} > 547~{\rm GeV}$.
- $^{316}\,\text{ABE}$ 95T looked for a cascade decay of five degenerate squarks into $\widetilde{\chi}^0_2$ which further decays into $\widetilde{\chi}^0_1$ and a photon. No signal is observed. Limits vary widely depending on the choice of parameters. For $\mu=-40$ GeV, $\tan\beta=1.5$, and heavy gluinos, the range $50{<}m_{\widetilde{a}}$ (GeV)<110 is excluded at 90% CL. See the paper for details.
- 317 ABE 92L assume five degenerate squark flavors and $m_{\widetilde{q}_L} = m_{\widetilde{q}_R}$. ABE 92L includes the effect of cascade decay, for a particular choice of parameters, $\mu = -250$ GeV, $\tan\beta = 2$. Results are weakly sensitive to these parameters over much of parameter space. No limit for $m_{\widetilde{q}} \leq 50$ GeV (but other experiments rule out that region). Limits are 10–20 GeV higher if $\mathrm{B}(\widetilde{q} \to q \widetilde{\gamma}) = 1$. Limit assumes GUT relations between gaugino masses and the gauge coupling; in particular that for $|\mu|$ not small, $m_{\widetilde{\chi}_1^0} \approx m_{\widetilde{g}}/6$. This last relation implies that as $m_{\widetilde{g}}$ increases, the mass of $\widetilde{\chi}_1^0$ will eventually exceed $m_{\widetilde{q}}$ so that no decay is possible. Even before that occurs, the signal will disappear; in particular no bounds can be obtained for $m_{\widetilde{g}} > 410$ GeV. $m_{H^+} = 500$ GeV.
- 318 ROY 92 reanalyzed CDF limits on di-lepton events to obtain limits on squark production in R-parity violating models. The 100% decay $\widetilde{q} \to q \widetilde{\chi}$ where $\widetilde{\chi}$ is the LSP, and the LSP decays either into $\ell q \overline{d}$ or $\ell \ell \overline{e}$ is assumed.
- ³¹⁹ NOJIRI 91 argues that a heavy squark should be nearly degenerate with the gluino in minimal supergravity not to overclose the universe.

Long-lived \tilde{q} (Squark) MASS LIMIT

The following are bounds on long-lived scalar quarks, assumed to hadronise into hadrons with lifetime long enough to escape the detector prior to a possible decay. Limits may depend on the mixing angle of mass eigenstates: $\tilde{q}_1 = \tilde{q}_L \cos\theta_q + \tilde{q}_R \sin\theta_q$. The coupling to the Z^0 boson vanishes for up-type squarks when $\theta_u = 0.98$, and for down type squarks when $\theta_d = 1.17$.

VALUE (GeV) <u>CL%</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>

• • • We do not use the following data for averages, fits, limits, etc. • • •

>95	95	³²⁰ HEISTER	03H ALEP	\widetilde{u}
>92	95	³²⁰ HEISTER	03H ALEP	\widetilde{d}
none 2-85	95	³²¹ ABREU	98P DLPH	\widetilde{u}_I
none 2-81	95	³²¹ ABREU	98P DLPH	\widetilde{u}_R^-
none 2-80	95	³²¹ ABREU	98P DLPH	\widetilde{u} , θ_{II} =0.98
none 2-83	95	³²¹ ABREU	98P DLPH	\widetilde{d}_I
none 5-40	95	³²¹ ABREU	98P DLPH	\widetilde{d}_R^-
none 5-38	95	³²¹ ABREU	98P DLPH	\widetilde{d} , $\theta_d = 1.17$

 $^{^{320}\,\}mathrm{HEISTER}$ 03H use e^+e^- data at and around the Z^0 peak to look for hadronizing stable squarks. Combining their results on searches for charged and neutral R-hadrons with JANOT 03, a lower limit of 15.7 GeV on the mass is obtained. Combining this further with the results of searches for tracks with anomalous ionization in data from 183 to 208 GeV yields the quoted bounds.

\widetilde{b} (Sbottom) MASS LIMIT

Limits in e^+e^- depend on the mixing angle of the mass eigenstate $\widetilde{b}_1=\widetilde{b}_L\cos\theta_b+\widetilde{b}_R\sin\theta_b$. Coupling to the Z vanishes for $\theta_b\sim 1.17$. As a consequence, no absolute constraint in the mass region $\lesssim 40$ GeV is available in the literature at this time from e^+e^- collisions. In the Listings below, we use $\Delta m=m_{\widetilde{b}_1}-m_{\widetilde{\chi}_1^0}$.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
none 35-222	95	³²² ABAZOV	06 R	D0	$\widetilde{b} ightarrow \ b \widetilde{\chi}_1^0$
>220	95	³²³ ABULENCIA	061	CDF	$\widetilde{g} \rightarrow \widetilde{b} \widetilde{b}, \Delta m > 6 \text{ GeV}, \widetilde{b}_1 \rightarrow$
					$b\widetilde{\chi}^0_1$, $m_{\widetilde{arphi}}<$ 270 GeV
> 95	95	³²⁴ ACHARD	04	L3	$\widetilde{b} ightarrow b\widetilde{\chi}_{1}^{0}$, θ_{b} =0, Δm > 15–25 GeV
> 81	95	³²⁴ ACHARD	04	L3	$\widetilde{b} \rightarrow b\widetilde{\chi}_1^{\overline{0}}$, all θ_b , $\Delta m > 15$ –25 GeV
> 7.5	95	³²⁵ JANOT	04	THEO	unstable $\hat{ar{b}}_1$, $e^+e^- o$ hadrons
> 93	95	³²⁶ ABDALLAH	03M	DLPH	$\widetilde{b} \rightarrow b\widetilde{\chi}^{0}$, $\theta_{b}=0$, $\Delta m > 7$ GeV
> 76	95	³²⁶ ABDALLAH	03M	DLPH	$\widetilde{b} \rightarrow b\widetilde{\chi}^0$, all θ_b , $\Delta m > 7$ GeV
> 85.1	95	³²⁷ ABBIENDI	02H	OPAL	$\widetilde{b} ightarrow \ b \widetilde{\chi}^0_1$, all $ heta_{m{b}}$, $\Delta m > \! \! 10$ GeV,
> 89	95	328 HEISTER	02K	ALEP	CDF $\widetilde{b} \to b\widetilde{\chi}_1^0$, all θ_b , $\Delta m > 8$ GeV,
none 3.5-4.5	95	³²⁹ SAVINOV	01	CLEO	\widetilde{B} meson
none 80–145		³³⁰ AFFOLDER	00 D	CDF	$\widetilde{b} ightarrow \ b \widetilde{\chi}_1^0, \ m_{\widetilde{\chi}_1^0} <$ 50 GeV
• • • We do	not us	se the following data			fits, limits, etc. • •
> 78	95	³³¹ ABDALLAH	04M	DLPH	$R, \ \widetilde{b}_L, \ \text{indirect}, \ \Delta m > 5 \ \text{GeV}$
none 50-82	95	³³² ABDALLAH			$\widetilde{b} \to b\widetilde{g}$, stable \widetilde{g} , all θ_h ,
		333 BERGER	03	THEO	$\Delta m >$ 10 GeV
> 71.5	95	334 HEISTER			\widetilde{b}_I , R decay
> 27.4	95	335 HEISTER			$\widetilde{b} \rightarrow b\widetilde{g}$, stable \widetilde{g} or \widetilde{b}
> 48	95	³³⁶ ACHARD	02	L3	\widetilde{b}_1 , \mathcal{R} decays

³²¹ ABREU 98P assumes that 40% of the squarks will hadronise into a charged hadron, and 60% into a neutral hadron which deposits most of its energy in hadron calorimeter. Data collected at \sqrt{s} =130–183 GeV.

- 322 ABAZOV 06R looked in 310 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with 2 or 3 jets and large \cancel{E}_T with at least 1 b-tagged jet and a veto against isolated leptons. No excess is observed relative to the SM background expectations. Limits are set on the sbottom pair production cross-section under the assumption that the only decay mode is into $b\widetilde{\chi}_1^0$. Exclusion contours are derived in the plane of sbottom versus neutralino masses, shown in their Fig. 2. The observed limit is more constraining than the expected one due to a lack of events corresponding to large sbottom masses. Supersedes the results of ABBOTT 99F.
- 323 ABULENCIA 06I searched in 156 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for multijet events with large $\not\!\!E_T$. They request at least 2 b-tagged jets and no isolated leptons. They investigate the production of gluinos decaying into $\widetilde{b}_1 \, b$ followed by $\widetilde{b}_1 \to b \, \widetilde{\chi}_1^0$. Both branching fractions are assumed to be 100% and the LSP mass to be 60 GeV. No significant excess was found compared to the background expectation. Upper limits on the cross-section are extracted and a limit is derived on the masses of sbottom and gluinos, see their Fig.3.
- ³²⁴ ACHARD 04 search for the production of $\widetilde{b}\widetilde{b}$ in acoplanar b-tagged di-jet final states in the 192–209 GeV data. See Fig. 6 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. This limit supersedes ACCIARRI 99V.
- 325 JANOT 04 reanalyzes $e^+e^- \to hadrons$ total cross section data with $\sqrt{s} \in [20,209]$ GeV from PEP, PETRA, TRISTAN, SLC, and LEP and constrains the mass of \widetilde{b}_1 assuming it decays quickly to hadrons.
- 326 ABDALLAH 03M looked for \widetilde{b} pair production in events with acoplanar jets and $\not\!\!E$ at $\sqrt{s}=189$ –208 GeV. The limit improves to 87 (98) GeV for all θ_b ($\theta_b=0$) for $\Delta m>10$ GeV. See Fig. 24 and Table 11 for other choices of Δm . These limits include and update the results of ABREU,P 00D.
- ³²⁷ ABBIENDI 02H search for events with two acoplanar jets and p_T in the 161–209 GeV data. The limit assumes 100% branching ratio and uses the exclusion at large Δm from CDF (AFFOLDER 00D). For θ_b =0, the bound improves to > 96.9 GeV. See Fig. 4 and Table 6 for the more general dependence on the limits on Δm . These results supersede ABBIENDI 99M.
- 328 HEISTER 02K search for bottom squarks in final states with acoplanar jets with b tagging, using 183–209 GeV data. The mass bound uses the CDF results from AFFOLDER 00D. See Fig. 5 for the more general dependence of the limits on Δm . Updates BARATE 01.
- SAVINOV 01 use data taken at \sqrt{s} =10.52 GeV, below the $B\overline{B}$ threshold. They look for events with a pair of leptons with opposite charge and a fully reconstructed hadronic D or D^* decay. These could originate from production of a light-sbottom hadron followed by $\widetilde{B} \to D^{(*)} \ell^- \widetilde{\nu}$, in case the $\widetilde{\nu}$ is the LSP, or $\widetilde{B} \to D^{(*)} \pi \ell^-$, in case of R. The mass range $3.5 \le M(\widetilde{B}) \le 4.5$ GeV was explored, assuming 100% branching ratio for either of the decays. In the $\widetilde{\nu}$ LSP scenario, the limit holds only for $M(\widetilde{\nu})$ less than about 1 GeV and for the D^* decays it is reduced to the range 3.9–4.5 GeV. For the R decay, the whole range is excluded.
- 330 AFFOLDER 00D search for final states with 2 or 3 jets and $\not\!\!E_T$, one jet with a b tag. See their Fig. 3 for the mass exclusion in the $m_{\widetilde t}$, $m_{\widetilde \chi_1^0}$ plane.
- 331 ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with \overline{UDD} couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5,\,\Delta m>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect \overline{UDD} decays using the neutralino constraint of 38.0 GeV, also derived in

- ABDALLAH 04M, and assumes no mixing. For indirect decays it remains at 78 GeV when the neutralino constraint is not used. Supersedes the result of ABREU 01D.
- 332 ABDALLAH 03C looked for events of the type $q \overline{q} R^{\pm} R^{\pm}$, $q \overline{q} R^{\pm} R^{0}$, or $q \overline{q} R^{0} R^{0}$ in $e^{+} e^{-}$ interactions at $\sqrt{s} = 189$ –208 GeV. The R^{\pm} bound states are identified by anomalous dE/dx in the tracking chambers and the R^{0} by missing energy due to their reduced energy loss in the calorimeters. Excluded mass regions in the $(m(\widetilde{b}), m(\widetilde{g}))$ plane for $m(\widetilde{g}) > 2$ GeV are obtained for several values of the probability for the gluino to fragment into R^{\pm} or R^{0} , as shown in their Fig. 19. The limit improves to 94 GeV for $\theta_{b} = 0$.
- 333 BERGER 03 studies the constraints on a \widetilde{b}_1 with mass in the 2.2–5.5 GeV region coming from radiative decays of $\Upsilon(\text{nS})$ into sbottomonium. The constraints apply only if \widetilde{b}_1 lives long enough to permit formation of the sbottomonium bound state. A small region of mass in the $m_{\widetilde{b}_1}-m_{\widetilde{g}}$ plane survives current experimental constraints from CLEO.
- HEISTER 03G searches for the production of \widetilde{b} pairs in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The limit holds for indirect decays mediated by R \overline{UDD} couplings. It improves to 90 GeV for indirect decays mediated by R $LL\overline{E}$ couplings and to 80 GeV for indirect decays mediated by R $LQ\overline{D}$ couplings. Supersedes the results from BARATE 01B.
- ³³⁵ HEISTER 03H use their results on bounds on stable squarks, on stable gluinos and on squarks decaying to a stable gluino from the same paper to derive a mass limit on \tilde{b} , see their Fig. 13. The limit for a long-lived \tilde{b}_1 is 92 GeV.
- $\overline{^{336}}$ ACHARD 02 searches for the production of squarks in the case of \cancel{R} prompt decays with \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit is computed for the minimal cross section and holds for indirect decays and reaches 55 GeV for direct decays.
- 337 BAEK 02 studies the constraints on a \widetilde{b}_1 with mass in the 2.2–5.5 GeV region coming from precision measurements of Z^0 decays. It is noted that CP-violating couplings in the MSSM parameters relax the strong constraints otherwised derived from CP conservation.
- 338 BECHER 02 studies the constraints on a \widetilde{b}_1 with mass in the 2.2–5.5 GeV region coming from radiative B meson decays, and sets limits on the off-diagonal flavor-changing couplings $q \, \widetilde{b} \, \widetilde{g} \, (q = d, s)$.
- 339 CHEUNG 02B studies the constraints on a \tilde{b}_1 with mass in the 2.2–5.5 GeV region and a gluino in the mass range 12–16 GeV, using precision measurements of Z^0 decays and e^+e^- annihilations at LEP2. Few detectable events are predicted in the LEP2 data for the model proposed by BERGER 01.
- 340 CHO 02 studies the constraints on a \widetilde{b}_1 with mass in the 2.2–5.5 GeV region coming from precision measurements of Z^0 decays. Strong constraints are obtained for *CP*-conserving MSSM couplings.
- 341 BERGER 01 reanalyzed interpretation of Tevatron data on bottom-quark production. Argues that pair production of light gluinos ($m\sim 12$ –16 GeV) with subsequent 2-body decay into a light sbottom ($m\sim 2$ –5.5 GeV) and bottom can reconcile Tevatron data with predictions of perturbative QCD for the bottom production rate. The sbottom must either decay hadronically via a R-parity- and B-violating interaction, or be long-lived. Constraints on the mass spectrum are derived from the measurements of time-averaged B^0 – \overline{B}^0 mixing.
- 342 ABBOTT 99F looked for events with two jets, with or without an associated muon from b decay, and $\not\!\!E_T$. See Fig. 2 for the dependence of the limit on $m_{\widetilde{\chi}_1^0}$. No limit for $m_{\widetilde{\chi}_1^0} >$ 47 GeV.

\tilde{t} (Stop) MASS LIMIT

Limits depend on the decay mode. In e^+e^- collisions they also depend on the mixing angle of the mass eigenstate $\widetilde{t}_1=\widetilde{t}_L\cos\theta_t+\widetilde{t}_R\sin\theta_t$. The coupling to the Z vanishes when $\theta_t=0.98$. In the Listings below, we use $\Delta m\equiv m_{\widetilde{t}_1}-m_{\widetilde{\chi}_1^0}$ or $\Delta m\equiv m_{\widetilde{t}_1}-m_{\widetilde{\nu}}$, depending on relevant decay mode. See also bounds in " \widetilde{q} (Squark) MASS LIMIT." Limits made obsolete by the most recent analyses of e^+e^- and $p\overline{p}$ collisions can be found in previous Editions of this Review.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
none 80–134	95	³⁴³ ABAZOV	07 B	D0	$\widetilde{t} ightarrow \ c \widetilde{\chi}_1^0, \ m_{\widetilde{\chi}_1^0} < ext{48 GeV}$
none 80-120	95	³⁴⁴ ABAZOV	04	D0	$\widetilde{t} \rightarrow b\ell\nu\widetilde{\chi}^0, m_{\widetilde{\chi}^0} = 50 \text{ GeV}$
> 90		³⁴⁵ ACHARD	04	L3	$\widetilde{t} ightarrow \ c \widetilde{\chi}_1^0$, all $ heta_t^{\lambda}, \Delta m >$
> 93		³⁴⁵ ACHARD	04	L3	$b \rightarrow b\ell\tilde{\nu}$, all θ_t ,
> 88		³⁴⁵ ACHARD	04	L3	$\Delta m > 15 \text{ GeV}$ $\widetilde{b} \rightarrow b \tau \widetilde{\nu}$, all $\theta_t, \Delta m > 15 \text{ GeV}$
> 75	95	³⁴⁶ ABDALLAH	03м	DLPH	$\widetilde{t} \rightarrow c\widetilde{\chi}_{1}^{0}, \ \theta_{t}=0, \ \Delta m > 2 \text{ GeV}$
> 71	95	³⁴⁶ ABDALLAH	03м	DLPH	$\widetilde{t} \rightarrow c\widetilde{\chi}^0$, all θ_t , $\Delta m > 2$ GeV
> 96	95	³⁴⁶ ABDALLAH			$\widetilde{t} \rightarrow c\widetilde{\chi}^0, \theta_t = 0, \Delta m > 10 \text{ GeV}$
> 92	95	³⁴⁶ ABDALLAH			$\widetilde{t} \rightarrow c\widetilde{\chi}^0$, all θ_t , $\Delta m > 10$ GeV
none 80-131	95	³⁴⁷ ACOSTA			$\widetilde{t} \rightarrow b\ell\widetilde{\nu}, m_{\widetilde{\nu}} \leq 63 \text{ GeV}$
>144	95	³⁴⁸ ABAZOV	02 C		$\widetilde{t} \rightarrow b\ell \widetilde{\nu}, m_{\widetilde{\nu}} = 45 \text{ GeV}$
> 95.7	95	³⁴⁹ ABBIENDI	02H		$c\widetilde{\chi}_1^0$, all θ_t , $\Delta m>10$ GeV
> 92.6	95	³⁴⁹ ABBIENDI	02H		$b\ell\widetilde{\widetilde{\nu}}$, all θ_t , $\Delta m > 10$ GeV
> 91.5	95	³⁴⁹ ABBIENDI	02H	OPAL	$b\tau\widetilde{\nu}$, all θ_t , $\Delta m > 10$ GeV
> 63	95	³⁵⁰ HEISTER	02K	ALEP	
> 92	95	³⁵⁰ HEISTER	02K	ALEP	$\widetilde{t} \rightarrow c \widetilde{\chi}_{1}^{0}$, all θ_{t} , $\Delta m > 8$
> 97	95	³⁵⁰ HEISTER	02K	ALEP	GeV, \bar{CDF} $\widetilde{t} \to b\ell\widetilde{\nu}$, all θ_t , $\Delta m > 8$ GeV, $D \emptyset$
> 78	95	350 HEISTER	02K	ALEP	$\widetilde{t} ightarrow b \widetilde{\chi}_1^0 W^*$, all θ_t , $\Delta m > 8$ GeV

• • We do not use the following data for averages, fits, limits, etc.

none 80-122	95	³⁵¹ ABAZOV	04 B	D0	$\widetilde{t} ightarrow c \widetilde{\chi}^0$, $m_{\widetilde{\chi}^0} <$ 45 GeV
> 77	95	³⁵² ABBIENDI	04F		R , direct, all $\overset{\curvearrowright}{\theta_t}$
> 77	95	³⁵³ ABDALLAH	04M	DLPH	R , indirect, all θ_t ,
>122	95	354 ACOSTA	04 B	CDF	$\Delta m > 5$ GeV R , direct, all θ_t
		³⁵⁵ AKTAS	04 B	H1	R, \tilde{t}_1
> 74.5		356 DAS	04	THEO	$\widetilde{t}\widetilde{t} \rightarrow b\ell\nu_{\ell}\chi^{0}\overline{b}q\overline{q}'\chi^{0}, m_{\chi_{1}^{0}}$
					$=$ 15 GeV, no $\overline{t} ightarrow c \chi^0$
none 50-87	95	³⁵⁷ ABDALLAH	03 C	DLPH	$\widetilde{t} \rightarrow c\widetilde{g}$, stable \widetilde{g} , all θ_t ,
		358 CHAKRAB	03	THEO	$\Delta M > 10 \text{ GeV}$ $p \overline{p} \rightarrow \widetilde{t} \widetilde{t}^*, \text{RPV}$
> 71.5	95	359 HEISTER			\widetilde{t}_I , \mathcal{R} decay
> 80	95	³⁶⁰ HEISTER			$\widetilde{t} \rightarrow c\widetilde{g}$, stable \widetilde{g} or \widetilde{t} , all θ_t ,
> 77	95	³⁶¹ ACHARD	02	L3	all ΔM \widetilde{t}_1 , R decays
		³⁶² AFFOLDER		CDF	$t \rightarrow \tilde{t} \chi_1^0$

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> 61	95	³⁶³ ABREU	001	DLPH	$\mathbb{R}(LL\overline{E}), \theta_t = 0.98, \Delta m > 4$
none 68-119	95	³⁶⁴ AFFOLDER	00 D	CDF	$\widetilde{t} \rightarrow c \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} < 40 \text{ GeV}$
none 84-120	95	365 AFFOLDER	00 G	CDF	$\widetilde{t}_1 \rightarrow b\ell\widetilde{\nu}, m_{\widetilde{\nu}}^{-1} < 45$
> 59	95	³⁶⁶ BARATE	00P	ALEP	Repl. by HEISTER 02K
>120	95	³⁶⁷ ABE	99м	CDF	$ ota \overline{ ho} ightarrow \widetilde{t}_1 \widetilde{t}_1$, $ ota \overline{t}_1$
none 61–91	95	³⁶⁸ ABACHI	96 B	D0	$\widetilde{t} \rightarrow c\widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} < 30 \text{ GeV}$
none 9-24.4	95	³⁶⁹ AID	96	H1	$ep \rightarrow \widetilde{t}\widetilde{t}$, R decays
>138	95	³⁷⁰ AID	96	H1	$e p ightarrow \widetilde{t}$, $ R$, $ \lambda { m cos} heta_{t} > 0.03$
> 45		³⁷¹ СНО	96	RVUE	B^0 - \overline{B}^0 and ϵ , θ_t = 0.98,
		270			$ an\!eta<\!2$
none 11–41	95	³⁷² BUSKULIC	95E	ALEP	$R(LL\overline{E}), \theta_t=0.98$
none 6.0-41.2	95	AKERS	94K	OPAL	$\widetilde{t} \rightarrow c\widetilde{\chi}_1^0$, $\theta_t = 0$, $\Delta m > 2$ GeV
none 5.0-46.0	95	AKERS	94K	OPAL	$\tilde{t} \rightarrow c \tilde{\chi}_1^0, \theta_t = 0, \Delta m > 5 \text{ GeV}$
none 11.2-25.5	95	AKERS			$\widetilde{t} \rightarrow c\widetilde{\chi}_{1}^{\overline{0}}, \theta_{t}=0.98, \Delta m > 2$
					GeV
none 7.9–41.2	95	AKERS	94K	OPAL	$\widetilde{t} \rightarrow c\widetilde{\chi}_{1}^{0}, \ \theta_{t} = 0.98, \ \Delta m > 5$
none 7.6–28.0	95	³⁷³ SHIRAI	94	VNS	$\stackrel{GeV}{\widetilde{t}} ightarrow c \widetilde{\chi}^0_1$, any $ heta_t$, $\Delta m > \!\! 10$
none 10–20	95	³⁷³ SHIRAI	94	VNS	$\widetilde{t} ightarrow c \widetilde{\chi}_1^0$, any $ heta_t$, $\Delta m > 2.5$
					GeV

- 343 ABAZOV 07B looked in 360 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with a pair of acoplanar heavy-flavor jets with $\not\!\!E_T$. No excess is observed relative to the SM background expectations. Limits are set on the production of $\widetilde t_1$ under the assumption that the only decay mode is into $c\,\widetilde\chi^0_1$, see their Fig. 4 for the limit in the $(m_{\widetilde t},m_{\widetilde\chi^0_1})$ plane. No limit can be obtained for $m_{\widetilde\chi^0_1}>54$ GeV. Supersedes the results of ABAZOV 04B.
- 344 ABAZOV 04 looked at $108.3pb^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.8$ TeV for events with $e+\mu+E_T$ as signature for the 3- and 4-body decays of stop into $b\ell\nu\widetilde{\chi}^0$ final states. For the $b\ell\widetilde{\nu}$ channel they use the results from ABAZOV 02C. No significant excess is observed compared to the Standard Model expectation and limits are derived on the mass of \widetilde{t}_1 for the 3- and 4-body decays in the $(m_{\widetilde{t}}$, $m_{\widetilde{\chi}^0})$ plane, see their Figure 4.
- 345 ACHARD 04 search in the 192–209 GeV data for the production of $\widetilde{t}\widetilde{t}$ in acoplanar di-jet final states and, in case of $b\ell\widetilde{\nu}$ ($b\tau\widetilde{\nu}$) final states, two leptons (taus). The limits for $\theta_t=$ 0 improve to 95, 96 and 93 GeV, respectively. All limits assume 100% branching ratio for the respective decay modes. See Fig. 6 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. These limits supersede ACCIARRI 99V.
- ³⁴⁶ ABDALLAH 03M looked for \widetilde{t} pair production in events with acoplanar jets and \cancel{E} at \sqrt{s} = 189–208 GeV. See Fig. 23 and Table 11 for other choices of Δm . These limits include and update the results of ABREU,P 00D.
- 347 ACOSTA 03C searched in 107 pb^{-1} of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for pair production of \widetilde{t} followed by the decay $\widetilde{t} \to b\ell\widetilde{\nu}$. They looked for events with two isolated leptons (e or μ), at least one jet and E_T . The excluded mass range is reduced for larger $m_{\widetilde{\nu}}$, and no limit is set for $m_{\widetilde{\nu}} > 88.4$ GeV (see Fig. 2).
- 348 ABAZOV 02C looked in $108.3 \mathrm{pb}^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for events with $e\,\mu E_T$, originating from associated production $\widetilde{t}\,\widetilde{t}$. Branching ratios are assumed to be 100%. The bound for the $b\ell\widetilde{\nu}$ decay weakens for large $\widetilde{\nu}$ mass (see Fig. 3), and no limit is set when $m_{\widetilde{\nu}} > 85$ GeV. See Fig. 4 for the limits in case of decays to a real $\widetilde{\chi}_1^\pm$, followed by $\widetilde{\chi}_1^\pm \to \ell\widetilde{\nu}$, as a function of $m_{\widetilde{\chi}_1^\pm}$.

- 349 ABBIENDI 02H looked for events with two acoplanar jets, p_T , and, in the case of $b\ell\widetilde{\nu}$ final states, two leptons, in the 161–209 GeV data. The bound for $c\widetilde{\chi}_1^0$ applies to the region where $\Delta m < m_W + m_b$, else the decay $\widetilde{t}_1 \to b\widetilde{\chi}_1^0 W^+$ becomes dominant. The limit for $b\ell\widetilde{\nu}$ assumes equal branching ratios for the three lepton flavors and for $b\tau\widetilde{\nu}$ 100% for this channel. For θ_t =0, the bounds improve to > 97.6 GeV $(c\widetilde{\chi}_1^0)$, > 96.0 GeV $(b\ell\widetilde{\nu})$, and > 95.5 $(b\tau\widetilde{\nu})$. See Figs. 5–6 and Table 5 for the more general dependence of the limits on Δm . These results supersede ABBIENDI 99M.
- 350 HEISTER 02K search for top squarks in final states with jets (with/without b tagging or leptons) or long-lived hadrons, using 183–209 GeV data. The absolute mass bound is obtained by varying the branching ratio of $\widetilde{t} \to c \widetilde{\chi}_1^0$ and the lepton fraction in $\widetilde{t} \to b \widetilde{\chi}_1^0 f \overline{f}'$ decays. The mass bound for $\widetilde{t} \to c \widetilde{\chi}_1^0$ uses the CDF results from AFFOLDER 00D and for $\widetilde{t} \to b \ell \widetilde{\nu}$ the DØ results from ABAZOV 02C. See Figs. 2–5 for the more general dependence of the limits on Δm . Updates BARATE 01 and BARATE 00P.
- 351 ABAZOV 04B looked in 85.2 pb^{-1} of $p\overline{p}$ collisions at $\sqrt{s}=1.8$ TeV for events with at least two acoplanar jets and E_T . No significant excess is observed compared to the Standard Model expectation and a limit is derived on the production of \widetilde{t}_1 , see their Figure 2 for the limit in the $(m_{\widetilde{t}}, m_{\widetilde{\chi}^0})$ plane. No limit can be obtained for $m_{\widetilde{\chi}^0} > 52$ GeV
- 352 ABBIENDI 04F use data from $\sqrt{s}=189$ –209 GeV. They derive limits on the stop mass under the assumption of R with $LQ\overline{D}$ or \overline{UDD} couplings. The limit quoted applies to direct decays with \overline{UDD} couplings when the stop decouples from the Z^0 and improves to 88 GeV for $\theta_t=0$. For $LQ\overline{D}$ couplings, the limit improves to 98 (100) GeV for λ'_{13k} or λ'_{23k} couplings and all θ_t ($\theta_t=0$). For λ'_{33k} couplings it is 96 (98) GeV for all θ_t ($\theta_t=0$). Supersedes the results of ABBIENDI 00.
- 353 ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or \overline{UDD} couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5$, $\Delta m>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for decoupling of the stop from the Z^0 and indirect \overline{UDD} decays using the neutralino constraint of 39.5 GeV for $LL\overline{E}$ and of 38.0 GeV for \overline{UDD} couplings, also derived in ABDALLAH 04M. For no mixing (decoupling) and indirect decays via $LL\overline{E}$ the limit improves to 92 (87) GeV if the constraint from the neutralino is used and to 88 (81) GeV if it is not used. For indirect decays via \overline{UDD} couplings it improves to 87 GeV for no mixing and using the constraint from the neutralino, whereas it becomes 81 GeV (67) GeV for no mixing (decoupling) if the neutralino constraint is not used. Supersedes the result of ABREU 01D.
- 354 ACOSTA 04B looked in 106 pb^{-1} of $p\overline{p}$ collisions at $\sqrt{s}=1.8$ TeV for R-parity violating decays of \widetilde{t}_1 with $LQ\overline{D}$ couplings. They search for events of the type $\widetilde{t}_1\overline{\widetilde{t}}_1 \to \ell \tau_h jj$ where $\ell=e,\mu$ originates from a leptonic τ decay and τ_h represents a hadronic decay of τ . They derive limits on the stop mass for direct decays after combining the results from e and μ and under the assumption that BR =1 for the decay to τ .
- 355 AKTAS 04B looked in 106 pb^{-1} of $e^{\pm}p$ collisions at $\sqrt{s}=319$ GeV and 301 GeV for resonant production of \tilde{t}_1 by R-parity violating $LQ\overline{D}$ couplings couplings with λ'_{131} , others being zero. They consider the decays $\tilde{t}_1 \rightarrow e^+d$ and $\tilde{t}_1 \rightarrow W\tilde{b}$ followed by $\tilde{b} \rightarrow \overline{\nu}_e d$ and assume gauginos too heavy to participate in the decays. They combine the channels $j \, ep'_T$, $j \, \mu p'_T$, $jj \, jp'_T$ to derive limits in the plane $(m_{\tilde{t}}, \lambda'_{131})$, see their Fig. 5.
- $^{356}\,\mathrm{DAS}$ 04 reanalyzes AFFOLDER 00G data and obtains constraints on $m_{\widetilde{t}_1}$ as a function of $\mathrm{B}(\widetilde{t}\to b\ell\nu\chi^0)\times\mathrm{B}(\widetilde{t}\to b\overline{q}\,q'\chi^0),\ \mathrm{B}(\widetilde{t}\to c\chi^0)$ and $m_{\chi^0}.$ Bound weakens for larger $\mathrm{B}(\widetilde{t}\to c\chi^0)$ and $m_{\chi^0}.$

- 357 ABDALLAH 03C looked for events of the type $q\bar{q}R^\pm R^\pm$, $q\bar{q}R^\pm R^0$ or $q\bar{q}R^0R^0$ in e^+e^- interactions at $\sqrt{s}=189$ –208 GeV. The R^\pm bound states are identified by anomalous dE/dx in the tracking chambers and the R^0 by missing energy, due to their reduced energy loss in the calorimeters. Excluded mass regions in the $(m(\tilde{t}),m(\tilde{g}))$ plane for $m(\tilde{g})>2$ GeV are obtained for several values of the probability for the gluino to fragment into R^\pm or R^0 , as shown in their Fig. 18. The limit improves to 90 GeV for $\theta_t=0$.
- ³⁵⁸ Theoretical analysis of e^+e^-+2 jet final states from the RPV decay of $\widetilde{t}\widetilde{t}^*$ pairs produced in $p\overline{p}$ collisions at \sqrt{S} =1.8 TeV. 95%CL limits of 220 (165) GeV are derived for B($\widetilde{t} \rightarrow eq$)=1 (0.5).
- HEISTER 03G searches for the production of \widetilde{t} pairs in the case of R prompt decays with $LL\overline{E}$, $LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The limit holds for indirect decays mediated by R \overline{UDD} couplings. It improves to 91 GeV for indirect decays mediated by R $LL\overline{E}$ couplings, to 97 GeV for direct (assuming $B(\widetilde{t}_L \to q\tau)=100\%$) and to 85 GeV for indirect decays mediated by R $LQ\overline{D}$ couplings. Supersedes the results from BARATE 01B.
- 360 HEISTER 03H use e^+e^- data from 183–208 GeV to look for the production of stop decaying into a c quark and a stable gluino hadronizing into charged or neutral R-hadrons. Combining these results with bounds on stable squarks and on a stable gluino LSP from the same paper yields the quoted limit. See their Fig. 13 for the dependence of the mass limit on the gluino mass and on θ_t .
- $\overline{^{361}}$ ACHARD 02 searches for the production of squarks in the case of \cancel{R} prompt decays with \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The limit is computed for the minimal cross section and holds for both direct and indirect decays.
- 362 AFFOLDER 01B searches for decays of the top quark into stop and LSP, in $t\overline{t}$ events. Limits on the stop mass as a function of the LSP mass and of the decay branching ratio are shown in Fig. 3. They exclude branching ratios in excess of 45% for SLP masses up to 40 GeV.
- ³⁶³ ABREU 00I searches for the production of stop in the case of R-parity violation with $LL\overline{E}$ couplings, for which only indirect decays are allowed. They investigate topologies with jets plus leptons in data from \sqrt{s} =183 GeV. The lower bound on the stop mass assumes a neutralino mass limit of 27 GeV, also derived in ABREU 00I.
- 364 AFFOLDER 00D search for final states with 2 or 3 jets and $\not\!\!E_T$, one jet with a c tag. See their Fig. 2 for the mass exclusion in the $(m_{\widetilde t}, m_{\widetilde \chi_1^0})$ plane. The maximum excluded $m_{\widetilde t}$ value is 119 GeV, for $m_{\widetilde \chi_1^0} =$ 40 GeV.
- 365 AFFOLDER 00G searches for \widetilde{t}_1 \widetilde{t}_1^* production, with $\widetilde{t}_1 \to b\ell\widetilde{\nu}$, leading to topologies with ≥ 1 isolated lepton (e or μ), $\not\!\!E_T$, and ≥ 2 jets with ≥ 1 tagged as b quark by a secondary vertex. See Fig. 4 for the excluded mass range as a function of $m_{\widetilde{\nu}}$. Cross-section limits for \widetilde{t}_1 \widetilde{t}_1^* , with $\widetilde{t}_1 \to b\chi_1^\pm$ ($\chi_1^\pm \to \ell^\pm \nu \widetilde{\chi}_1^0$), are given in Fig. 2.
- 366 BARATE 00P use data from $\sqrt{s}=189-202$ GeV to explore the region of small mass difference between the stop and the neutralino by searching heavy stable charged particles or tracks with large impact parameters. For prompt decays, they make use of acoplanar jets from BARATE 99Q, updated up to 202 GeV. The limit is reached for $\Delta m=1.6$ GeV and a decay length of 1 cm. If the MSSM relation between the decay width and Δm is used, the limit improves to 63 GeV. It is set for $\Delta m=1.9$ GeV. $\tan\beta=2.6$, and $\theta_{\widetilde{t}}=0.98$, and large negative μ .
- 367 ABE 99M looked in 107 pb $^{-1}$ of $p\overline{p}$ collisions at \sqrt{s} =1.8 TeV for events with like sign dielectrons and two or more jets from the sequential decays $\widetilde{q} \to q\widetilde{\chi}_1^0$ and $\widetilde{\chi}_1^0 \to eq\overline{q}'$, assuming R coupling $L_1Q_jD_k^c$, with j=2,3 and k=1,2,3. They assume B($\widetilde{t}_1 \to c\widetilde{\chi}_1^0$)=1, B($\widetilde{\chi}_1^0 \to eq\overline{q}'$)=0.25 for both e^+ and e^- , and $m_{\widetilde{\chi}_1^0} \geq m_{\widetilde{t}_1}/2$. The limit improves for heavier $\widetilde{\chi}_1^0$.

- ³⁶⁸ ABACHI 96B searches for final states with 2 jets and missing E_T . Limits on $m_{\widetilde{t}}$ are given as a function of $m_{\widetilde{\chi}_1^0}$. See Fig. 4 for details.
- ³⁶⁹ AID 96 considers photoproduction of $\tilde{t}\tilde{t}$ pairs, with 100% *R*-parity violating decays of \tilde{t} to eq, with q=d, s, or b quarks.
- 370 AID 96 considers production and decay of \tilde{t} via the *R*-parity violating coupling $\lambda' L_1 \, Q_3 \, d_1^c$.
- 371 CHO 96 studied the consistency among the $B^0-\overline{B}^0$ mixing, ϵ in $K^0-\overline{K}^0$ mixing, and the measurements of V_{cb} , V_{ub}/V_{cb} . For the range 25.5 GeV $< m_{\widetilde t_1} < m_Z/2$ left by AKERS 94K for $\theta_t=0.98$, and within the allowed range in M_2 - μ parameter space from chargino, neutralino searches by ACCIARRI 95E, they found the scalar top contribution to $B^0-\overline{B}^0$ mixing and ϵ to be too large if $\tan\beta<2$. For more on their assumptions, see the paper and their reference 10.
- ³⁷² BUSKULIC 95E looked for $Z \to \widetilde{t}\overline{\widetilde{t}}$, where $\widetilde{t} \to c\chi_1^0$ and χ_1^0 decays via R-parity violating interactions into two leptons and a neutrino.
- $^{373}\,\text{SHIRAI}$ 94 bound assumes the cross section without the s-channel Z-exchange and the QCD correction, underestimating the cross section up to 20% and 30%, respectively. They assume $m_{\text{C}}\!=\!1.5\,\,\text{GeV}.$

Heavy \tilde{g} (Gluino) MASS LIMIT

For $m_{\widetilde{g}} > 60\text{--}70$ GeV, it is expected that gluinos would undergo a cascade decay via a number of neutralinos and/or charginos rather than undergo a direct decay to photinos as assumed by some papers. Limits obtained when direct decay is assumed are usually higher than limits when cascade decays are included. Limits made obsolete by the most recent analyses of $p\,\overline{p}$ collisions can be found in previous Editions of this *Review*.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>241	95	³⁷⁴ ABAZOV	06 C	D0	jets+ $ ot\!$
>337	95	³⁷⁴ ABAZOV	06 C	D0	$\text{jets+}\cancel{E}_T, an\beta=3, \\ \mu < 0, A_0=0, \ m_{\widetilde{s}}=m_{\widetilde{q}}$
>270	95	³⁷⁵ ABULENCIA	061	CDF	$\widetilde{g} \rightarrow \widetilde{b}b, \Delta m > 6 \text{ GeV}, \widetilde{b}_1 \rightarrow$
					$b\widetilde{\chi}_1^0$, $m_{\widetilde{b}_1}$ <220 GeV
>195	95	³⁷⁶ AFFOLDER	02	CDF	Jets $+ ot \!$
>300	95	³⁷⁶ AFFOLDER	02	CDF	$Jets + \not\!\! E_T, \ m_{\widetilde{q}} = m_{\widetilde{g}}$
>129	95	377 ABBOTT	01 D	D0	$\ell\ell+{ m jets}+E_T$, $ aneta<10$, $m_0<300$ GeV, $\mu<0$, $A_0=0$
>175	95	377 ABBOTT	01 D	D0	$\ell\ell$ +jets+ $\not\!\!E_T$, tan β =2, large m_0 , μ < 0, A_0 =0
>255	95	377 ABBOTT	01 D	D0	$\begin{array}{c} \ell\ell + \mathrm{jets} + E_T, \ \tan\beta = 2, \\ m_{\widetilde{g}} = m_{\widetilde{q}}, \ \mu < 0, \ A_0 = 0 \end{array}$
>168	95	³⁷⁸ AFFOLDER	01 J	CDF	$\ell\ell+$ Jets $+\cancel{E}_T$, $ aneta=2$, $\mu=-800$ GeV, $m_{\widetilde{g}}\gg m_{\widetilde{g}}$
>221	95	³⁷⁸ AFFOLDER	01 J	CDF	$\begin{array}{c} \mathbf{q} & \mathbf{g} \\ \ell\ell + Jets + \mathbf{E}_T, \ tan\beta = 2, \\ \mu = -800 \ GeV, \ m_{\widetilde{\boldsymbol{q}}} = m_{\widetilde{\boldsymbol{g}}} \end{array}$
>190	95	³⁷⁹ ABBOTT	99L	D0	Jets+ $\not\!\!E_T$, tan β =2, μ <0, A =0
>260	95	³⁷⁹ АВВОТТ	99L	D0	$ \int_{-\infty}^{\infty} \int_{T}^{\infty} m_{\widetilde{g}} = m_{\widetilde{q}} $

• • •	We do not use	the following data	for averages, fits,	limits, etc. • • •
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>224	95	³⁸⁰ ABAZOV	02F	D0	$\mathbb{R} \lambda_{2ik}'$ indirect decays,
					tan $eta=$ 2, any $m_{\widetilde{m{q}}}$
>265	95	³⁸⁰ ABAZOV	02F	D0	$\mathbb{R} \lambda'_{2jk}$ indirect decays,
					tan $eta{=}2$, $m_{\widetilde{m{q}}}{=}m_{\widetilde{m{g}}}$
		381 ABAZOV	02G	D0	$p\overline{p} ightarrow $
		382 CHEUNG	02 B	THEO	
		383 BERGER	01	THEO	$p\overline{p} \rightarrow X + b$ -quark
>240	95	³⁸⁴ ABBOTT	99	D0	$\widetilde{g} \rightarrow \widetilde{\chi}_2^0 X \rightarrow \widetilde{\chi}_1^0 \gamma X$,
					$m_{\widetilde{\chi}_2^0} - m_{\widetilde{\chi}_1^0} > 20 \text{ GeV}$
>320	95	³⁸⁴ ABBOTT	99	D0	$\widetilde{g} \rightarrow \widetilde{\chi}_1^0 X \rightarrow \widetilde{G} \gamma X$
>227	95	³⁸⁵ АВВОТТ	99K	D0	any $m_{\widetilde{q}}$, R , $ aneta=2$, $\mu<0$
>212	95	³⁸⁶ ABACHI	95 C	D0	$m_{\widetilde{g}} \geq m_{\widetilde{q}}$; with cascade decays
>144	95	³⁸⁶ ABACHI	95 C	D0	Any $m_{\widetilde{a}}$; with cascade decays
,		³⁸⁷ ABE	95T		$\widetilde{g} \rightarrow \widetilde{\chi}_2^0 \rightarrow \widetilde{\chi}_1^0 \gamma$
					. 2 1
010	00	388 HEBBEKER	93	RVUE	•
>218	90	³⁸⁹ ABE	92L	CDF	$m_{\widetilde{q}} \leq m_{\widetilde{g}}$; with cascade
> 100		³⁹⁰ ROY	92	RVUE	decay $p \overline{p} \rightarrow \widetilde{g} \widetilde{g}; R$
>100		³⁹¹ NOJIRI	92 91	COSM	$pp \rightarrow gg, p$
none 4–53	90	³⁹² ALBAJAR		UA1	Any $m_{\sim} > m_{\sim}$
none 4–75	90	³⁹² ALBAJAR		UA1	Any $m_{\widetilde{q}} > m_{\widetilde{g}}$
					$m_{\widetilde{q}} = m_{\widetilde{g}}$
none 16–58	90	³⁹³ ANSARI	87 D	UA2	$m_{\widetilde{q}} \lesssim 100 { m GeV}$

374 ABAZOV 06C looked in 310 pb $^{-1}$ of $p\bar{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with acoplanar jets or multijets with large E_T . No significant excess was found compared to the background expectation. A limit is derived on the masses of squarks and gluinos for specific MSUGRA parameter values, see Figure 3. Similar results would be obtained for a large class of parameter sets. Supersedes the results of ABBOTT 99L.

 376 AFFOLDER 02 searched in \sim 84 pb $^{-1}$ of $p\overline{p}$ collisions for events with \geq 3 jets and E_T , arising from the production of gluinos and/or squarks. Limits are derived by scanning the parameter space, for $m_{\widetilde{q}} \geq m_{\widetilde{g}}$ in the framework of minimal Supergravity, assuming five flavors of degenerate squarks, and for $m_{\widetilde{q}} < m_{\widetilde{g}}$ in the framework of constrained MSSM, assuming conservatively four flavors of degenerate squarks. See Fig. 3 for the variation of the limit as function of the squark mass. Supersedes the results of ABE 97K.

 377 ABBOTT 01D looked in $\sim 108~{\rm pb}^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for events with $e\,e,$ $\mu\,\mu,$ or $e\,\mu$ accompanied by at least 2 jets and E_T . Excluded regions are obtained in the MSUGRA framework from a scan over the parameters 0< m_0 <300 GeV, 10< $m_{1/2}$ <110 GeV, and 1.2 <tan β <10.

³⁷⁸ AFFOLDER 01J searched in $\sim 106~{\rm pb}^{-1}$ of $p\overline{p}$ collisions for events with 2 like-sign leptons (e or μ), \geq 2 jets and E_T , expected to arise from the production of gluinos and/or squarks with cascade decays into $\widetilde{\chi}^{\pm}$ or $\widetilde{\chi}^0_2$. Spectra and decay rates are evaluated

- in the framework of minimal Supergravity, assuming five flavors of degenerate squarks and a pseudoscalar Higgs mass m_A =500 GeV. The limits are derived for $\tan\beta$ =2, μ =-800 GeV, and scanning over $m_{\widetilde{g}}$ and $m_{\widetilde{q}}$. See Fig. 2 for the variation of the limit as function of the squark mass. These limits supersede the results of ABE 96D.
- 379 ABBOTT 99L consider events with three or more jets and large E_T . Spectra and decay rates are evaluated in the framework of minimal Supergravity, assuming five flavors of degenerate squarks, and scanning the space of the universal gaugino $(m_{1/2})$ and scalar (m_0) masses See their Figs. 2–3 for the dependence of the limit on the relative value of $m_{\widetilde{a}}$ and $m_{\widetilde{p}}$.
- 380 ABAZOV 02F looked in 77.5 pb $^{-1}$ of $p\overline{p}$ collisions at 1.8 TeV for events with $\geq 2\mu + \geq$ 4jets, originating from associated production of squarks followed by an indirect R decay (of the $\widetilde{\chi}_1^0$) via $LQ\overline{D}$ couplings of the type $\lambda'_{2j\,k}$ where j=1,2 and k=1,2,3. Bounds are obtained in the MSUGRA scenario by a scan in the range $0 \leq M_0 \leq 400$ GeV, $60 \leq m_{1/2} \leq 120$ GeV for fixed values $A_0=0$, $\mu<0$, and $\tan\beta=2$ or 6. The bounds are weaker for $\tan\beta=6$. See Figs. 2,3 for the exclusion contours in $m_{1/2}$ versus m_0 for $\tan\beta=2$ and 6. respectively.
- 381 ABAZOV 02G search for associated production of gluinos and squarks in 92.7 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV, using events with one electron, \geq 4 jets, and large E_T . The results are compared to a MSUGRA scenario with μ <0, $A_0{=}0$, and $\tan\beta{=}3$ and allow to exclude a region of the $(m_0,m_{1/2})$ shown in Fig. 11.
- 382 CHEUNG 02B studies the constraints on a \tilde{b}_1 with mass in the 2.2–5.5 GeV region and a gluino in the mass range 12–16 GeV, using precision measurements of Z^0 decays and e^+e^- annihilations at LEP2. Few detectable events are predicted in the LEP2 data for the model proposed by BERGER 01.
- 383 BERGER 01 reanalyzed interpretation of Tevatron data on bottom-quark production. Argues that pair production of light gluinos ($m\sim$ 12–16 GeV) with subsequent 2-body decay into a light sbottom ($m\sim$ 2–5.5 GeV) and bottom can reconcile Tevatron data with predictions of perturbative QCD for the bottom production rate. The sbottom must either decay hadronically via a R-parity- and B-violating interaction, or be long-lived.
- ³⁸⁴ ABBOTT 99 searched for $\gamma \not\!\! E_T + \geq 2$ jet final states, and set limits on $\sigma(p\overline{p} \to \widetilde{g} + X) \cdot B(\widetilde{g} \to \gamma \not\!\! E_T X)$. The quoted limits correspond to $m_{\widetilde{q}} \geq m_{\widetilde{g}}$, with $B(\widetilde{\chi}_2^0 \to \widetilde{\chi}_1^0 \gamma) = 1$ and $B(\widetilde{\chi}_1^0 \to \widetilde{G} \gamma) = 1$, respectively. They improve to 310 GeV (360 GeV in the case of $\gamma \, \widetilde{G}$ decay) for $m_{\widetilde{p}} = m_{\widetilde{g}}$.
- ABBOTT 99K uses events with an electron pair and four jets to search for the decay of the $\widetilde{\chi}_1^0$ LSP via \mathcal{R} $LQ\overline{D}$ couplings. The particle spectrum and decay branching ratios are taken in the framework of minimal supergravity. An excluded region at 95% CL is obtained in the $(m_0,m_{1/2})$ plane under the assumption that A_0 =0, μ <0, $\tan\beta$ =2 and any one of the couplings $\lambda'_{1jk} > 10^{-3}$ (j=1,2 and k=1,2,3) and from which the above limit is computed. For equal mass squarks and gluinos, the corresponding limit is 277 GeV. The results are essentially independent of A_0 , but the limit deteriorates rapidly with increasing $\tan\beta$ or μ >0.
- 386 ABACHI 95C assume five degenerate squark flavors with with $m_{\widetilde{q}_L} = m_{\widetilde{q}_R}$. Sleptons are assumed to be heavier than squarks. The limits are derived for fixed $\tan\beta = 2.0~\mu = -250~\text{GeV}$, and $m_{H^+} = 500~\text{GeV}$, and with the cascade decays of the squarks and gluinos calculated within the framework of the Minimal Supergravity scenario. The bounds are weakly sensitive to the three fixed parameters for a large fraction of parameter space.
- 387 ABE 95T looked for a cascade decay of gluino into $\widetilde{\chi}^0_2$ which further decays into $\widetilde{\chi}^0_1$ and a photon. No signal is observed. Limits vary widely depending on the choice of parameters. For $\mu=-40$ GeV, $\tan\beta=1.5$, and heavy squarks, the range $50 < m_{\widetilde{g}}$ (GeV) <140 is excluded at 90% CL. See the paper for details.

- 388 HEBBEKER 93 combined jet analyses at various $e^+\,e^-$ colliders. The 4-jet analyses at TRISTAN/LEP and the measured $\alpha_{\rm S}$ at PEP/PETRA/TRISTAN/LEP are used. A constraint on effective number of quarks $N\!\!=\!6.3\pm1.1$ is obtained, which is compared to that with a light gluino, $N\!\!=\!8.$
- 389 ABE 92L bounds are based on similar assumptions as ABACHI 95C. Not sensitive to $m_{\rm gluino}$ <40 GeV (but other experiments rule out that region).
- 390 ROY 92 reanalyzed CDF limits on di-lepton events to obtain limits on gluino production in *R*-parity violating models. The 100% decay $\tilde{g} \to q \overline{q} \tilde{\chi}$ where $\tilde{\chi}$ is the LSP, and the LSP decays either into $\ell q \overline{d}$ or $\ell \ell \overline{e}$ is assumed.
- 391 NOJIRI 91 argues that a heavy gluino should be nearly degenerate with squarks in minimal supergravity not to overclose the universe.
- The limits of ALBAJAR 87D are from $p\overline{p} \to \widetilde{g}\widetilde{g}X$ ($\widetilde{g} \to q\overline{q}\widetilde{\gamma}$) and assume $m_{\widetilde{q}} > m_{\widetilde{g}}$. These limits apply for $m_{\widetilde{\gamma}} \lesssim 20$ GeV and $\tau(\widetilde{g}) < 10^{-10}$ s.
- 393 The limit of ANSARI 87D assumes $m_{\widetilde{q}} > m_{\widetilde{g}}$ and $m_{\widetilde{\gamma}} pprox 0$.

Long-lived/light \tilde{g} (Gluino) MASS LIMIT

Limits on light gluinos ($m_{\widetilde{g}}$ < 5 GeV), or gluinos which leave the detector before decaying.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
ullet $ullet$ We do not use the following data for averages, fits, limits, etc. $ullet$ $ullet$					
>12		³⁹⁴ BERGER			hadron scattering data
none 2-18	95	³⁹⁵ ABDALLAH	03 C	DLPH	$e^+e^- ightarrow \ q\overline{q}\widetilde{g}\widetilde{g}$, stable \widetilde{g}
> 5		³⁹⁶ ABDALLAH	03 G	DLPH	QCD beta function
		³⁹⁷ HEISTER	03	ALEP	Color factors
>26.9	95	³⁹⁸ HEISTER	03H	ALEP	$e^+e^- ightarrow q \overline{q} \widetilde{g} \widetilde{g}$
> 6.3		³⁹⁹ ЈАNОТ	03	RVUE	$\Delta\Gamma_{had}$ <3.9 MeV
		⁴⁰⁰ MAFI	00	THEO	$pp o {\sf jets} + p_T'$
		⁴⁰¹ ALAVI-HARA	TI99E	KTEV	$pN \rightarrow R^0$, with $R^0 \rightarrow \rho^0 \widetilde{\gamma}$
		400			and $\mathit{R}^0 ightarrow \ \pi^0 \widetilde{\gamma}$
		402 BAER			Stable \widetilde{g} hadrons
					$ hoBe o \ \mathit{R}^0 o \ \eta\widetilde{\gamma}$
		⁴⁰⁴ ACKERSTAF	F 98V	OPAL	$e^+e^- \rightarrow \widetilde{\chi}_1^+\widetilde{\chi}_1^-$
					$pN \rightarrow R^0 \stackrel{1}{\rightarrow} \rho^0 \widetilde{\gamma}$
		406 ALBUQUERG	97	E761	$R^+(uud\widetilde{g}) \rightarrow S^0(uds\widetilde{g})\pi^+,$
					$X^-(ssd\widetilde{g}) ightarrow S^0\pi^-$
> 6.3	95	407 BARATE			Color factors
> 5	99	408 CSIKOR	97	RVUE	eta function, $Z o$ jets
> 1.5	90	409 DEGOUVEA	97	THEO	$Z \rightarrow jjjj$
		⁴¹⁰ FARRAR			
none 1.9–13.6	95	⁴¹¹ AKERS	95 R	OPAL	Z decay into a long-lived
		410 -			$(\widetilde{g}q\overline{q})^{\pm}$
< 0.7		412 CLAVELLI	95		quarkonia
none 1.5–3.5			94	RVUE	$\Upsilon(1S) ightarrow \gamma + gluinonium$
not 3–5		414 LOPEZ	93 C		
≈ 4		415 CLAVELLI	92		$\alpha_{\it S}$ running
		416 ANTONIADIS	91	RVUE	$\alpha_{\it S}$ running
> 1		411 ANTONIADIS	91		$pN \rightarrow \text{missing energy}$
		⁴¹⁸ NAKAMURA	89	SPEC	R - Δ^{++}

```
<sup>419</sup> ARNOLD
                                                                      EMUL \pi^- (350 GeV). \sigma \simeq A^1
> 3.8
                          90
                                   <sup>419</sup> ARNOLD
                                                                      EMUL \pi^{-} (350 GeV). \sigma \simeq A^{0.72}
                          90
> 3.2
                                   <sup>420</sup> TUTS
                                                                      CUSB \Upsilon(1S) 
ightarrow \gamma + \mathsf{gluinonium}
                          90
none 0.6-2.2
                                                              86C ARG 1 \times 10^{-11} \lesssim \tau \lesssim 1 \times 10^{-9} \text{s} 86 BDMP 1 \times 10^{-10} < \tau < 1 \times 10^{-7} \text{s}
                                   <sup>421</sup> ALBRECHT
none 1 -4.5
                          90
                                   <sup>422</sup> BADIER
                          90
none 1-4
                                   <sup>423</sup> BARNETT
                                                                      RVUE p\overline{p} \rightarrow \text{gluino gluino gluon}
none 3-5
                                   424 VOLOSHIN
                                                                      RVUE If (quasi) stable; \tilde{g} u u d
none
                                   <sup>425</sup> COOPER-...
                                                              85B
none 0.5-2
                                                                     BDMP For m_{\widetilde{a}}=300 GeV
                                   <sup>425</sup> COOPER-...
none 0.5-4
                                                              85B
                                                                      BDMP For m_{\widetilde{a}} <65 GeV
                                   <sup>425</sup> COOPER-...
                                                              85B
                                                                      BDMP For m_{\tilde{a}} = 150 \text{ GeV}
none 0.5-3
                                   <sup>426</sup> DAWSON
                                                                      RVUE \tau > 10^{-7} \text{ s}
none 2-4
                                   <sup>426</sup> DAWSON
                                                                      RVUE For m_{\widetilde{a}} = 100 \text{ GeV}
none 1-2.5
                                                              85
                                   <sup>427</sup> FARRAR
                          90
none 0.5-4.1
                                                                      RVUE FNAL beam dump
                                   <sup>428</sup> GOLDMAN
> 1
                                                              85
                                                                      RVUE Gluinonium
                                   <sup>429</sup> HABER
                                                                      RVUE
                                                              85
>1-2
                                   <sup>430</sup> BALL
                                                                      CALO
                                   <sup>431</sup> BRICK
                                                              84
                                                                      RVUE
                                   <sup>432</sup> FARRAR
                                                              84
                                                                      RVUE
                                   <sup>433</sup> BERGSMA
                                                                     RVUE For m_{\widetilde{q}} < 100 GeV
> 2
                                                              83C
                                   <sup>434</sup> CHANOWITZ
                                                              83
                                                                      RVUE \tilde{g}u\overline{d}, \tilde{g}uud
                                   <sup>435</sup> KANE
>2-3
                                                                      RVUE Beam dump
>1.5-2
                                         FARRAR
                                                              78
                                                                      RVUE R-hadron
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- ³⁹⁴BERGER 05 include the light gluino in proton PDF and perform global analysis of hadronic data. Effects on the running of α_s also included. Strong dependency on $\alpha_s(m_Z)$. Bound quoted for $\alpha_s(m_Z) = 0.118$.
- 395 ABDALLAH 03C looked for events of the type $q\overline{q}R^{\pm}R^{\pm}$, $q\overline{q}R^{\pm}R^{0}$ or $q\overline{q}R^{0}R^{0}$ in $e^{+}e^{-}$ interactions at 91.2 GeV collected in 1994. The R^{\pm} bound states are identified by anomalous dE/dx in the tracking chambers and the R^{0} by missing energy, due to their reduced energy loss in the calorimeters. The upper value of the excluded range depends on the probability for the gluino to fragment into R^{\pm} or R^{0} , see their Fig. 17. It improves to 23 GeV for 100% fragmentation to R^{\pm} .
- 396 ABDALLAH 03G used $e^+\,e^-$ data at and around the Z^0 peak, above the Z^0 up to $\sqrt{s}=202$ GeV and events from radiative return to cover the low energy region. They perform a direct measurement of the QCD beta-function from the means of fully inclusive event observables. Compared to the energy range, gluinos below 5 GeV can be considered massless and are firmly excluded by the measurement.
- 397 HEISTER 03 use $e^+\,e^-$ data from 1994 and 1995 at and around the Z^0 peak to measure the 4-jet rate and angular correlations. The comparison with QCD NLO calculations allow $\alpha_S(M_Z)$ and the color factor ratios to be extracted and the results are in agreement with the expectations from QCD. The inclusion of a massless gluino in the beta functions yields $T_R \ / \ C_F = 0.15 \pm 0.06 \pm 0.06$ (expectation is $T_R \ / \ C_F = 3/8$), excluding a massless gluino at more than 95% CL. As no NLO calculations are available for massive gluinos, the earlier LO results from BARATE 97L for massive gluinos remain valid.
- ³⁹⁸ HEISTER 03H use e^+e^- data at and around the Z^0 peak to look for stable gluinos hadronizing into charged or neutral R-hadrons with arbitrary branching ratios. Combining these results with bounds on the Z^0 hadronic width from electroweak measurements (JANOT 03) to cover the low mass region the quoted lower limit on the mass of a long-lived gluino is obtained.
- ³⁹⁹ JANOT 03 excludes a light gluino from the upper limit on an additional contribution to the Z hadronic width. At higher confidence levels, $m_{\widetilde{\varrho}} > 5.3(4.2)$ GeV at $3\sigma(5\sigma)$ level.
- 400 MAFI 00 reanalyzed CDF data assuming a stable heavy gluino as the LSP, with model for *R*-hadron-nucleon scattering. Gluino masses between 35 GeV and 115 GeV are excluded

- based on the CDF Run I data. Combined with the analysis of BAER 99, this allows a LSP gluino mass between 25 and 35 GeV if the probability of fragmentation into charged R-hadron P>1/2. The cosmological exclusion of such a gluino LSP are assumed to be avoided as in BAER 99. Gluino could be NLSP with $au_{\widetilde{m{g}}} \sim 100$ yrs, and decay to gluon
- ⁴⁰¹ ALAVI-HARATI 99E looked for R^0 bound states, yielding $\pi^+\pi^-$ or π^0 in the final state. The experiment is sensitive to values of $\Delta m = m_{R^0} m_{\widetilde{\gamma}}$ larger than 280 MeV and 140 MeV for the two decay modes, respectively, and to R^0 mass and lifetime in the ranges 0.8–5 GeV and 10^{-10} – 10^{-3} s. The limits obtained depend on B($R^0 \rightarrow$ $\pi^+\pi^-$ photino) and B($R^0 o\pi^0$ photino) on the value of $m_{R^0}/m_{\widetilde{\gamma}}$, and on the ratio of production rates $\sigma(R^0)/\sigma(K_I^0)$. See Figures in the paper for the excluded R^0 production rates as a function of Δm , R^0 mass and lifetime. Using the production rates expected from perturbative QCD, and assuming dominance of the above decay channels over the suitable phase space, R^0 masses in the range 0.8–5 GeV are excluded at 90%CL for a large fraction of the sensitive lifetime region. ALAVI-HARATI 99E updates and supersedes the results of ADAMS 97B.
- 402 BAER 99 set constraints on the existence of stable \widetilde{g} hadrons, in the mass range $m_{\widetilde{g}} > 3$ GeV. They argue that strong-interaction effects in the low-energy annihilation rates could leave small enough relic densities to evade cosmological constraints up to $m_{\widetilde{\sigma}} < 10$ TeV. They consider jet+ $\not\!\!E_T$ as well as heavy-ionizing charged-particle signatures from production of stable \widetilde{g} hadrons at LEP and Tevatron, developing modes for the energy loss of \widetilde{g} hadrons inside the detectors. Results are obtained as a function of the fragmentation probability P of the \tilde{g} into a charged hadron. For P < 1/2, and for various energyloss models, OPAL and CDF data exclude gluinos in the 3 $< m_{\widetilde{g}}({
 m GeV}) <$ 130 mass range. For P>1/2, gluinos are excluded in the mass ranges $3< m_{\widetilde{g}}({\rm GeV})<23$ and $50 < m_{\widetilde{\varrho}}(\text{GeV}) < 200.$
- ⁴⁰³ FANTI 99 looked for R^0 bound states yielding high P_T $\eta \to 3\pi^0$ decays. The experiment is sensitive to a region of R^0 mass and lifetime in the ranges of 1–5 GeV and 10^{-10} – 10^{-3} s. The limits obtained depend on B($R^0 \to \eta \widetilde{\gamma}$), on the value of $m_{R^0}/m_{\widetilde{\gamma}}$, and on the ratio of production rates $\sigma(R^0)/\sigma(K_L^0)$. See Fig. 6–7 for the excluded production rates as a function of R^0 mass and lifetime.
- $404\,\text{ACKERSTAFF}$ 98V excludes the light gluino with universal gaugino mass where charginos,
- neutralinos decay as $\widetilde{\chi}_1^{\pm}, \widetilde{\chi}_2^0 \to q \overline{q} \widetilde{g}$ from total hadronic cross sections at \sqrt{s} =130–172 GeV. See paper for the case of nonuniversal gaugino mass. 405 ADAMS 97B looked for $\rho^0 \to \pi^+\pi^-$ as a signature of R^0 =($\widetilde{g} g$) bound states. The experiment is sensitive to an R^0 mass range of 1.2–4.5 GeV and to a lifetime range of 1.2–10 to 3 10^{-10} – 10^{-3} sec. Precise limits depend on the assumed value of $m_{R^0}/m_{\widetilde{\gamma}}$. See Fig. 7 for the excluded mass and lifetime region.
- 406 ALBUQUERQUE 97 looked for weakly decaying baryon-like states which contain a light gluino, following the suggestions in FARRAR 96. See their Table 1 for limits on the production fraction. These limits exclude gluino masses in the range 100-600 MeV for the predicted lifetimes (FARRAR 96) and production rates, which are assumed to be comparable to those of strange or charmed baryons.
- $^{
 m 407}$ BARATE 97L studied the QCD color factors from four-jet angular correlations and the differential two-jet rate in Z decay. Limit obtained from the determination of $n_f =$ 4.24 \pm 0.29 \pm 1.15, assuming T_F/C_F =3/8 and C_A/C_F =9/4.
- ⁴⁰⁸ CSIKOR 97 combined the α_s from $\sigma(e^+e^- \rightarrow hadron)$, τ decay, and jet analysis in Z decay. They exclude a light gluino below 5 GeV at more than 99.7%CL.
- $^{
 m 409}$ DEGOUVEA 97 reanalyzed AKERS 95A data on Z decay into four jets to place constraints on a light stable gluino. The mass limit corresponds to the pole mass of 2.8 GeV. The analysis, however, is limited to the leading-order QCD calculation.

- ⁴¹⁰ FARRAR 96 studied the possible $R^0 = (\tilde{g} \, g)$ component in Fermilab E799 experiment and used its bound B($K_L^0 \to \pi^0 \, \nu \, \overline{\nu}$) $\leq 5.8 \times 10^{-5}$ to place constraints on the combination of R^0 production cross section and its lifetime.
- ⁴¹¹ AKERS 95R looked for Z decay into $q\overline{q}\widetilde{g}\widetilde{g}$, by searching for charged particles with dE/dx consistent with \widetilde{g} fragmentation into a state $(\widetilde{g}q\overline{q})^{\pm}$ with lifetime $\tau>10^{-7}$ sec. The fragmentation probability into a charged state is assumed to be 25%.
- ⁴¹² CLAVELLI 95 updates the analysis of CLAVELLI 93, based on a comparison of the hadronic widths of charmonium and bottomonium S-wave states. The analysis includes a parametrization of relativistic corrections. Claims that the presence of a light gluino improves agreement with the data by slowing down the running of α_s .
- ⁴¹³CAKIR 94 reanalyzed TUTS 87 and later unpublished data from CUSB to exclude pseudo-scalar gluinonium $\eta_{\widetilde{g}}(\widetilde{g}\,\widetilde{g})$ of mass below 7 GeV. it was argued, however, that the perturbative QCD calculation of the branching fraction $\Upsilon \to \eta_{\widetilde{g}} \gamma$ is unreliable for $m_{\eta_{\widetilde{g}}} < 3$ GeV. The gluino mass is defined by $m_{\widetilde{g}} = (m_{\eta_{\widetilde{q}}})/2$. The limit holds for any gluino lifetime.
- 414 LOPEZ 93C uses combined restraint from the radiative symmetry breaking scenario within the minimal supergravity model, and the LEP bounds on the (M_2,μ) plane. Claims that the light gluino window is strongly disfavored.
- 415 CLAVELLI 92 claims that a light gluino mass around 4 GeV should exist to explain the discrepancy between $\alpha_{\rm S}$ at LEP and at quarkonia (Υ), since a light gluino slows the running of the QCD coupling.
- 416 ANTONIADIS 91 argue that possible light gluinos (< 5 GeV) contradict the observed running of $\alpha_{\rm S}$ between 5 GeV and m_{Z} . The significance is less than 2 s.d.
- 417 ANTONIADIS 91 interpret the search for missing energy events in 450 GeV/ $^{\prime}c$ $^{\prime}p$ $^{\prime}N$ collisions, AKESSON 91, in terms of light gluinos.
- 418 NAKAMURA 89 searched for a long-lived ($\tau \gtrsim 10^{-7}$ s) charge-(±2) particle with mass $\lesssim 1.6$ GeV in proton-Pt interactions at 12 GeV and found that the yield is less than 10^{-8} times that of the pion. This excludes $R\text{-}\Delta^{++}$ (a $\tilde{g}\,u\,u\,u$ state) lighter than 1.6 GeV.
- 419 The limits assume $m_{\widetilde{q}}=$ 100 GeV. See their figure 3 for limits vs. $m_{\widetilde{q}}.$
- 420 The gluino mass is defined by half the bound $\widetilde{g}\widetilde{g}$ mass. If zero gluino mass gives a $\widetilde{g}\widetilde{g}$ of mass about 1 GeV as suggested by various glueball mass estimates, then the low-mass bound can be replaced by zero. The high-mass bound is obtained by comparing the data with nonrelativistic potential-model estimates.
- 421 ALBRECHT 86C search for secondary decay vertices from $\chi_{b1}(1P) \to \widetilde{g}\,\widetilde{g}\,g$ where \widetilde{g} 's make long-lived hadrons. See their figure 4 for excluded region in the $m_{\widetilde{g}}-m_{\widetilde{g}}$ and $m_{\widetilde{g}}-m_{\widetilde{q}}$ plane. The lower $m_{\widetilde{g}}$ region below ~ 2 GeV may be sensitive to fragmentation effects. Remark that the \widetilde{g} -hadron mass is expected to be ~ 1 GeV (glueball mass) in the zero \widetilde{g} mass limit.
- ⁴²² BADIER 86 looked for secondary decay vertices from long-lived \widetilde{g} -hadrons produced at 300 GeV π^- beam dump. The quoted bound assumes \widetilde{g} -hadron nucleon total cross section of 10μ b. See their figure 7 for excluded region in the $m_{\widetilde{g}}-m_{\widetilde{q}}$ plane for several assumed total cross-section values.
- 423 BARNETT 86 rule out light gluinos (m=3-5 GeV) by calculating the monojet rate from gluino gluino gluon events (and from gluino gluino events) and by using UA1 data from $p\bar{p}$ collisions at CERN.
- ⁴²⁴ VOLOSHIN 86 rules out stable gluino based on the cosmological argument that predicts too much hydrogen consisting of the charged stable hadron \tilde{g} uud. Quasi-stable ($\tau > 1. \times 10^{-7}$ s) light gluino of $m_{\widetilde{g}} < 3$ GeV is also ruled out by nonobservation of the stable charged particles, \tilde{g} uud, in high energy hadron collisions.
- ⁴²⁵COOPER-SARKAR 85B is BEBC beam-dump. Gluinos decaying in dump would yield $\widetilde{\gamma}$'s in the detector giving neutral-current-like interactions. For $m_{\widetilde{q}} >$ 330 GeV, no limit is set.

- 426 DAWSON 85 first limit from neutral particle search. Second limit based on FNAL beam dump experiment.
- 427 FARRAR 85 points out that BALL 84 analysis applies only if the \widetilde{g} 's decay before interacting, i.e. $m_{\widetilde{q}}~<\!80 m_{\widetilde{g}}^{-1.5}$. FARRAR 85 finds $m_{\widetilde{g}}~<\!0.5$ not excluded for $m_{\widetilde{q}}=30\text{--}1000$ GeV and $m_{\widetilde{g}}~<\!1.0$ not excluded for $m_{\widetilde{q}}=100\text{--}500$ GeV by BALL 84 experiment.
- ⁴²⁸ GOLDMAN 85 use nonobservation of a pseudoscalar \widetilde{g} - \widetilde{g} bound state in radiative ψ decay.
- 429 HABER 85 is based on survey of all previous searches sensitive to low mass \tilde{g} 's. Limit makes assumptions regarding the lifetime and electric charge of the lightest supersymmetric particle.
- 430 BALL 84 is FNAL beam dump experiment. Observed no interactions of $\widetilde{\gamma}$ in the calorimeter, where $\widetilde{\gamma}$'s are expected to come from pair-produced \widetilde{g} 's. Search for long-lived $\widetilde{\gamma}$ interacting in calorimeter 56m from target. Limit is for $m_{\widetilde{q}}=40$ GeV and production cross section proportional to A^{0.72}. BALL 84 find no \widetilde{g} allowed below 4.1 GeV at CL = 90%. Their figure 1 shows dependence on $m_{\widetilde{q}}$ and A. See also KANE 82.
- ⁴³¹ BRICK 84 reanalyzed FNAL 147 GeV HBC data for R- Δ (1232)⁺⁺ with $\tau > 10^{-9}$ s and $p_{\text{lab}} > 2$ GeV. Set CL = 90% upper limits 6.1, 4.4, and 29 microbarns in pp, π^+p , K^+p collisions respectively. R- Δ^{++} is defined as being \widetilde{g} and 3 up quarks. If mass = 1.2–1.5 GeV, then limits may be lower than theory predictions.
- 432 FARRAR 84 argues that $m_{\widetilde{g}}~<\!100$ MeV is not ruled out if the lightest R-hadrons are long-lived. A long lifetime would occur if R-hadrons are lighter than $\widetilde{\gamma}$'s or if $m_{\widetilde{q}}~>\!100$ GeV.
- 433 BERGSMA 83C is reanalysis of CERN-SPS beam-dump data. See their figure 1.
- 434 CHANOWITZ 83 find in bag-model that charged s-hadron exists which is stable against strong decay if $m_{\widetilde{g}}$ <1 GeV. This is important since tracks from decay of neutral s-hadron cannot be reconstructed to primary vertex because of missed $\widetilde{\gamma}$. Charged s-hadron leaves track from vertex.
- 435 KANE 82 inferred above \tilde{g} mass limit from retroactive analysis of hadronic collision and beam dump experiments. Limits valid if \tilde{g} decays inside detector.

LIGHT \widetilde{G} (Gravitino) MASS LIMITS FROM COLLIDER EXPERIMENTS

The following are bounds on light ($\ll 1\,\mathrm{eV}$) gravitino indirectly inferred from its coupling to matter suppressed by the gravitino decay constant.

Unless otherwise stated, all limits assume that other supersymmetric particles besides the gravitino are too heavy to be produced. The gravitino is assumed to be undetected and to give rise to a missing energy (\cancel{E}) signature.

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
\bullet \bullet We do not	use the	following data for aver	ages, fits, li	mits, etc. • • •
$> 1.09 \times 10^{-5}$	95	436 ABDALLAH 09	5B DLPH	$e^+e^- ightarrow \ \widetilde{G} \ \widetilde{G} \gamma$
$> 1.35 \times 10^{-5}$	95		4E L3	$e^+e^- ightarrow \widetilde{G} \widetilde{G} \gamma$
$> 1.3 \times 10^{-5}$			3c ALEP	$e^+e^- ightarrow \widetilde{G} \widetilde{G} \gamma$
$>11.7 \times 10^{-6}$	95		2H CDF	
$> 8.7 \times 10^{-6}$	95		OD OPAL	$e^+e^- ightarrow\ \widetilde{G}\widetilde{G}\gamma$
$>10.0 \times 10^{-6}$	95	441 ABREU 00	0z DLPH	Superseded by ABDAL- LAH 05B
$>11 \times 10^{-6}$	95	442 AFFOLDER 00	0」 CDF	$p \overline{p} \rightarrow \widetilde{G} \widetilde{G} + \text{jet}$
$> 8.9 \times 10^{-6}$	95	441 ACCIARRI 99	9R L3	Superseded by ACHARD 04E
$> 7.9 \times 10^{-6}$	95		8V L3	$e^+e^- ightarrow\ \widetilde{G}\widetilde{G}\gamma$
$> 8.3 \times 10^{-6}$	95	443 BARATE 98	8J ALEP	$e^+e^- \rightarrow \widetilde{G}\widetilde{G}\gamma$

- 436 ABDALLAH 05B use data from $\sqrt{s}=180$ –208 GeV. They look for events with a single photon + \cancel{E} final states from which a cross section limit of $\sigma < 0.18~pb$ at 208 GeV is obtained, allowing a limit on the mass to be set. Supersedes the results of ABREU 00Z.
- 437 ACHARD 04E use data from $\sqrt{s}=$ 189–209 GeV. They look for events with a single photon $+ \not\!\! E$ final states from which a limit on the Gravitino mass is set corresponding to $\sqrt{F}~>$ 238 GeV. Supersedes the results of ACCIARRI 99R.
- 438 HEISTER 03C use the data from $\sqrt{s}=$ 189–209 GeV to search for γE_T final states.
- 439 ACOSTA 02H looked in 87 pb^{-1} of $p\overline{p}$ collisions at \sqrt{s} =1.8 TeV for events with a high- E_T photon and E_T . They compared the data with a GMSB model where the final state could arise from $q\overline{q} \to \widetilde{G}\widetilde{G}\gamma$. Since the cross section for this process scales as $1/|F|^4$, a limit at 95% CL is derived on $|F|^{1/2} >$ 221 GeV. A model independent limit for the above topology is also given in the paper.
- 440 ABBIENDI,G 00D searches for $\gamma E\!\!\!\!/$ final states from $\sqrt{s} = 189$ GeV.
- ⁴⁴¹ ABREU 00Z, ACCIARRI 99R search for γE final states using data from \sqrt{s} =189 GeV.
- ⁴⁴² AFFOLDER 00J searches for final states with an energetic jet (from quark or gluon) and large E_T from undetected gravitinos.
- 443 Searches for γE final states at \sqrt{s} =183 GeV.

Supersymmetry Miscellaneous Results

Results that do not appear under other headings or that make nonminimal assumptions.

VALUE DOCUMENT ID TECN COMMENT • • • We do not use the following data for averages, fits, limits, etc. • • • ⁴⁴⁴ ABULENCIA 06P CDF $\ell \gamma E_T$, $\ell \ell \gamma$, GMSB ⁴⁴⁵ ACOSTA 04E CDF 04 ISTR $K^{-} \to \pi^{-} \pi^{0} P$ 446 TCHIKILEV ⁴⁴⁷ AFFOLDER 02D CDF $p\overline{p}
ightarrow \gamma b (E_T)$ ⁴⁴⁸ AFFOLDER 01H CDF $p\overline{p} \rightarrow \gamma \gamma X$ ⁴⁴⁹ ABBOTT 00G D0 $p\overline{p} \rightarrow 3\ell + E_T$, R, $LL\overline{E}$ 00C DLPH $e^+e^- \rightarrow \gamma + S/P$ 450 ABREU.P ⁴⁵¹ ABACHI ⁴⁵² BARBER 84B RVUE ⁴⁵³ HOFFMAN CNTR $\pi p \rightarrow n(e^+e^-)$ 83

- 444 ABULENCIA 06P searched in 305 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for an excess of events with $\ell\gamma \not\!\!E_T$ and $\ell\ell\gamma$ ($\ell=e,\,\mu$). No significant excess was found compared to the background expectation. No events are found such as the $e\,e\,\gamma\gamma\not\!\!E_T$ event observed in ABF 991.
- in ABE 991. 445 ACOSTA 04E looked in 107 pb^{-1} of $p\overline{p}$ collisions at $\sqrt{s}=1.8$ TeV for events with two same sign leptons without selection of other objects nor E_T . No significant excess is observed compared to the Standard Model expectation and constraints are derived on the parameter space of MSUGRA models, see Figure 4.
- ⁴⁴⁶ Looked for the scalar partner of a goldstino in decays $K^- \to \pi^- \pi^0 P$ from a 25 GeV K^- beam produced at the IHEP 70 GeV proton synchrotron. The sgoldstino is assumed to be sufficiently long-lived to be invisible. A 90% CL upper limit on the decay branching ratio is set at $\sim 9.0 \times 10^{-6}$ for a sgoldstino mass range from 0 to 200 MeV, excluding the interval near $m(\pi^0)$, where the limit is $\sim 3.5 \times 10^{-5}$.
- 447 AFFOLDER 02D looked in 85 pb $^{-1}$ of $p\overline{p}$ collisions at \sqrt{s} =1.8 TeV for events with a high- E_T photon, and a b-tagged jet with or without $\not\!\!E_T$. They compared the data with models where the final state could arise from cascade decays of gluinos and/or squarks into $\widetilde{\chi}^\pm$ and $\widetilde{\chi}^0_2$ or direct associated production of $\widetilde{\chi}^0_2\widetilde{\chi}^\pm_2$, followed by $\widetilde{\chi}^0_2\to \gamma\widetilde{\chi}^0_1$ or a GMSB model where $\widetilde{\chi}^0_1\to \gamma\widetilde{G}$. It is concluded that the experimental sensitivity is

- insufficient to detect the associated production or the GMSB model, but some sensitivity may exist to the cascade decays. A model independent limit for the above topology is also given in the paper.
- ⁴⁴⁸ AFFOLDER 01H searches for $p\overline{p} \to \gamma\gamma X$ events, where the di-photon system originates from sgoldstino production, in 100 pb $^{-1}$ of data. Upper limits on the cross section times branching ratio are shown as function of the di-photon mass >70 GeV in Fig. 5. Excluded regions are derived in the plane of the sgoldstino mass versus the supersymmetry breaking scale for two representative sets of parameter values, as shown in Figs. 6 and 7.
- 449 ABBOTT 00G searches for trilepton final states ($\ell = e, \mu$) with $\not\!E_T$ from the indirect decay of gauginos via $LL\overline{E}$ couplings. Efficiencies are computed for all possible production and decay modes of SUSY particles in the framework of the Minimal Supergravity scenario. See Figs. 1–4 for excluded regions in the $m_{1/2}$ versus m_0 plane.
- ⁴⁵⁰ ABREU,P 00C look for the *CP*-even (*S*) and *CP*-odd (*P*) scalar partners of the goldstino, expected to be produced in association with a photon. The S/P decay into two photons or into two gluons and both the tri-photon and the photon + two jets topologies are investigated. Upper limits on the production cross section are shown in Fig. 5 and the excluded regions in Fig. 6. Data collected at \sqrt{s} = 189–202 GeV.
- ⁴⁵¹ ABACHI 97 searched for $p\overline{p} \to \gamma \gamma \not \!\!\! E_T + X$ as supersymmetry signature. It can be caused by selectron, sneutrino, or neutralino production with a radiative decay of their decay products. They placed limits on cross sections.
- 452 BARBER 84B consider that $\widetilde{\mu}$ and \widetilde{e} may mix leading to $\mu \to e \widetilde{\gamma} \widetilde{\gamma}.$ They discuss mass-mixing limits from decay dist. asym. in LBL-TRIUMF data and e^+ polarization in SIN data.
- ⁴⁵³ HOFFMAN 83 set CL = 90% limit $d\sigma/dt$ B(e^+e^-) < 3.5 × 10⁻³² cm²/GeV² for spin-1 partner of Goldstone fermions with 140 < m <160 MeV decaying $\rightarrow e^+e^-$ pair.

REFERENCES FOR Supersymmetric Particle Searches

ABAZOV	07B	PL B645 119	V.M. Abazov <i>et al.</i>	`	Collab.)
ABAZOV	06C	PL B638 119	V.M. Abazov <i>et al.</i>		Collab.)
ABAZOV	06D	PL B638 441	V.M. Abazov <i>et al.</i>		Collab.)
ABAZOV	06I	PRL 97 111801	V.M. Abazov <i>et al.</i>	(D0	Collab.)
ABAZOV	06P	PRL 97 161802	V.M. Abazov et al.	(D0	Collab.)
ABAZOV	06R	PRL 97 171806	V.M. Abazov et al.	(D0	Collab.)
ABBIENDI	06B	EPJ C46 307	G. Abbiendi <i>et al.</i>		Collab.)
ABDALLAH	06C	EPJ C45 589	J. Abdallah <i>et al.</i>	(DELPHI	Collab.)
ABULENCIA	061	PRL 96 171802	A. Abulencia et al.		Collab.)
ABULENCIA	06M	PRL 96 211802	A. Abulencia et al.		Collab.)
ABULENCIA	06P	PRL 97 031801	A. Abulencia et al.	(CDF	Collab.)
ACHTERBERG	06	ASP 26 129	A. Achterberg et al.	(AMANDA	
ACKERMANN	06	ASP 24 459	M. Ackermann et al.	(AMANDA	
AKERIB	06	PR D73 011102R	D.S. Akerib et al.		Collab.)
AKERIB	06A	PRL 96 011302	D.S. Akerib et al.	`	Collab.)
BENOIT	06	PL B637 156	A. Benoit <i>et al.</i>	,	,
DEBOER	06	PL B636 13	W. de Boer <i>et al.</i>		
LEE	06	PL B633 201	H.S. Lee et al.	(KIMS	Collab.)
SHIMIZU	06A	PL B633 195	Y. Shimizu et al.	(,
SMITH	06	PL B642 567	N.J.T. Smith, A.S. Murphy,	T.J. Summer	
ABAZOV	05A	PRL 94 041801	V.M. Abazov et al.		Collab.)
ABAZOV	05U	PRL 95 151805	V.M. Abazov et al.	`	Collab.)
ABDALLAH	05B	EPJ C38 395	J. Abdallah <i>et al.</i>	(DELPHI	Collab.)
ABULENCIA	05A	PRL 95 252001	A. Abulencia et al.		Collab.)
ACOSTA	05E	PR D71 031104R	D. Acosta et al.	`	Collab.)
ACOSTA	05R	PRL 95 131801	D. Acosta et al.	`	Collab.)
AKERIB	05	PR D72 052009	D.S. Akerib et al.		Collab.)
AKTAS	05	PL B616 31	A. Aktas <i>et al.</i>		Collab.)
ALNER	05	PL B616 17	G.J. Alner et al.	(UK Dark Matter	
ALNER	05A	ASP 23 444	G.J. Alner et al.	UK Dark Matter	,
ANGLOHER	05	ASP 23 325	G. Angloher et al.	` (CRESST-II	
BARNABE-HE	.05	PL B624 186	M. Barnabe-Heider <i>et al.</i>	`(PICASSO	
BERGER	05	PR D71 014007	E.L. Berger et al.	,	,
CHEKANOV	05A	EPJ C44 463	S. Chekanov et al.	(ZEUS	Collab.)
				,	,

ELLIS	05	PR D71 095007	J. Ellis et al.	
GIRARD	05	PL B621 233	T.A. Girard et al.	(SIMPLE Collab.)
KLAPDOR-K		PL B609 226	H.V. Klapdor-Kleingrothaus, I	.V. Krivosheina, C. Tomei
SANGLARD ABAZOV	05 04	PR D71 122002 PL B581 147	V. Sanglard <i>et al.</i> V.M. Abazov <i>et al.</i>	(EDELWEISS Collab.)
ABAZOV	04B	PRL 93 011801	V.M. Abazov <i>et al.</i>	(D0 Collab.) (D0 Collab.)
ABBIENDI	04	EPJ C32 453	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	04F	EPJ C33 149	G. Abbiendi et al.	(OPAL Collab.)
ABBIENDI	04H	EPJ C35 1	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	04N	PL B602 167	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABDALLAH ABDALLAH	04H 04M	EPJ C34 145 EPJ C36 1	J. Abdallah <i>et al.</i> J. Abdallah <i>et al.</i>	(DELPHI Collab.) (DELPHI Collab.)
Also	OTIVI	EPJ C37 129 (erratum)	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ACHARD	04	PL B580 37 `	P. Achard et al.	` (L3 Collab.)
ACHARD	04E	PL B587 16	P. Achard et al.	(L3)
ACOSTA	04B	PRL 92 051803	D. Acosta <i>et al.</i>	(CDF Collab.)
ACOSTA AKERIB	04E 04	PRL 93 061802 PRL 93 211301	D. Acosta <i>et al.</i> D. Akerib <i>et al.</i>	(CDF Collab.) (CDMSII Collab.)
AKTAS	04B	PL B599 159	A. Aktas <i>et al.</i>	(H1 Collab.)
AKTAS	04D	EPJ C36 425	A. Aktas <i>et al.</i>	(H1 Collab.)
BELANGER	04	JHEP 0403 012	G. Belanger et al.	,
BOTTINO	04	PR D69 037302	A. Bottino <i>et al.</i>	
DAS DESAI	04 04	PL B596 293	S.P. Das, A. Datta, M. Maity S. Desai <i>et al.</i>	y (Super-Kamiokande Collab.)
ELLIS	04	PR D70 083523 PR D69 015005	J. Ellis <i>et al.</i>	(Super-Namilokande Collab.)
ELLIS	04B	PR D70 055005	J. Ellis <i>et al.</i>	
GIULIANI	04	PL B588 151	F. Giuliani, T.A. Girard	
HEISTER	04	PL B583 247	A. Heister <i>et al.</i>	(ALEPH Collab.)
JANOT	04	PL B594 23	P. Janot	
PIERCE TCHIKILEV	04A 04	PR D70 075006 PL B602 149	A. Pierce O.G. Tchikilev <i>et al.</i>	(ISTRA+ Coolab.)
ABBIENDI	03H	EPJ C29 479	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	03L	PL B572 8	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABDALLAH	03C	EPJ C26 505	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABDALLAH	03D	EPJ C27 153	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABDALLAH	03F 03G	EPJ C28 15 EPJ C29 285	J. Abdallah <i>et al.</i> J. Abdallah <i>et al.</i>	(DELPHI Collab.)
ABDALLAH ABDALLAH	03M	EPJ C31 421	J. Abdallah <i>et al.</i>	(DELPHI Collab.) (DELPHI Collab.)
ACOSTA	03C	PRL 90 251801	D. Acosta <i>et al.</i>	(CDF Collab.)
ACOSTA	03E	PRL 91 171602	D. Acosta et al.	(CDF Collab.)
ADLOFF	03	PL B568 35	C. Adloff et al.	(H1 Collab.)
AHMED AKERIB	03 03	ASP 19 691 PR D68 082002	B. Ahmed <i>et al.</i> D. Akerib <i>et al.</i>	(UK Dark Matter Collab.)
BAER	03	JCAP 0305 006	H. Baer, C. Balazs	(CDMS Collab.)
BAER	03A	JCAP 0309 007	H. Baer <i>et al.</i>	
BERGER	03	PL B552 223	E. Berger et al.	
BOTTINO	03	PR D68 043506	A. Bottino et al.	
BOTTINO CHAKRAB	03A 03	PR D67 063519 PR D68 015005	A. Bottino, N. Fornengo, S. S. Chakrabarti, M. Guchait, N.	•
CHATTOPAD		PR D68 035005	U. Chattopadhyay, A. Corsett	
CHEKANOV	03B	PR D68 052004	S. Chekanov et al.	(ZEUS Collab.)
ELLIS	03	ASP 18 395	J. Ellis, K.A. Olive, Y. Santo	
ELLIS	03B	NP B652 259	J. Ellis <i>et al.</i>	
ELLIS ELLIS	03C 03D	PL B565 176 PL B573 162	J. Ellis <i>et al.</i> J. Ellis <i>et al.</i>	
ELLIS	03E	PR D67 123502	J. Ellis <i>et al.</i>	
HEISTER	03	EPJ C27 1	A. Heister et al.	(ALEPH)
HEISTER	03C	EPJ C28 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
HEISTER	03G	EPJ C31 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
HEISTER HOOPER	03H 03	EPJ C31 327 PL B562 18	A. Heister <i>et al.</i> D. Hooper, T. Plehn	(ALEPH Collab.)
JANOT	03	PL B564 183	P. Janot	
KLAPDOR-K		ASP 18 525	H.V. Klapdor-Kleingrothaus et	t al.
LAHANAS	03	PL B568 55	A. Lahanas, D. Nanopoulos	
LEP ALEDH DI	03 ELDHI	SLAC-R-701, LEPEWWG		(LEP Collabs.)
TAKEDA	03 -	L3, OPAL, the LEP EWN PL B572 145	A. Takeda <i>et al.</i>	
ABAZOV	02C	PRL 88 171802	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABAZOV	02F	PRL 89 171801	V.M. Abazov et al.	(D0 Collab.)
ABAZOV	02G	PR D66 112001	V.M. Abazov et al.	(D0 Collab.)
ABAZOV	02H	PRL 89 261801	V.M. Abazov et al.	(D0 Collab.)

ABBIENDI	02	EPJ C23 1	G. Abbiendi <i>et al.</i>	(OPAI	Collab.)
ABBIENDI	02B	PL B526 233	G. Abbiendi <i>et al.</i>	` .	Collab.)
ABBIENDI	02H	PL B545 272	G. Abbiendi <i>et al.</i>	` .	Collab.)
Also		PL B548 258 (erratum)	G. Abbiendi <i>et al.</i>	` .	Collab.)
ABRAMS	02	PR D66 122003	D. Abrams <i>et al.</i>	(CDMS	
ACHARD	02	PL B524 65	P. Achard et al.	` .	Collab.)
ACOSTA	02H	PRL 89 281801	D. Acosta et al.	`	Collab.)
AFFOLDER	02	PRL 88 041801	T. Affolder et al.	` ·	Collab.)
AFFOLDER	02D	PR D65 052006	T. Affolder et al.		Collab.)
ANGLOHER	02	ASP 18 43	G. Angloher et al.	(CRÈSST	
ARNOWITT	02	hep-ph/0211417	R. Arnowitt, B. Dutta	,	,
BAEK	02	PL B541 161	S. Baek		
BAER	02	JHEP 0207 050	H. Baer et al.		
BECHER	02	PL B540 278	T. Becher et al.		
BENOIT	02	PL B545 43	A. Benoit et al.	(EDELWEISS	
CHEKANOV	02	PR D65 092004	S. Chekanov <i>et al.</i>	(ZEUS	Collab.)
CHEUNG	02B	PRL 89 221801	K. Cheung, WY. Keung		
CHO	02	PRL 89 091801	GC. Cho		
ELLIS	02	PL B525 308	J. Ellis, D.V. Nanopoulos, K		
ELLIS	02B	PL B532 318	J. Ellis, A. Ferstl, K.A. Olive	e	
ELLIS	02C	PL B539 107	J. Ellis, K.A. Olive, Y. Santo	OSO	
GHODBANE	02	NP B647 190	N. Ghodbane <i>et al.</i>		
HEISTER	02	PL B526 191	A. Heister <i>et al.</i>	(ALEPH	,
HEISTER	02E	PL B526 206	A. Heister <i>et al.</i>	(ALEPH	
HEISTER	02F	EPJ C25 1	A. Heister <i>et al.</i>	(ALEPH	- :
HEISTER	02J	PL B533 223	A. Heister <i>et al.</i>	(ALEPH	- :
HEISTER	02K	PL B537 5	A. Heister <i>et al.</i>	(ALEPH	- :
HEISTER	02N	PL B544 73	A. Heister <i>et al.</i>	(ALEPH	
HEISTER	02R	EPJ C25 339	A. Heister <i>et al.</i>	(ALEPH	Collab.)
KIM	02	PL B527 18	H.B. Kim et al.		
KIM	02B	JHEP 0212 034	Y.G. Kim et al.		
LAHANAS	02	EPJ C23 185	A. Lahanas, V.C. Spanos	(2001-	
MORALES	02B	ASP 16 325	A. Morales et al.	(COSME	
MORALES	02C	PL B532 8	A. Morales <i>et al.</i>	. ``.	Collab.)
ABBIENDI	01	PL B501 12	G. Abbiendi <i>et al.</i>	` .	Collab.)
ABBOTT	01D	PR D63 091102	B. Abbott <i>et al.</i>		Collab.)
ABREU	01	EPJ C19 29	P. Abreu <i>et al.</i>	(DELPHI	
ABREU	01B	EPJ C19 201	P. Abreu <i>et al.</i>	(DELPHI	
ABREU	01C	PL B502 24	P. Abreu <i>et al.</i>	(DELPHI	
ABREU	01D	PL B500 22	P. Abreu et al.	(DELPHI	- :
ABREU	01G	PL B503 34	P. Abreu <i>et al.</i> M. Acciarri <i>et al.</i>	(DELPHI	- :
ACCIARRI ADAMS	01 01	EPJ C19 397	T. Adams <i>et al.</i>	(L3 (NuTeV	Collab.)
ADLOFF	01 01B	PRL 87 041801 EPJ C20 639	C. Adloff <i>et al.</i>	` .	- :
AFFOLDER	01B	PR D63 091101	T. Affolder <i>et al.</i>	`	Collab.) Collab.)
AFFOLDER	01B	PR D64 092002	T. Affolder <i>et al.</i>	` ·	Collab.)
AFFOLDER	01J	PRL 87 251803	T. Affolder <i>et al.</i>	` ·	Collab.)
BALTZ	013	PRL 86 5004	E. Baltz, P. Gondolo	(CDI	Collab.)
BARATE	01	PL B499 67	R. Barate <i>et al.</i>	(ALEPH	Collab)
BARATE	01B	EPJ C19 415	R. Barate <i>et al.</i>	(ALEPH	Collab.)
BARGER	01C	PL B518 117	V. Barger, C. Kao	(/	Coa.b.)
BAUDIS	01	PR D63 022001	L. Baudis <i>et al.</i>	(Heidelberg-Moscow	Collab.)
BENOIT	01	PL B513 15	A. Benoit <i>et al.</i>	(EDELWEISS	
BERGER	01	PRL 86 4231	E. Berger <i>et al.</i>	(======================================	
BERNABEI	01	PL B509 197	R. Bernabei <i>et al.</i>	(DAMA	Collab.)
BOTTINO	01	PR D63 125003	A. Bottino et al.	`	,
BREITWEG	01	PR D63 052002	J. Breitweg et al.	(ZEUS	Collab.)
CORSETTI	01	PR D64 125010	A. Corsetti, P. Nath	(====	
DJOUADI	01	JHEP 0108 055	A. Djouadi, M. Drees, J.L. I	Kneur	
ELLIS	01B	PL B510 236	J. Ellis <i>et al.</i>		
ELLIS	01C	PR D63 065016	J. Ellis, A. Ferstl, K.A. Olive	9	
GOMEZ	01	PL B512 252	M.E. Gomez, J.D. Vergados		
LAHANAS	01	PL B518 94	A. Lahanas, D.V. Nanopoulo	s, V. Spanos	
ROSZKOWSKI	01	JHEP 0108 024	L. Roszkowski, R. Ruiz de A	ustri, T. Nihei	
SAVINOV	01	PR D63 051101	V. Savinov et al.		Collab.)
ABBIENDI	00	EPJ C12 1	G. Abbiendi et al.	(OPAL	Collab.)
ABBIENDI	00G	EPJ C14 51	G. Abbiendi <i>et al.</i>	(OPAL	Collab.)
ABBIENDI	00H	EPJ C14 187	G. Abbiendi <i>et al.</i>		Collab.)
Also		EPJ C16 707 (erratum)	G. Abbiendi <i>et al.</i>	` .	Collab.)
ABBIENDI	00J	EPJ C12 551	G. Abbiendi <i>et al.</i>		Collab.)
ABBIENDI	00R	EPJ C13 553	G. Abbiendi <i>et al.</i>	(OPAL	Collab.)

ABBIENDI,G	00D	EPJ C18 253	G. Abbiendi <i>et al.</i>	(ODAL Callab.)
				(OPAL Collab.)
ABBOTT	00C	PRL 84 2088	B. Abbott <i>et al.</i>	(D0 Collab.)
ABBOTT	00G	PR D62 071701R	B. Abbott <i>et al.</i>	(D0 Collab.)
ABREU	001	EPJ C13 591	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00J	PL B479 129	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00Q	PL B478 65	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00\$			` '
-		PL B485 45	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00T	PL B485 95	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00U	PL B487 36	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00V	EPJ C16 211	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00W	PL B489 38	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	00Z	EPJ C17 53	P. Abreu <i>et al.</i>	(DELPHI Collab.)
				` '
ABREU,P	00C	PL B494 203	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU,P	00D	PL B496 59	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABUSAIDI	00	PRL 84 5699	R. Abusaidi <i>et al.</i>	(CDMS Collab.)
ACCIARRI	00C	EPJ C16 1	M. Acciarri et al.	` (L3 Collab.)
ACCIARRI	00D	PL B472 420	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	00K	PL B482 31	M. Acciarri et al.	(L3 Collab.)
ACCIARRI	00P	PL B489 81	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCOMANDO	00	NP B585 124	E. Accomando <i>et al.</i>	
AFFOLDER	00D	PRL 84 5704	T. Affolder <i>et al.</i>	(CDF Collab.)
AFFOLDER	00G	PRL 84 5273	T. Affolder et al.	(CDF Collab.)
AFFOLDER	00J	PRL 85 1378	T. Affolder et al.	(CDF Collab.)
AFFOLDER	00K	PRL 85 2056	T. Affolder et al.	(CDF Collab.)
BARATE	00G	EPJ C16 71	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	00H	EPJ C13 29	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	001	EPJ C12 183	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	00P	PL B488 234	R. Barate <i>et al.</i>	(ALEPH Collab.)
	00	PL B480 23	R. Bernabei <i>et al.</i>	\ <u>'</u>
BERNABEI				(DAMA Collab.)
BERNABEI	00C	EPJ C18 283	R. Bernabei <i>et al.</i>	(DAMA Collab.)
BERNABEI	00D	NJP 2 15	R. Bernabei <i>et al.</i>	(DAMA Collab.)
BOEHM	00B	PR D62 035012	C. Boehm, A. Djouadi, M. Drees	;
BREITWEG	00E	EPJ C16 253	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
CHO	00B	NP B574 623	GC. Cho, K. Hagiwara	(=====)
COLLAR	00	PRL 85 3083	J.I. Collar et al.	(SIMDLE Collab.)
				(SIMPLE Collab.)
ELLIS	00	PR D62 075010	J. Ellis et al.	
FENG	00	PR D62 075010 PL B482 388	J. Ellis <i>et al.</i> J.L. Feng, K.T. Matchev, F. Wile	czek
FENG	00	PL B482 388 PR D62 023515	J.L. Feng, K.T. Matchev, F. Wild A. Lahanas, D.V. Nanopoulos, V.	.C. Spanos
FENG LAHANAS LEP	00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016	J.L. Feng, K.T. Matchev, F. Wil A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL	
FENG LAHANAS LEP MAFI	00 00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby	.C. Spanos
FENG LAHANAS LEP MAFI MALTONI	00 00 00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni <i>et al.</i>	.C. Spanos .PHI, L3, OPAL, SLD+)
FENG LAHANAS LEP MAFI MALTONI MORALES	00 00 00 00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al.	.C. Spanos
FENG LAHANAS LEP MAFI MALTONI	00 00 00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni <i>et al.</i>	.C. Spanos .PHI, L3, OPAL, SLD+)
FENG LAHANAS LEP MAFI MALTONI MORALES	00 00 00 00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al.	.C. Spanos .PHI, L3, OPAL, SLD+)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG	00 00 00 00 00 00 00	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al.	.C. Spanos .PHI, L3, OPAL, SLD+) (IGEX Collab.) (UK Dark Matter Col.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI	00 00 00 00 00 00 00 00 99	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al.	.C. Spanos PHI, L3, OPAL, SLD+) (IGEX Collab.) (UK Dark Matter Col.) (OPAL Collab.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI	00 00 00 00 00 00 00 00 99 99F	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al.	.C. Spanos PHI, L3, OPAL, SLD+) (IGEX Collab.) (UK Dark Matter Col.) (OPAL Collab.) (OPAL Collab.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI	00 00 00 00 00 00 00 00 99 99F 99M	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al.	.C. Spanos PHI, L3, OPAL, SLD+) (IGEX Collab.) (UK Dark Matter Col.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI ABBIENDI	00 00 00 00 00 00 00 00 99 99F 99M 99T	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95 EPJ C11 619	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al.	(IGEX Collab.) (UK Dark Matter Col.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.)
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FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABE ABE ABE ABREU ABREU	00 00 00 00 00 00 00 00 99 99F 99M 99T 99 99F 99J 99L 99I 99A 99A	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95 EPJ C11 619 PRL 82 29 PR D60 031101 PRL 83 2896 PRL 83 4476 PRL 83 4937 PR D59 092002 PRL 83 2133 EPJ C11 383 EPJ C6 385	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. B. Abbott et al. F. Abe et al. F. Abe et al. P. Abreu et al. P. Abreu et al. P. Abreu et al.	(IGEX Collab.) (IGEX Collab.) (UK Dark Matter Col.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (CDF Collab.) (CDF Collab.) (DELPHI Collab.)
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FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABE ABE ABREU ABREU ABREU ABREU	00 00 00 00 00 00 00 00 99 99F 99M 99T 99 99F 99J 99L 99I 99A 99C 99F	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95 EPJ C11 619 PRL 82 29 PR D60 031101 PRL 83 2896 PRL 83 4476 PRL 83 4937 PR D59 092002 PRL 83 2133 EPJ C11 383 EPJ C1 385 EPJ C7 595 PL B446 62	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. B. Abbott et al. F. Abe et al. F. Abe et al. P. Abreu et al.	(UK Dark Matter Col.) (UK Dark Matter Col.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (CDF Collab.) (CDF Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENTI ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABE ABE ABE ABREU ABREU ACCIARRI	00 00 00 00 00 00 00 00 99 99F 99M 99F 99F 99J 99B 99B 99B 99C 99F	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95 EPJ C11 619 PRL 82 29 PR D60 031101 PRL 83 2896 PRL 83 4476 PRL 83 4937 PR D59 092002 PRL 83 2133 EPJ C11 383 EPJ C1 385 EPJ C7 595 PL B446 62 PL B456 283	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. B. Abbott et al. P. Abreu et al. M. Acciarri et al.	(UK Dark Matter Col.) (OPAL Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (CDF Collab.) (CDF Collab.) (DELPHI Collab.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABE ABE ABE ABE ABE ABREU ABREU ABREU ACCIARRI ACCIARRI	00 00 00 00 00 00 00 00 99 99F 99M 99T 99 99F 99J 99B 99B 99B 99B 99C 99B	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95 EPJ C11 619 PRL 82 29 PR D60 031101 PRL 83 2896 PRL 83 4476 PRL 83 4937 PR D59 092002 PRL 83 2133 EPJ C11 383 EPJ C1 385 EPJ C7 595 PL B446 62 PL B456 283 PL B459 354	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. B. Abbott et al. P. Abreu et al. M. Acciarri et al. M. Acciarri et al.	(UK Dark Matter Col.) (OPAL Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (CDF Collab.) (CDF Collab.) (CDF Collab.) (DELPHI Collab.) (L3 Collab.)
FENG LAHANAS LEP MAFI MALTONI MORALES PDG SPOONER ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENTI ABBOTT ABBOTT ABBOTT ABBOTT ABBOTT ABE ABE ABE ABREU ABREU ACCIARRI	00 00 00 00 00 00 00 00 99 99F 99M 99F 99F 99J 99B 99B 99B 99C 99F	PL B482 388 PR D62 023515 CERN-EP-2000-016 PR D62 035003 PL B476 107 PL B489 268 EPJ C15 1 PL B473 330 EPJ C6 1 EPJ C8 23 PL B456 95 EPJ C11 619 PRL 82 29 PR D60 031101 PRL 83 2896 PRL 83 4476 PRL 83 4937 PR D59 092002 PRL 83 2133 EPJ C11 383 EPJ C1 385 EPJ C7 595 PL B446 62 PL B456 283	J.L. Feng, K.T. Matchev, F. Wilk A. Lahanas, D.V. Nanopoulos, V. LEP Collabs. (ALEPH, DEL A. Mafi, S. Raby M. Maltoni et al. A. Morales et al. D.E. Groom et al. N.J.C. Spooner et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. B. Abbott et al. P. Abreu et al. M. Acciarri et al.	(UK Dark Matter Col.) (OPAL Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (D0 Collab.) (CDF Collab.) (CDF Collab.) (DELPHI Collab.)
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BELLI BERNABEI FANTI	99C 99 99	NP B563 97 PL B450 448 PL B446 117	P. Belli <i>et al.</i> R. Bernabei <i>et al.</i> V. Fanti <i>et al.</i>	(DAMA Collab.) (DAMA Collab.) (CERN NA48 Collab.)
MALTONI	99B	PL B463 230	M. Maltoni, M.I. Vysotsky	(1
OOTANI ABBOTT	99 98	PL B461 371 PRL 80 442	W. Ootani <i>et al.</i> B. Abbott <i>et al.</i>	(D0 Collab.)
ABBOTT	98C	PRL 80 1591	B. Abbott <i>et al.</i>	(D0 Collab.)
ABBOTT ABBOTT	98E 98J	PRL 80 2051 PRL 81 38	B. Abbott <i>et al.</i> B. Abbott <i>et al.</i>	(D0 Collab.) (D0 Collab.)
ABE	98J	PRL 80 5275	F. Abe <i>et al.</i>	(CDF Collab.)
ABE ABREU	98S 98	PRL 81 4806 EPJ C1 1	F. Abe <i>et al.</i> P. Abreu <i>et al.</i>	(CDF Collab.) (DELPHI Collab.)
ABREU	98P	PL B444 491	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI ACCIARRI	98F 98J	EPJ C4 207 PL B433 163	M. Acciarri <i>et al.</i> M. Acciarri <i>et al.</i>	(L3 Collab.) (L3 Collab.)
ACCIARRI	98V	PL B444 503	M. Acciarri et al.	(L3 Collab.)
ACKERSTAFF	98K 98L	EPJ C4 47	K. Ackerstaff <i>et al.</i> K. Ackerstaff <i>et al.</i>	(OPAL Collab.) (OPAL Collab.)
ACKERSTAFF ACKERSTAFF	96L 98P	EPJ C2 213 PL B433 195	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	98V	EPJ C2 441	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BARATE BARATE	98H 98J	PL B420 127 PL B429 201	R. Barate <i>et al.</i> R. Barate <i>et al.</i>	(ALEPH Collab.) (ALEPH Collab.)
BARATE	98K	PL B433 176	R. Barate et al.	(ALEPH Collab.)
BARATE BARATE	98S 98X	EPJ C4 433 EPJ C2 417	R. Barate <i>et al.</i> R. Barate <i>et al.</i>	(ALEPH Collab.) (ALEPH Collab.)
BERNABEI	98	PL B424 195	R. Bernabei et al.	(DAMA Collab.)
BERNABEI BREITWEG	98C 98	PL B436 379 PL B434 214	R. Bernabei <i>et al.</i> J. Breitweg <i>et al.</i>	(DAMA Collab.) (ZEUS Collab.)
ELLIS	98 98	PR D58 095002	J. Ellis <i>et al.</i>	(ZEO3 Collab.)
ELLIS	98B	PL B444 367	J. Ellis, T. Falk, K. Olive C. Caso <i>et al.</i>	
PDG ABACHI	98 97	EPJ C3 1 PRL 78 2070	S. Abachi <i>et al.</i>	(D0 Collab.)
ABBOTT	97B	PRL 79 4321	B. Abbott et al.	(D0 Collab.)
ABE ACCIARRI	97K 97U	PR D56 R1357 PL B414 373	F. Abe <i>et al.</i> M. Acciarri <i>et al.</i>	(CDF Collab.) (L3 Collab.)
ACKERSTAFF	97H	PL B396 301	K. Ackerstaff et al.	(OPAL Collab.)
ADAMS ALBUQUERQ	97B 97	PRL 79 4083 PRL 78 3252	J. Adams <i>et al.</i> I.F. Albuquerque <i>et al.</i>	(FNAL KTeV Collab.) (FNAL E761 Collab.)
BAER	97	PR D57 567	H. Baer, M. Brhlik	(TIVIL ETGI COMBD.)
BARATE BARATE	97K 97L	PL B405 379 ZPHY C76 1	R. Barate <i>et al.</i> R. Barate <i>et al.</i>	(ALEPH Collab.) (ALEPH Collab.)
BERNABEI	97L 97	ASP 7 73	R. Bernabei <i>et al.</i>	(DAMA Collab.)
CARENA	97 07	PL B390 234	M. Carena, G.F. Giudice, C.E.M	
CSIKOR DATTA	97 97	PRL 78 4335 PL B395 54	F. Csikor, Z. Fodor A. Datta, M. Guchait, N. Parua	(EOTV, CERN) (ICTP, TATA)
DEGOUVEA	97	PL B400 117	A. de Gouvea, H. Murayama	
DERRICK EDSJO	97 97	ZPHY C73 613 PR D56 1879	M. Derrick <i>et al.</i> J. Edsjo, P. Gondolo	(ZEUS Collab.)
ELLIS	97	PL B394 354	J. Ellis, J.L. Lopez, D.V. Nanop	
HEWETT KALINOWSKI	97 97	PR D56 5703 PL B400 112	J.L. Hewett, T.G. Rizzo, M.A. I J. Kalinowski, P. Zerwas	Joncheski
TEREKHOV	97	PL B412 86	I. Terekhov	(ALAT)
ABACHI ABACHI	96 96B	PRL 76 2228 PRL 76 2222	S. Abachi <i>et al.</i> S. Abachi <i>et al.</i>	(D0 Collab.) (D0 Collab.)
ABE	96	PRL 77 438	F. Abe <i>et al.</i>	(CDF Collab.)
ABE ABE	96D 96K	PRL 76 2006 PRL 76 4307	F. Abe <i>et al.</i> F. Abe <i>et al.</i>	(CDF Collab.) (CDF Collab.)
AID	96	ZPHY C71 211	S. Aid <i>et al.</i>	(H1 Collab.)
ADNOWITT	96C	PL B380 461	S. Aid <i>et al.</i>	(H1 Collab.)
ARNOWITT BAER	96 96	PR D54 2374 PR D53 597	R. Arnowitt, P. Nath H. Baer, M. Brhlik	
BERGSTROM	96 06	ASP 5 263	L. Bergstrom, P. Gondolo	(TOKALL OCII)
CHO FARRAR	96 96	PL B372 101 PRL 76 4111	G.C. Cho, Y. Kizukuri, N. Oshir G.R. Farrar	mo (TOKAH, OCH) (RUTG)
LEWIN	96	ASP 6 87	J.D. Lewin, P.F. Smith	,,,,,
TEREKHOV ABACHI	96 95C	PL B385 139 PRL 75 618	I. Terkhov, L. Clavelli S. Abachi <i>et al.</i>	(ALAT) (D0 Collab.)
ABE	95N	PRL 74 3538	F. Abe <i>et al.</i>	(CDF Collab.)
ABE ACCIARRI	95T 95E	PRL 75 613 PL B350 109	F. Abe <i>et al.</i> M. Acciarri <i>et al.</i>	(CDF Collab.) (L3 Collab.)
AKERS	95A	ZPHY C65 367	R. Akers et al.	(OPAL Collab.)

AKERS	95R	ZPHY C67 203	R. Akers <i>et al.</i> (OPAL Collab.)
BEREZINSKY	95_	ASP 5 1	V. Berezinsky et al.
BUSKULIC	95E	PL B349 238	D. Buskulic <i>et al.</i> (ALEPH Collab.)
CLAVELLI FALK	95 95	PR D51 1117 PL B354 99	L. Clavelli, P.W. Coulter (ALAT)
LOSECCO	95 95	PL B342 392	T. Falk, K.A. Olive, M. Srednicki (MINN, UCSB) J.M. LoSecco (NDAM)
AKERS	94K	PL B337 207	R. Akers <i>et al.</i> (OPAL Collab.)
BECK	94	PL B336 141	M. Beck et al. (MPIH, KIAE, SASSO)
CAKIR	94	PR D50 3268	M.B. Cakir, G.R. Farrar (RUTG)
FALK	94	PL B339 248	T. Falk, K.A. Olive, M. Srednicki (UCSB, MINN)
SHIRAI	94	PRL 72 3313	J. Shirai et al. (VENUS Collab.)
ADRIANI ALITTI	93M 93	PRPL 236 1 NP B400 3	O. Adriani et al. (L3 Collab.) J. Alitti et al. (UA2 Collab.)
CLAVELLI	93	PR D47 1973	L. Clavelli, P.W. Coulter, K.J. Yuan (ALAT)
DREES	93	PR D47 376	M. Drees, M.M. Nojiri (DESY, SLAC)
DREES	93B	PR D48 3483	M. Drees, M.M. Nojiri
FALK	93	PL B318 354	T. Falk et al. (UCB, UCSB, MINN)
HEBBEKER	93	ZPHY C60 63	T. Hebbeker (CERN)
KELLEY	93	PR D47 2461	S. Kelley et al. (TAMU, ALAH)
LOPEZ MIZUTA	93C 93	PL B313 241 PL B298 120	J.L. Lopez, D.V. Nanopoulos, X. Wang (TAMU, HARC+) S. Mizuta, M. Yamaguchi (TOHO)
MORI	93	PR D48 5505	M. Mori et al. (KEK, NIIG, TOKY, TOKA+)
ABE	92L	PRL 69 3439	F. Abe <i>et al.</i> (CDF Collab.)
BOTTINO	92	MPL A7 733	A. Bottino <i>et al.</i> (TORI, ZARA)
Also		PL B265 57	A. Bottino et al. (TORI, INFN)
CLAVELLI	92	PR D46 2112	L. Clavelli (ALAT)
DECAMP	92	PRPL 216 253	D. Decamp <i>et al.</i> (ALEPH Collab.)
LOPEZ MCDONALD	92 92	NP B370 445 PL B283 80	J.L. Lopez, D.V. Nanopoulos, K.J. Yuan (TAMU) J. McDonald, K.A. Olive, M. Srednicki (LISB+)
ROY	92	PL B283 270	D.P. Roy (CERN)
ABREU	91F	NP B367 511	P. Abreu <i>et al.</i> (DELPHI Collab.)
AKESSON	91	ZPHY C52 219	T. Akesson <i>et al.</i> (HELIOS Collab.)
ALEXANDER	91F	ZPHY C52 175	G. Alexander <i>et al.</i> (OPAL Collab.)
ANTONIADIS	91	PL B262 109	I. Antoniadis, J. Ellis, D.V. Nanopoulos (EPOL+)
BOTTINO	91 91	PL B265 57	A. Bottino et al. (TORI, INFN)
GELMINI GRIEST	91	NP B351 623 PR D43 3191	G.B. Gelmini, P. Gondolo, E. Roulet (ÚCLA, TRST) K. Griest, D. Seckel
KAMIONKOW.	-	PR D44 3021	M. Kamionkowski (CHIC, FNAL)
MORI	91B	PL B270 89	M. Mori et al. (Kamiokande Collab.)
NOJIRI	91	PL B261 76	M.M. Nojiri (KEK)
OLIVE	91	NP B355 208	K.A. Olive, M. Srednicki (MINN, UCSB)
ROSZKOWSKI		PL B262 59	L. Roszkowski (CERN)
SATO ADACHI	91 90C	PR D44 2220 PL B244 352	N. Sato <i>et al.</i> (Kamiokande Collab.) I. Adachi <i>et al.</i> (TOPAZ Collab.)
GRIEST	90	PR D41 3565	K. Griest, M. Kamionkowski, M.S. Turner (UCB+)
BARBIERI	89C	NP B313 725	R. Barbieri, M. Frigeni, G. Giudice
NAKAMURA	89	PR D39 1261	T.T. Nakamura et al. (KYOT, TMTC)
OLIVE	89	PL B230 78	K.A. Olive, M. Srednicki (MINN, UCSB)
ELLIS	88D	NP B307 883	J. Ellis, R. Flores
GRIEST OLIVE	88B 88	PR D38 2357 PL B205 553	K. Griest K.A. Olive, M. Srednicki (MINN, UCSB)
SREDNICKI	88	NP B310 693	M. Srednicki, R. Watkins, K.A. Olive (MINN, UCSB)
ALBAJAR	87D	PL B198 261	C. Albajar <i>et al.</i> (UA1 Collab.)
ANSARI	87D	PL B195 613	R. Ansari et al. (UA2 Collab.)
ARNOLD	87	PL B186 435	R.G. Arnold et al. (BRUX, DUUC, LOUC+)
NG	87	PL B188 138	K.W. Ng, K.A. Olive, M. Srednicki (MINN, UCSB)
TUTS	87 06.C	PL B186 233	P.M. Tuts <i>et al.</i> (CUSB Collab.)
ALBRECHT BADIER	86C 86	PL 167B 360 ZPHY C31 21	H. Albrecht <i>et al.</i> (ARGUS Collab.) J. Badier <i>et al.</i> (NA3 Collab.)
BARNETT	86	NP B267 625	R.M. Barnett, H.E. Haber, G.L. Kane (LBL, UCSC+)
GAISSER	86	PR D34 2206	T.K. Gaisser, G. Steigman, S. Tilav (BART, DELA)
VOLOSHIN	86	SJNP 43 495	M.B. Voloshin, L.B. Okun (ITEP)
COOPER	050	Translated from YAF 43	
COOPER DAWSON	85B	PL 160B 212	A.M. Cooper-Sarkar <i>et al.</i> (WA66 Collab.)
FARRAR	85 85	PR D31 1581 PRL 55 895	S. Dawson, E. Eichten, C. Quigg (LBL, FNAL) G.R. Farrar (RUTG)
GOLDMAN	85	Physica 15D 181	T. Goldman, H.E. Haber (LANL, UCSC)
HABER	85	PRPL 117 75	H.E. Haber, G.L. Kane (UCSC, MICH)
BALL	84	PRL 53 1314	R.C. Ball et al. (MICH, FIRZ, OSU, FNAL+)
BARBER	84B	PL 139B 427	J.S. Barber, R.E. Shrock (STON)
BRICK	84	PR D30 1134	D.H. Brick et al. (BROW, CAVE, IIT+)

ELLIS	84	NP B238 453	J. Ellis <i>et al.</i>	(CERN)
FARRAR	84	PRL 53 1029	G.R. Farrar	(RUTG)
BERGSMA	83C	PL 121B 429	F. Bergsma et al.	(CHARM Collab.)
CHANOWITZ	83	PL 126B 225	M.S. Chanowitz, S. Sharpe	(UCB, LBL)
GOLDBERG	83	PRL 50 1419	H. Goldberg	(NEAS)
HOFFMAN	83	PR D28 660	C.M. Hoffman et al.	(LANL, ARZS)
KRAUSS	83	NP B227 556	L.M. Krauss	(HARV)
VYSOTSKII	83	SJNP 37 948	M.I. Vysotsky	(ITEP)
		Translated from	YAF 37 1597.	
KANE	82	PL 112B 227	G.L. Kane, J.P. Leveille	(MICH)
CABIBBO	81	PL 105B 155	N. Cabibbo, G.R. Farrar, L. Maiani	(ROMA, RUTG)
FARRAR	78	PL 76B 575	G.R. Farrar, P. Fayet	(CIT)
Also		PL 79B 442	G.R. Farrar, P. Fayet	(CIT)